



ASTR/PHYS 1060: The Universe

Chapter 11: The Sun

Chapter 11 Reading Assignment due today at 10:45am

In-class/HW Assignment due Monday, October 15th
(or today for +5pts bonus)

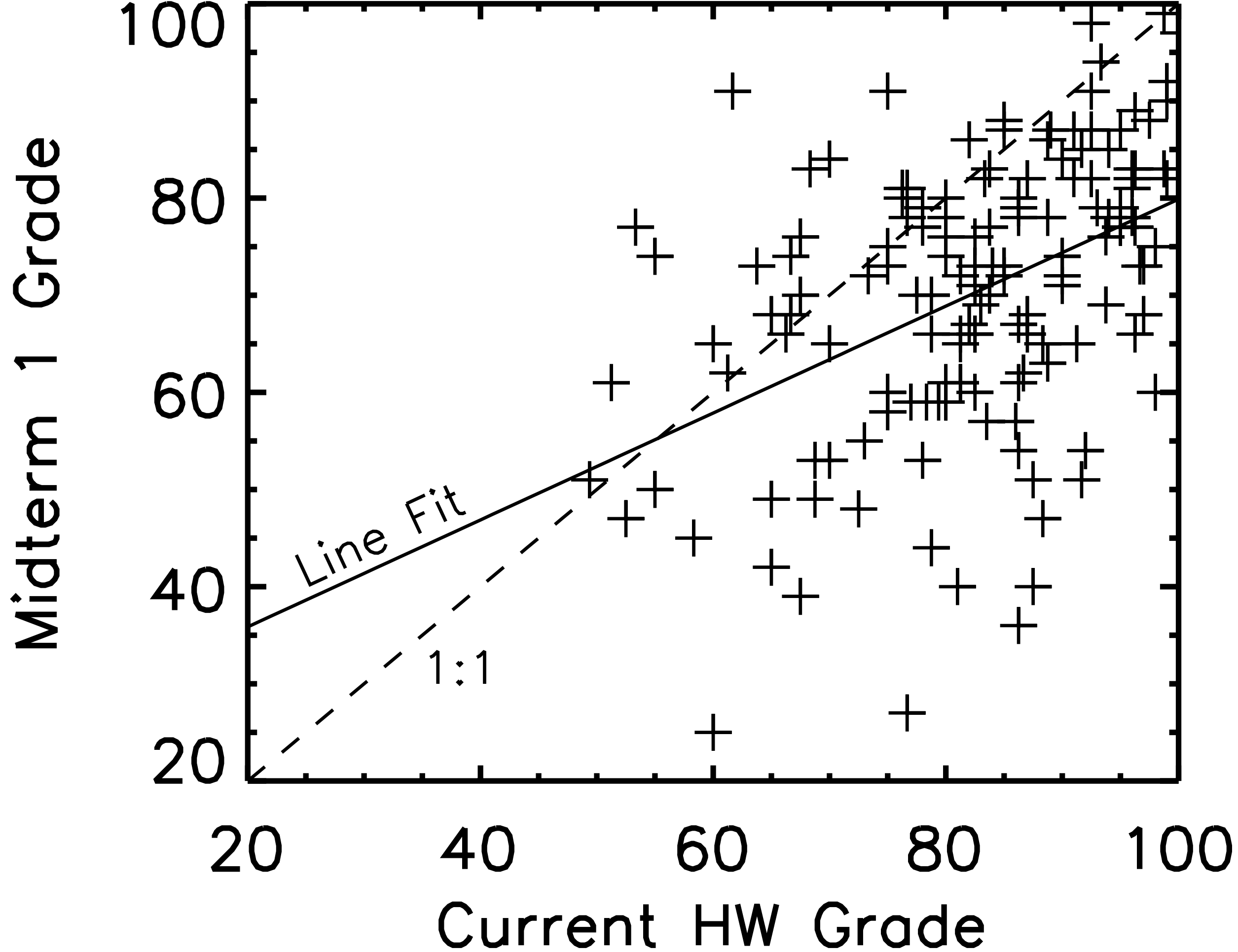
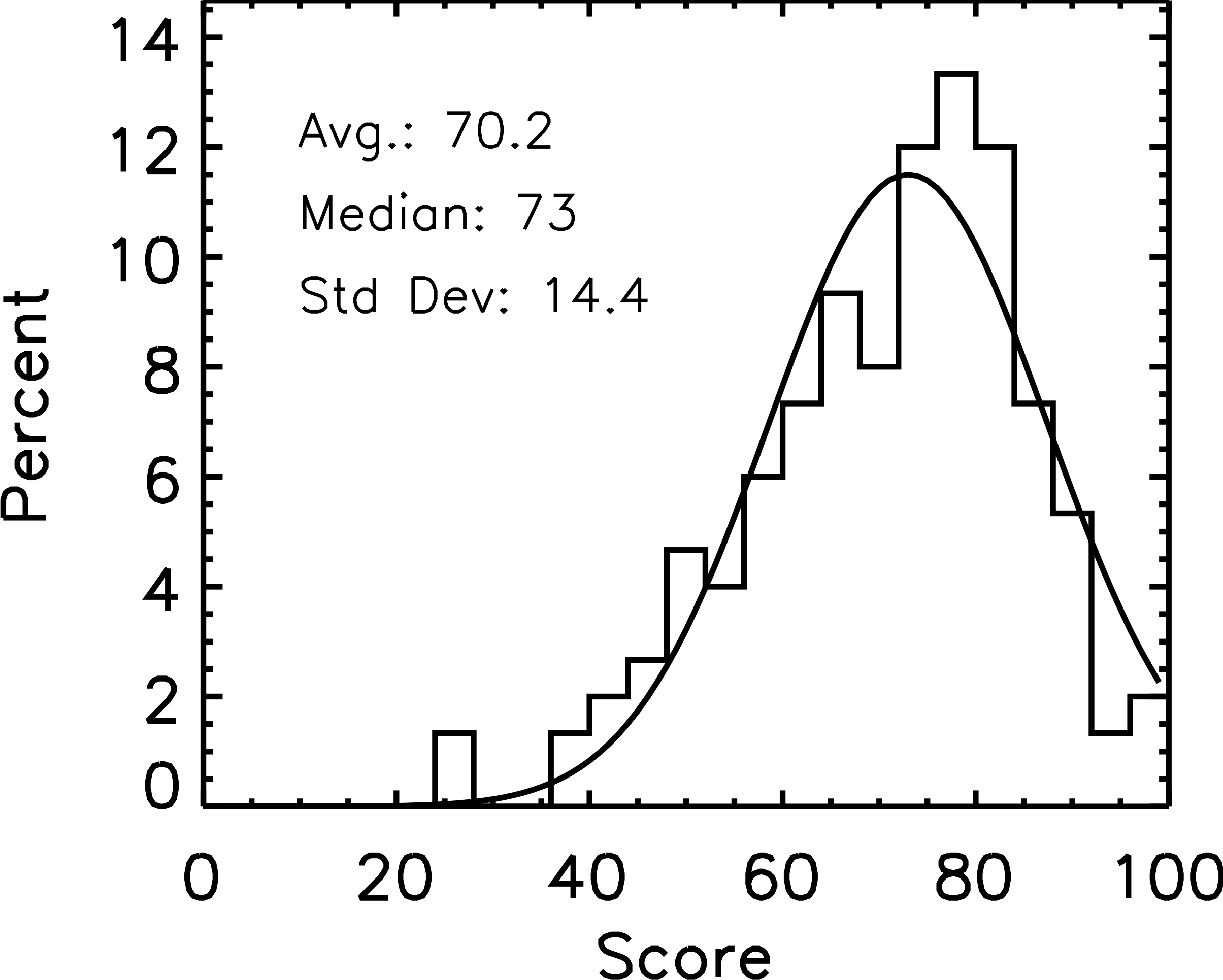
Are your grades in Canvas correct???

Planetarium Extra Credit
Opportunity!
(see the syllabus)

Oct. 11th or 13th at 6:45 pm
for the "Gateway to the Stars
Program."

Free tickets available here,
\$2 otherwise

Midterm 1 Results



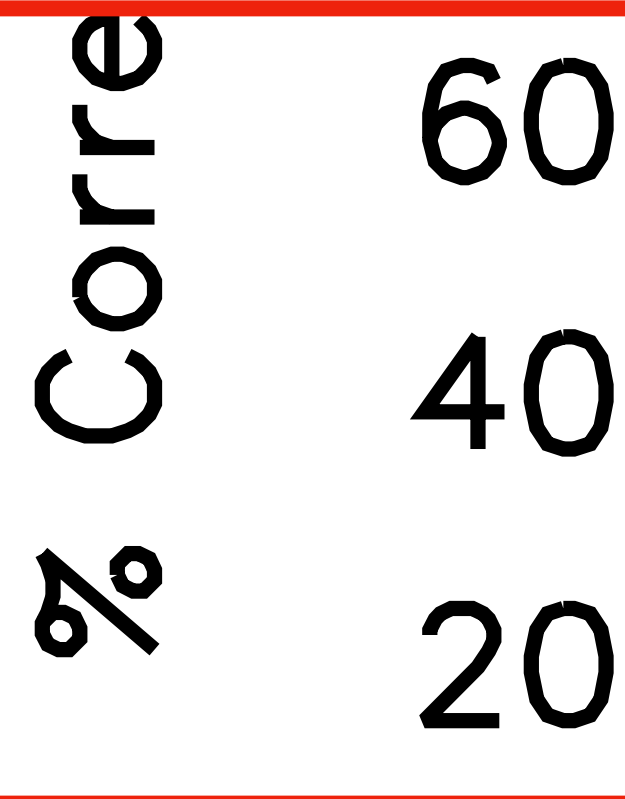
Multiple Choice Questions

_____ The diffraction limit of a 4-meter telescope is _____ than that of a 2-meter telescope.

- A. two times smaller
- B. four times smaller
- C. four times larger
- D. two times larger



_____ spacecraft feel weightless in space because:
_____ orbit.



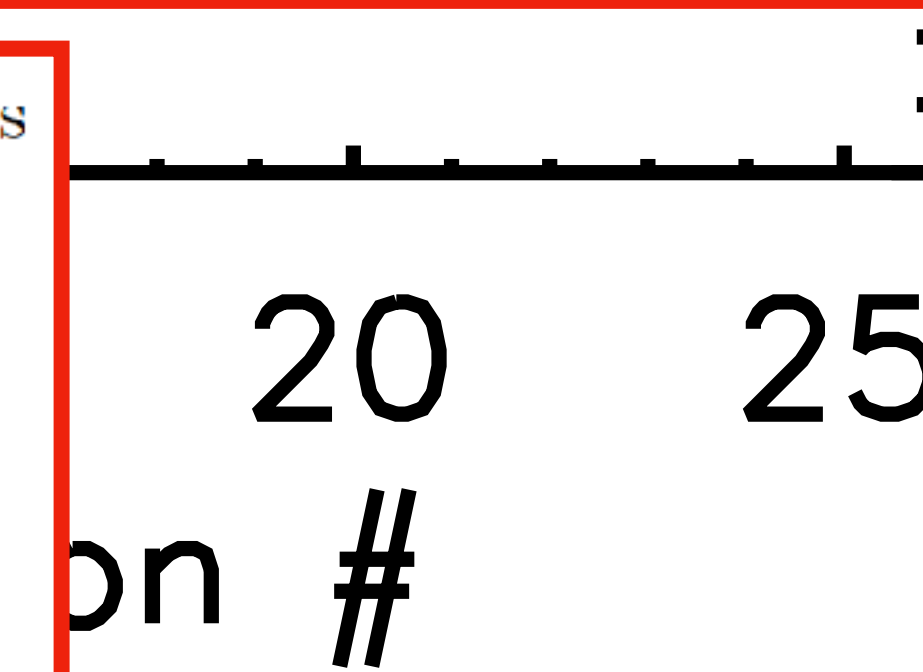
B. the added gravitational pull of the Moon balances Earth's gravitational pull.

_____ The solid form of a volatile material is generally referred to as:

- A. rock.
- B. a metal.
- C. a silicate.
- D. ice.

_____ When we determine the angular resolution of an interferometric array of radio telescopes using the formula $\theta = \lambda/D$, the variable D stands for the:

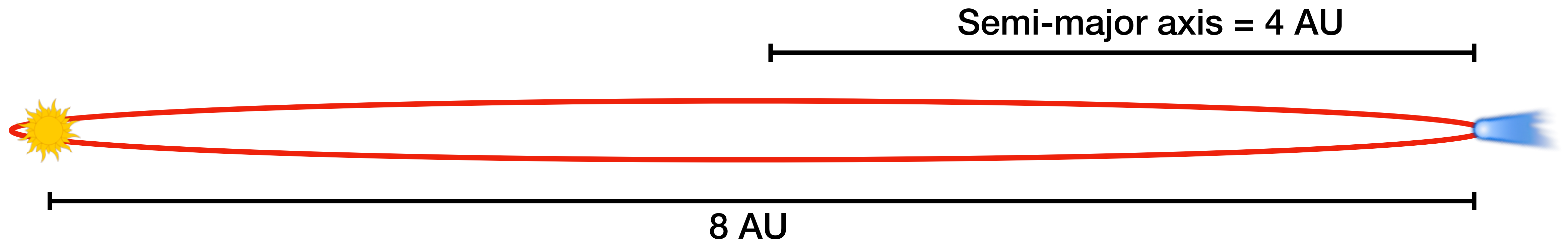
- A. number of telescopes.
- B. focal length of the telescopes.
- C. separation between the telescopes.
- D. diameter of the telescopes.



Kepler's Law Short Answer Question

28. An asteroid is discovered to have an extremely elliptical orbit, such that its major axis is much much larger than its minor axis. It is found that its farthest distance from the Sun is 8 AU. How long does it take to complete one orbit?

Solution: Because the orbit is highly elliptical, the foci are close to the outer edges of the ellipse, with the Sun at one of the foci. This means that the asteroid passes very near the Sun on closest approach, so its farthest distance of 8 AU corresponds to the entire length of the major axis, yielding a semi-major axis $a = 4$ AU. Plugging this into Kepler's third law, $P(\text{in years})^2 = a(\text{in AU})^3$, we find the asteroid's period $P = \sqrt{4^3} = \sqrt{64} = 8$ years.



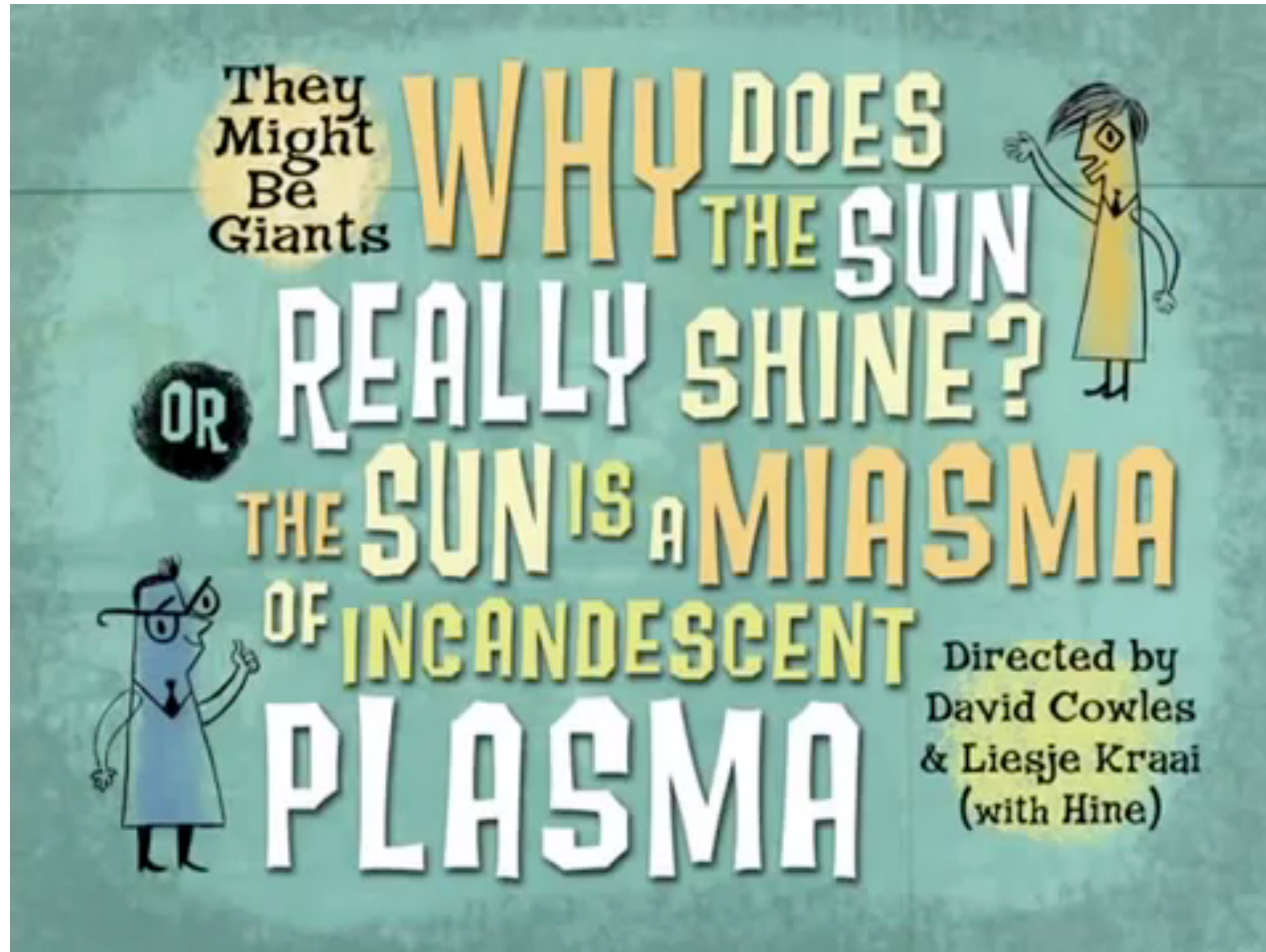
Why does the Sun shine?

<https://www.youtube.com/watch?v=3JdWISF195Y>

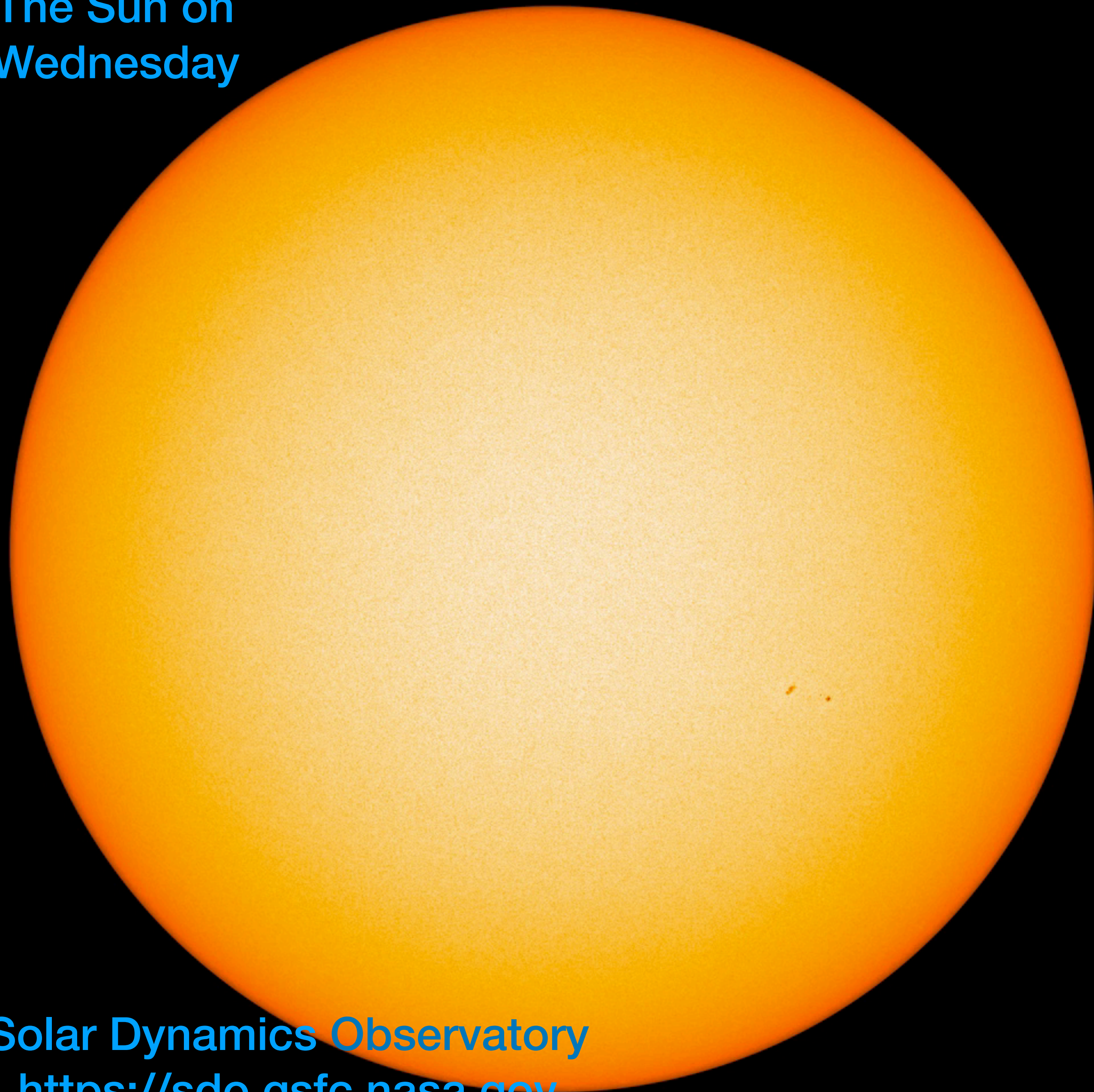


A Small Correction...

<https://www.youtube.com/watch?v=sLkGSV9WDMMA>



The Sun on
Wednesday



Solar Dynamics Observatory
<https://sdo.gsfc.nasa.gov>

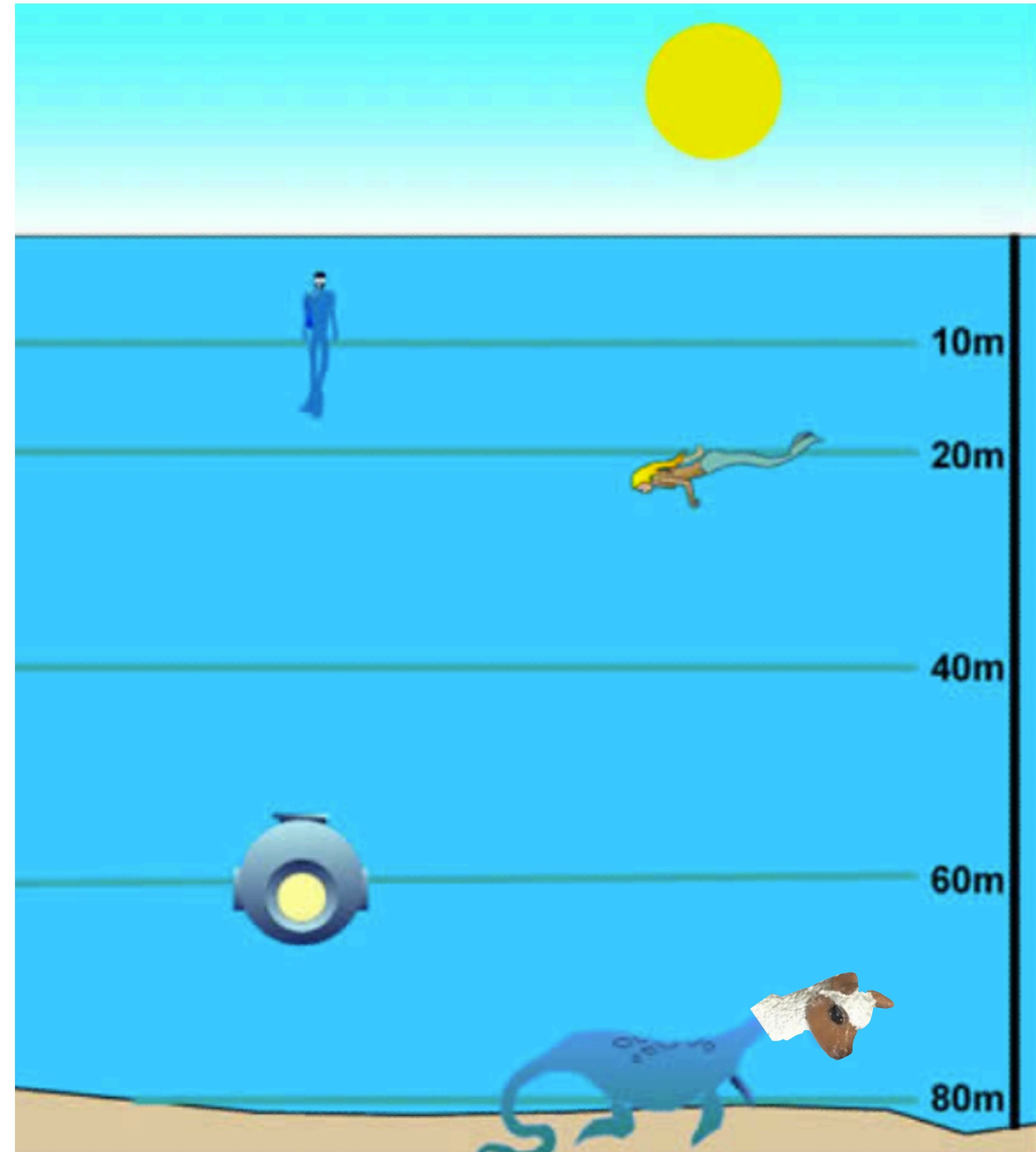
SDO/HMI

Sun Facts

- Diameter: 1.4 million km (~100x Earth's)
- Mass: 2×10^{30} kg (300,000x Earth's)
- Total Luminosity: 4×10^{26} Watts
- Energy Reaching Earth: 1400 Watts/meter²
- Surface Temperature: 5800 K

The greatest pressure is experienced by the...

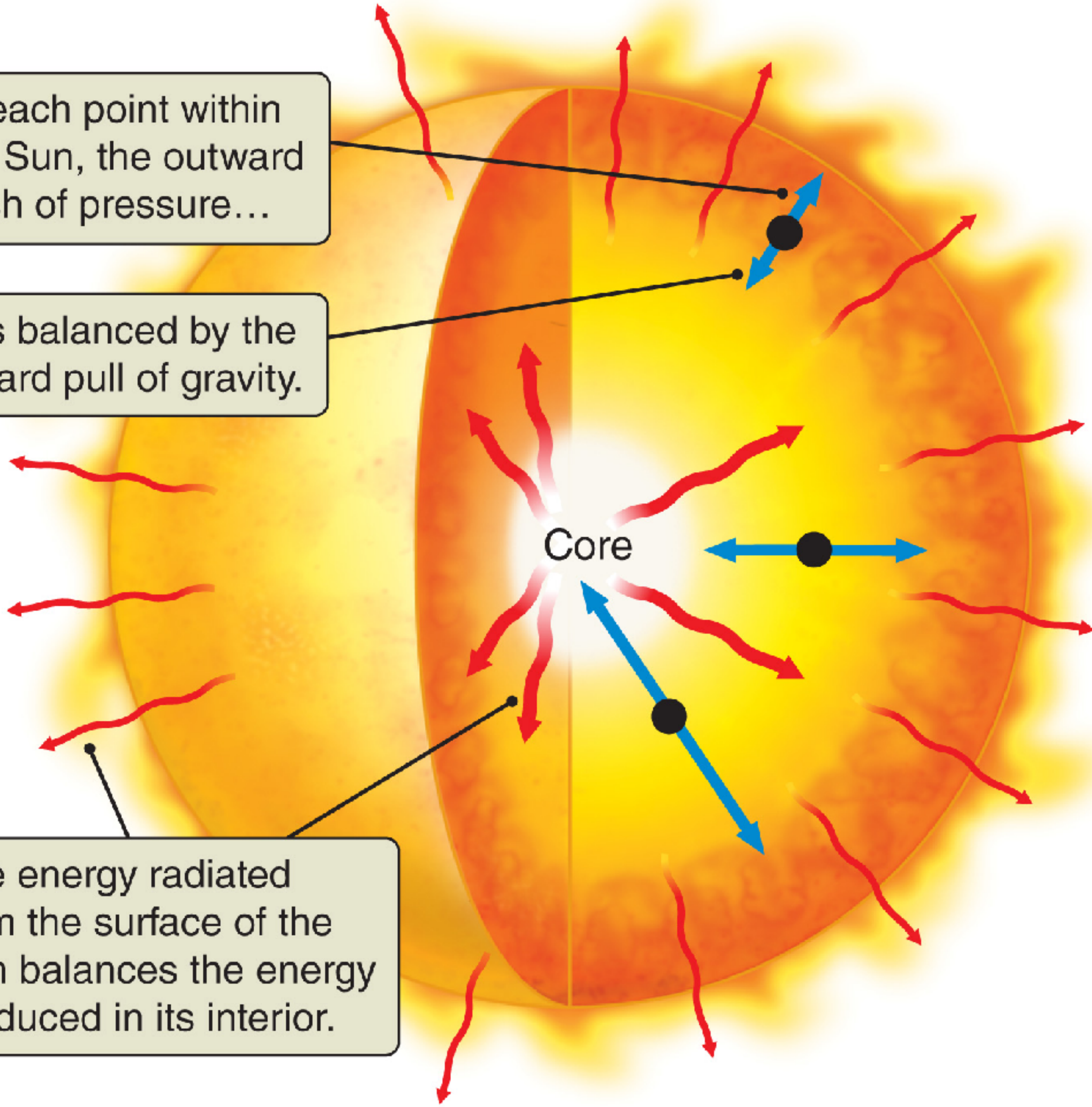
- A) Diver
- B) Mermaid
- C) Bathysphere
- D) Loch Ness Llama



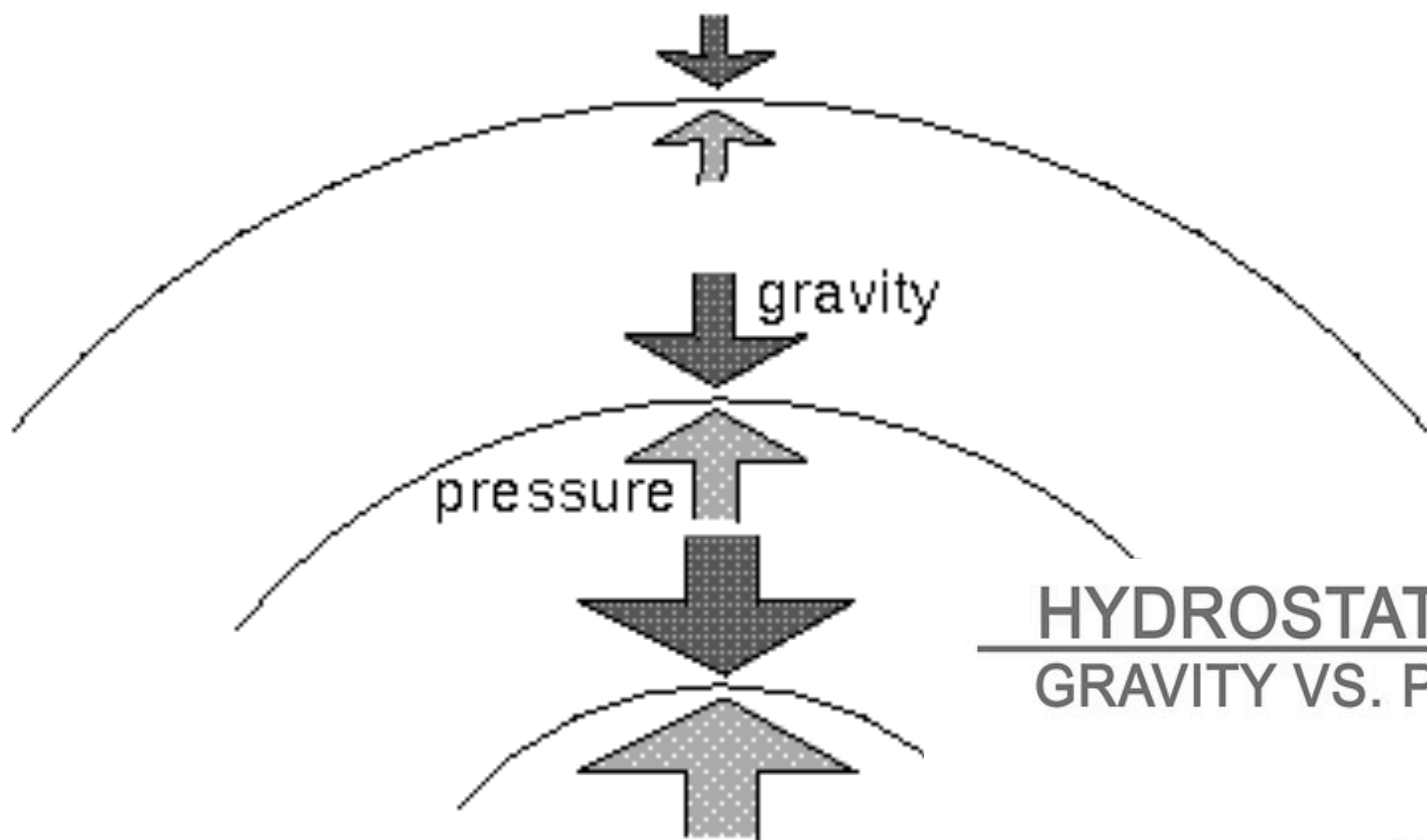
At each point within the Sun, the outward push of pressure...

...is balanced by the inward pull of gravity.

The energy radiated from the surface of the Sun balances the energy produced in its interior.

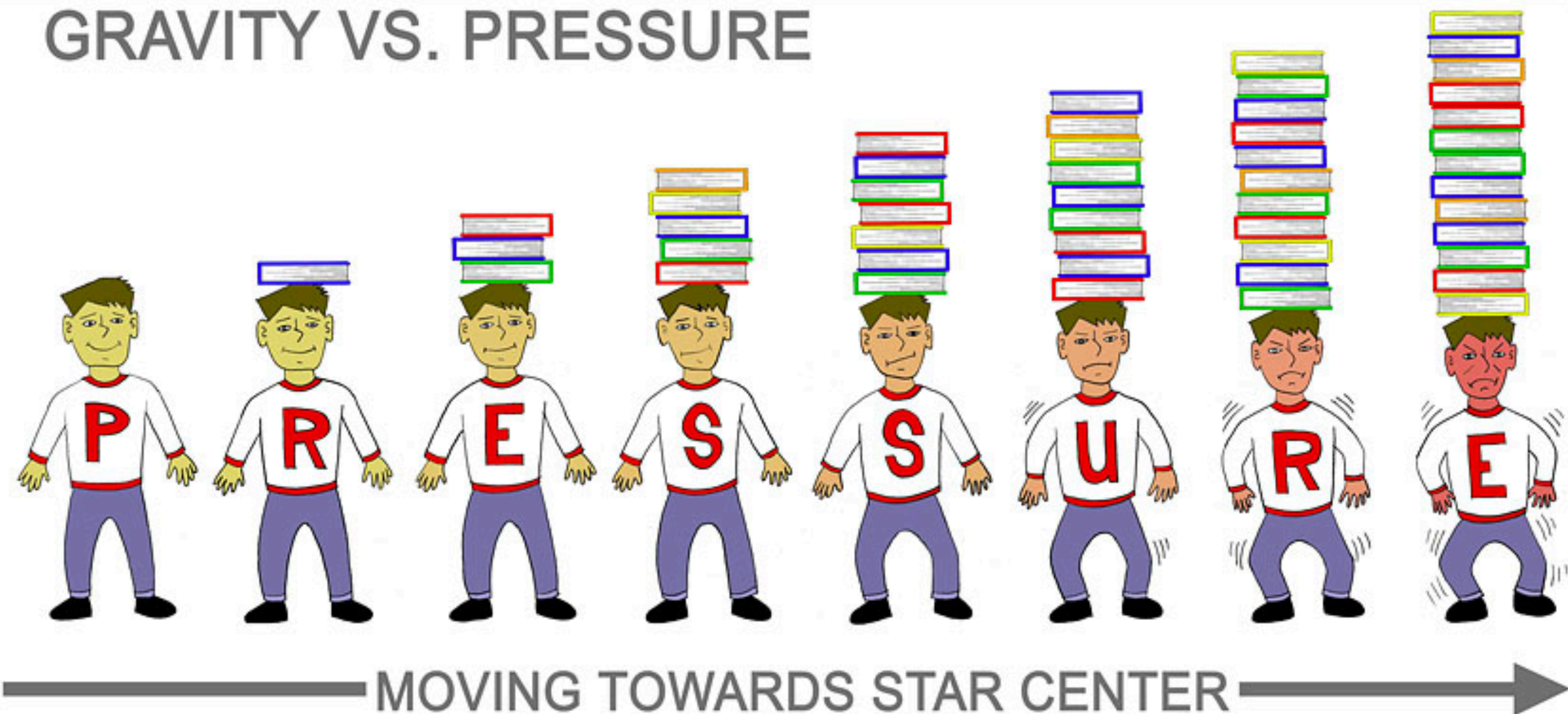


Hydrostatic Equilibrium

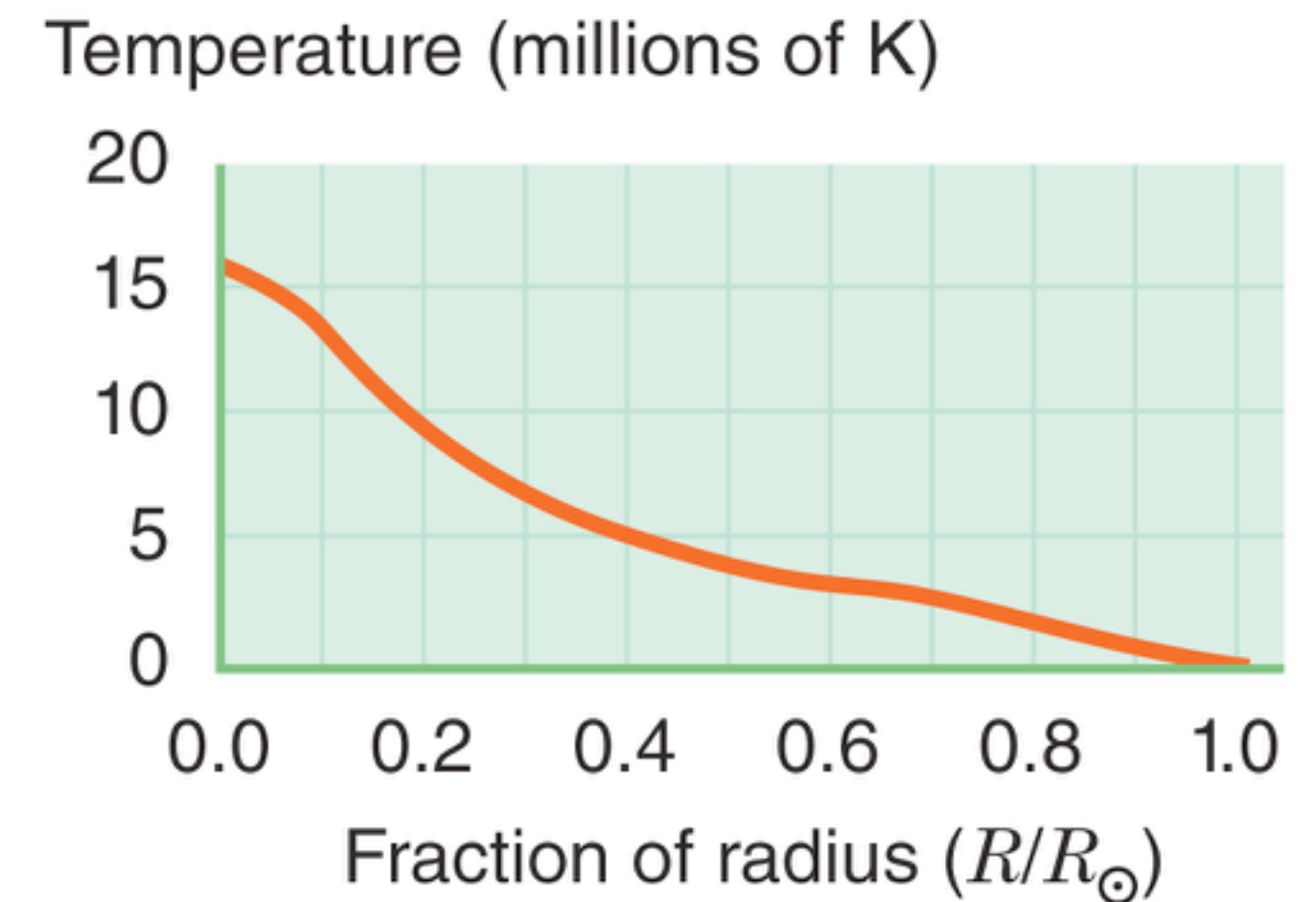
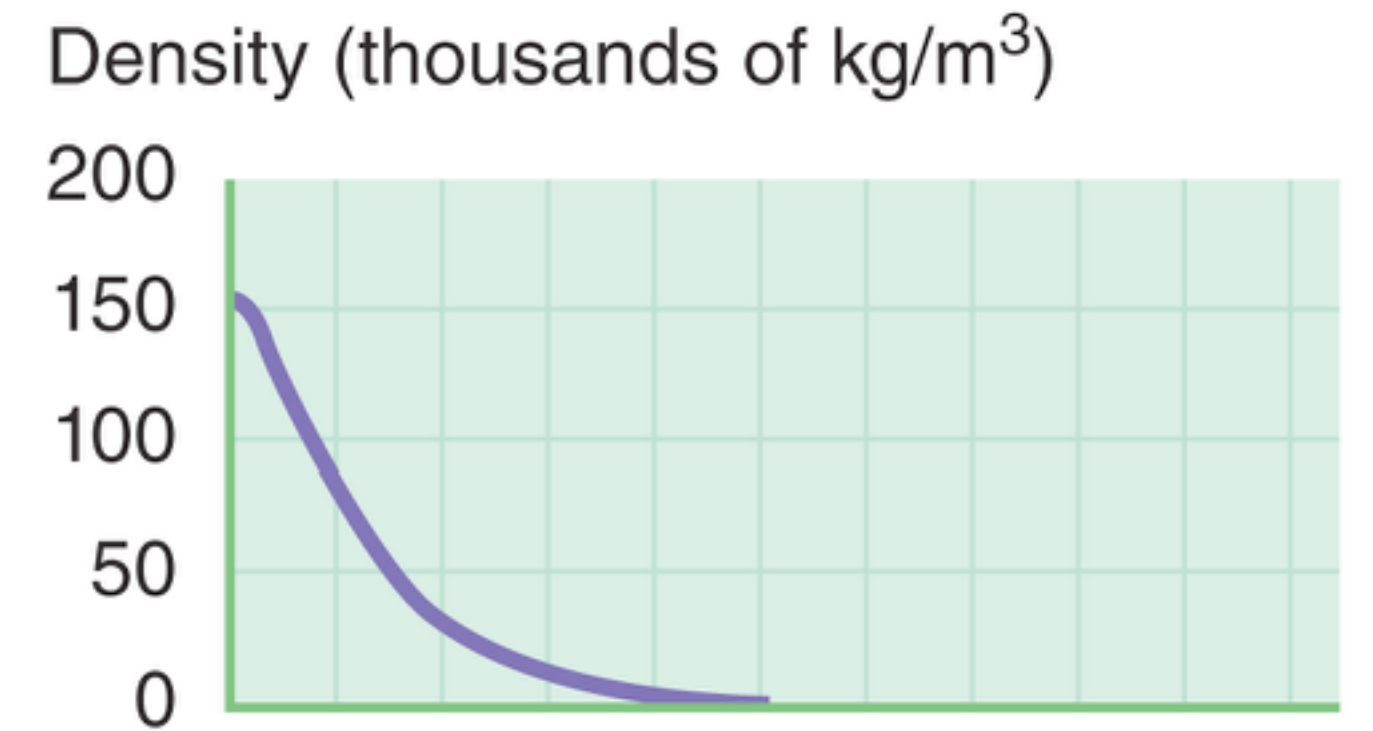
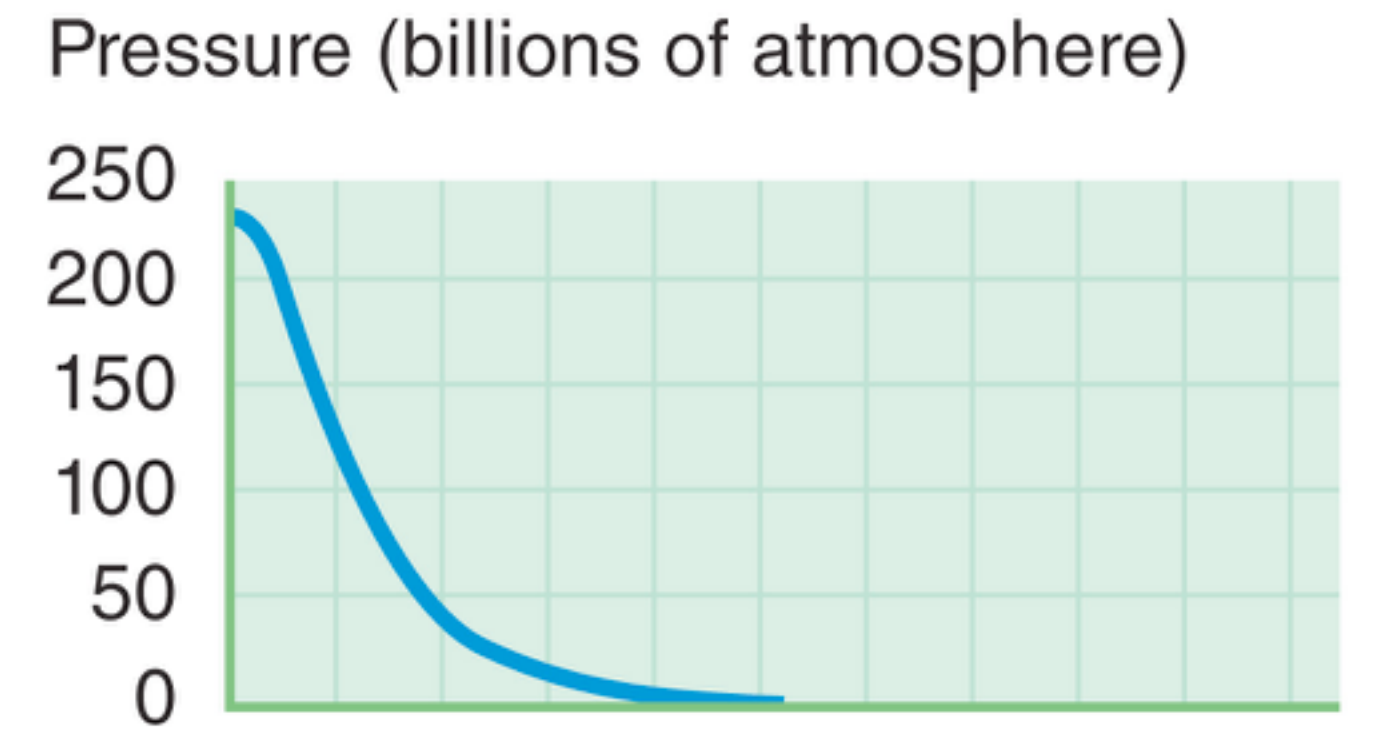
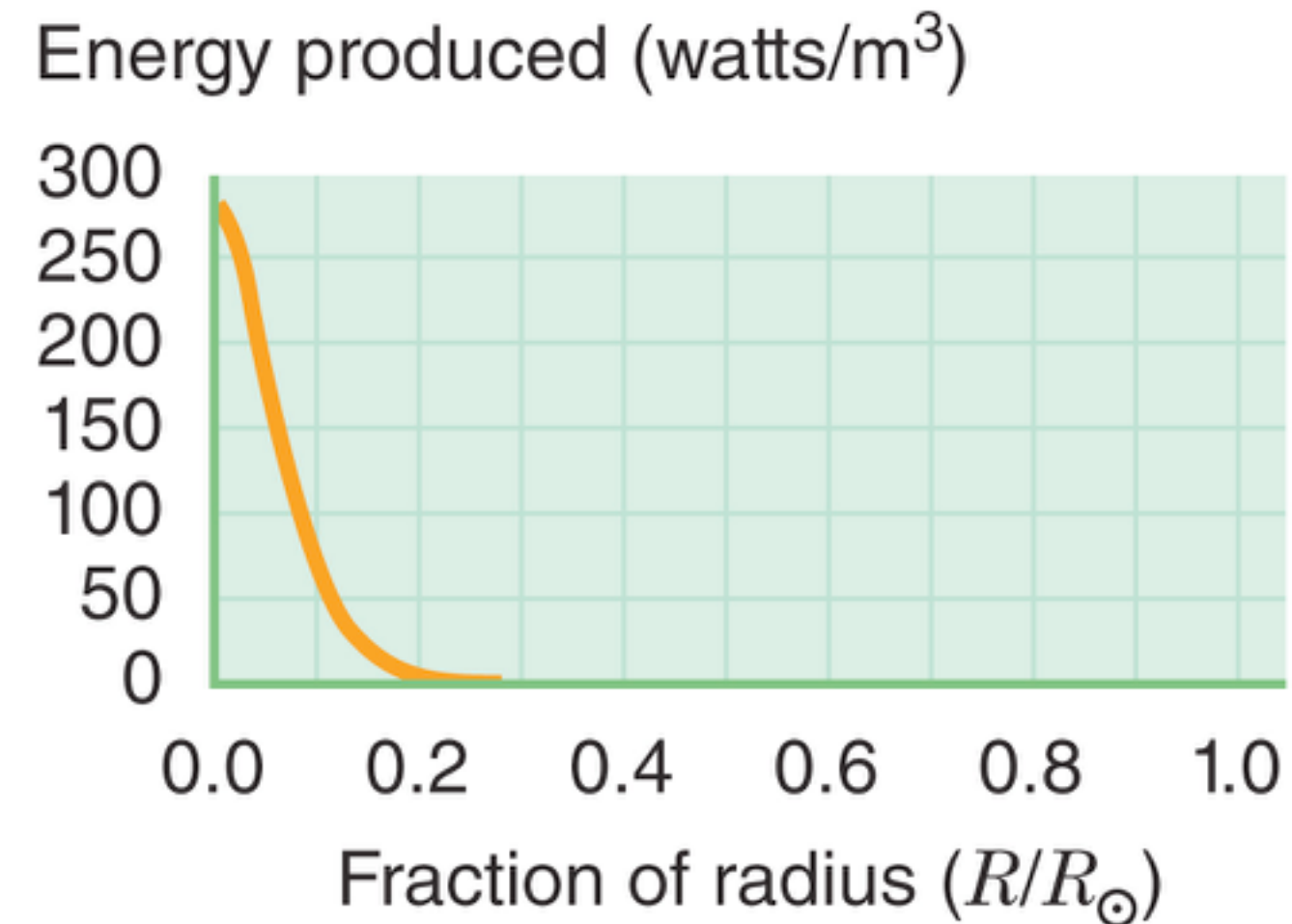
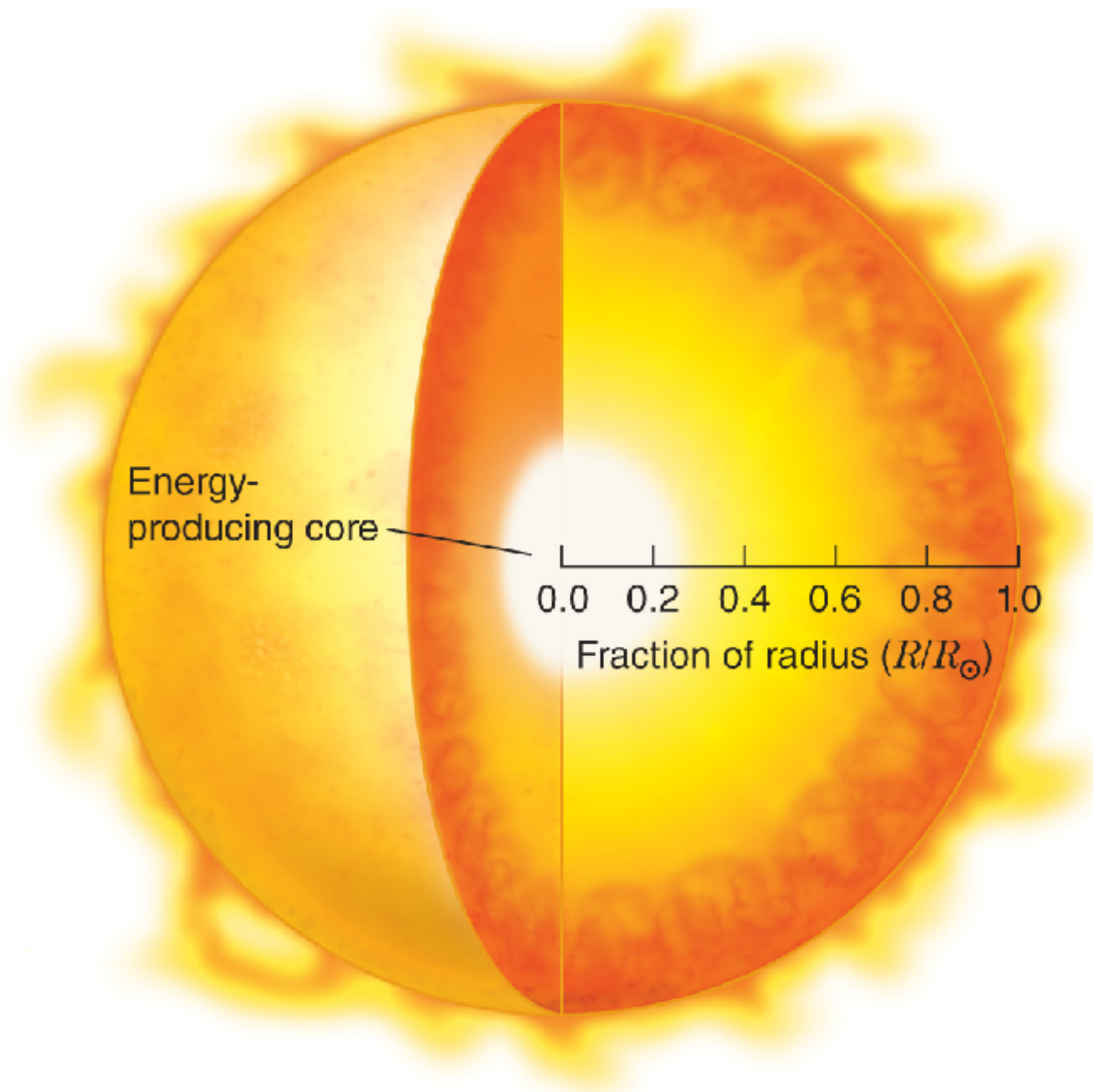


HYDROSTATIC EQUILIBRIUM IN A STAR

GRAVITY VS. PRESSURE



Pressure, Temperature, Density, Energy Produced all change with radius



The Sun's Luminosity Comes From:

- A) Chemical Burning (loss of electron energy)**
- B) Nuclear Fusion**
- C) Gravitational Collapse**

Fusion happens at the center of the sun because:



A) High Temperature

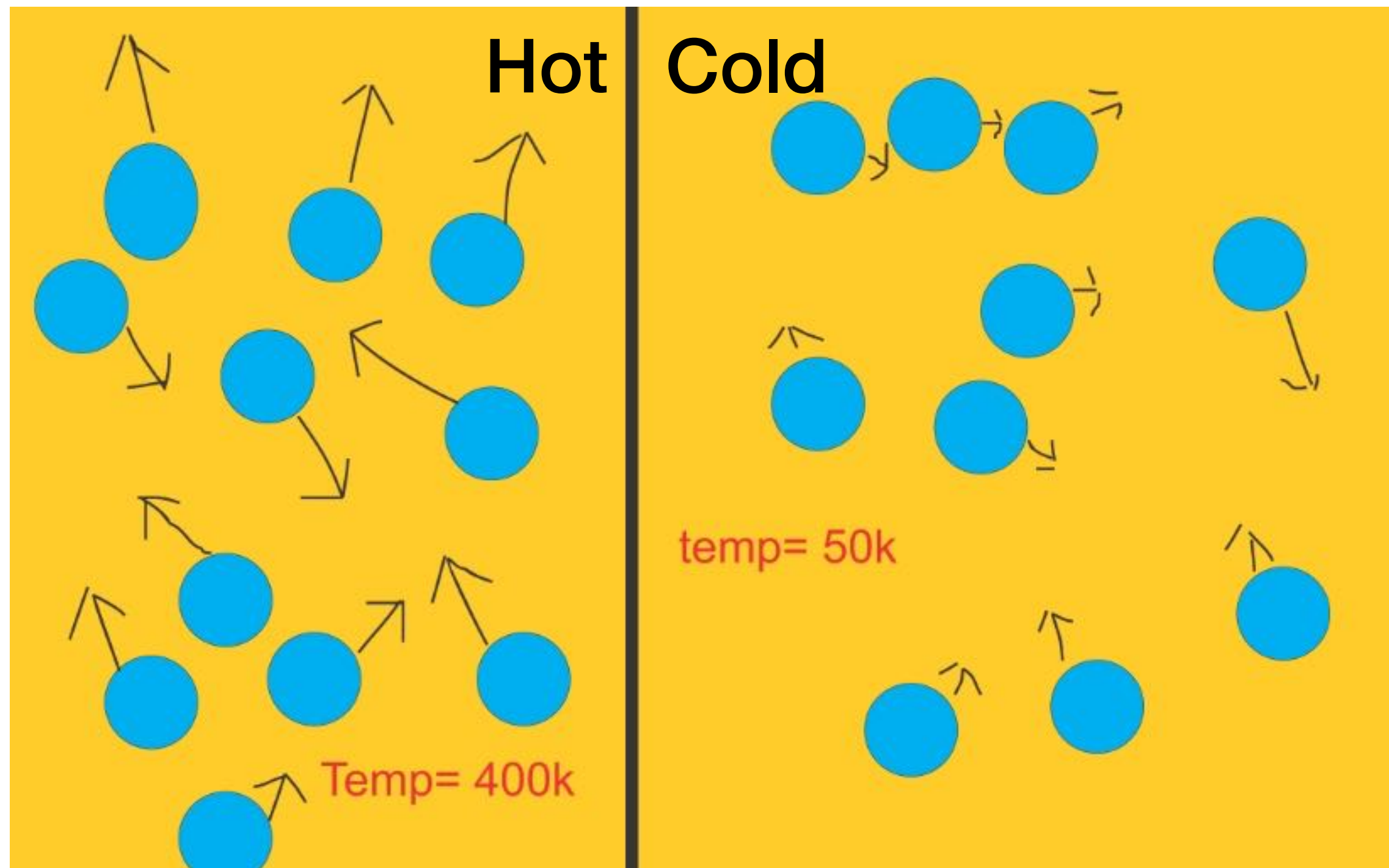


B) High Density

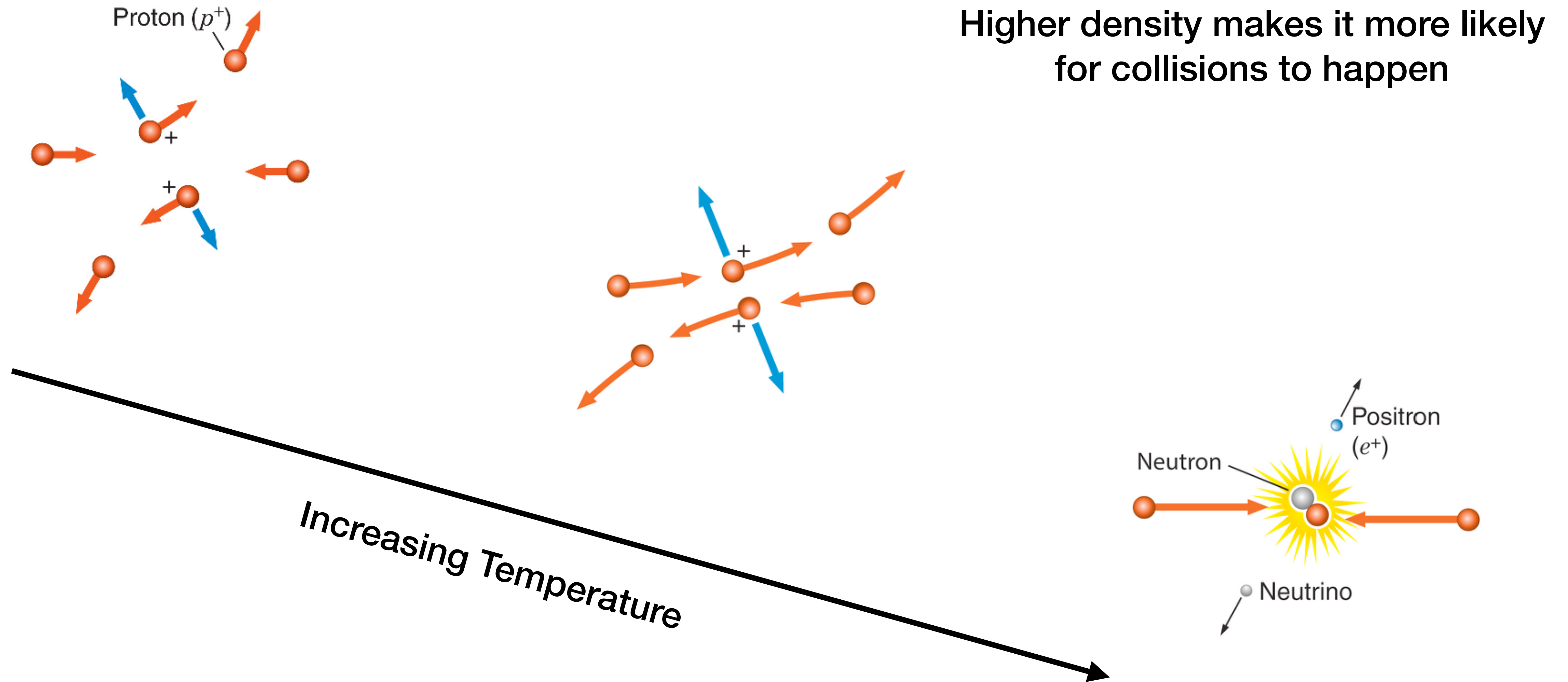
C) High Density & Temperature

Higher Temperature \rightarrow
Faster Movement

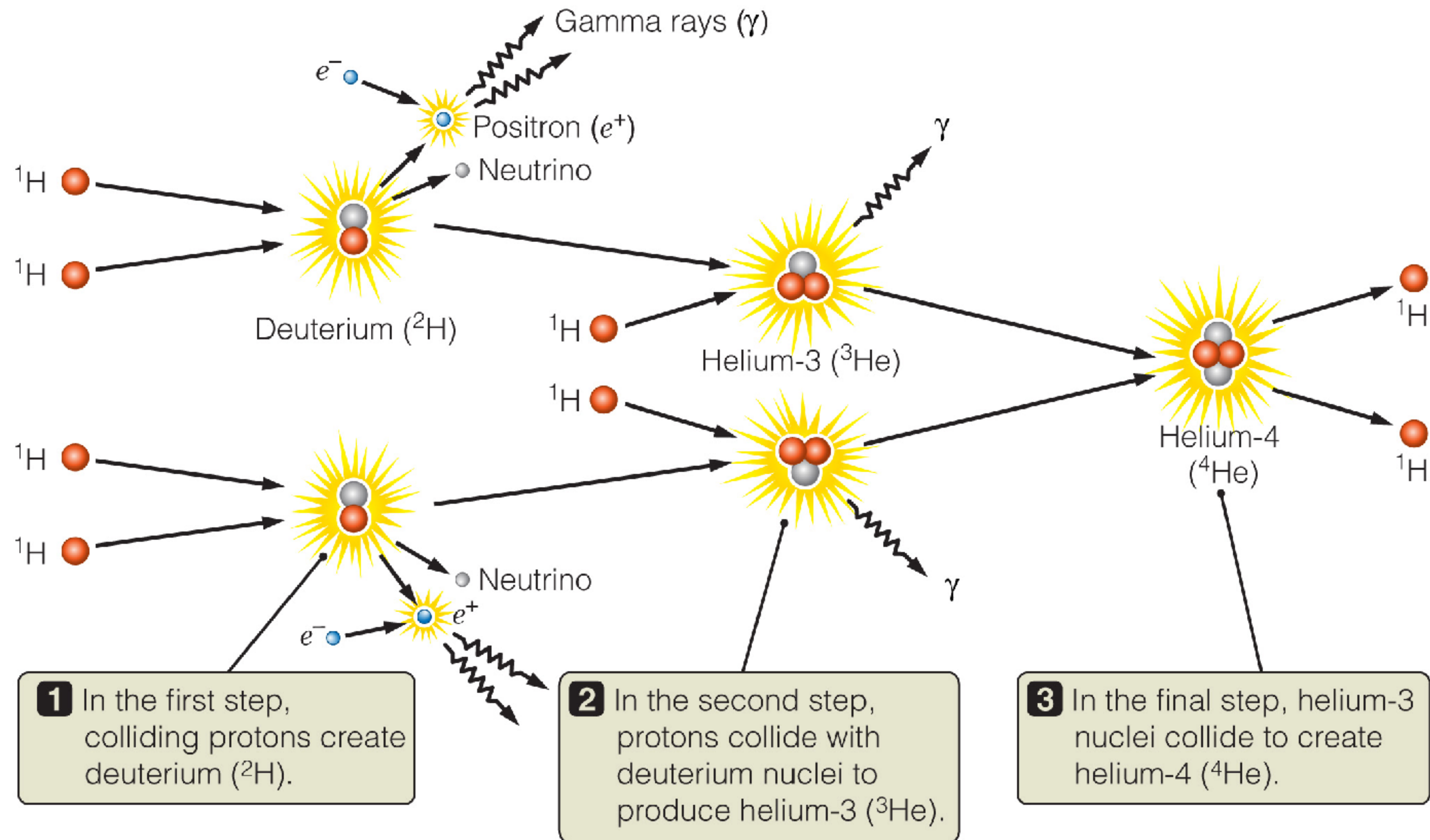
High Density \rightarrow More
likely to run into things



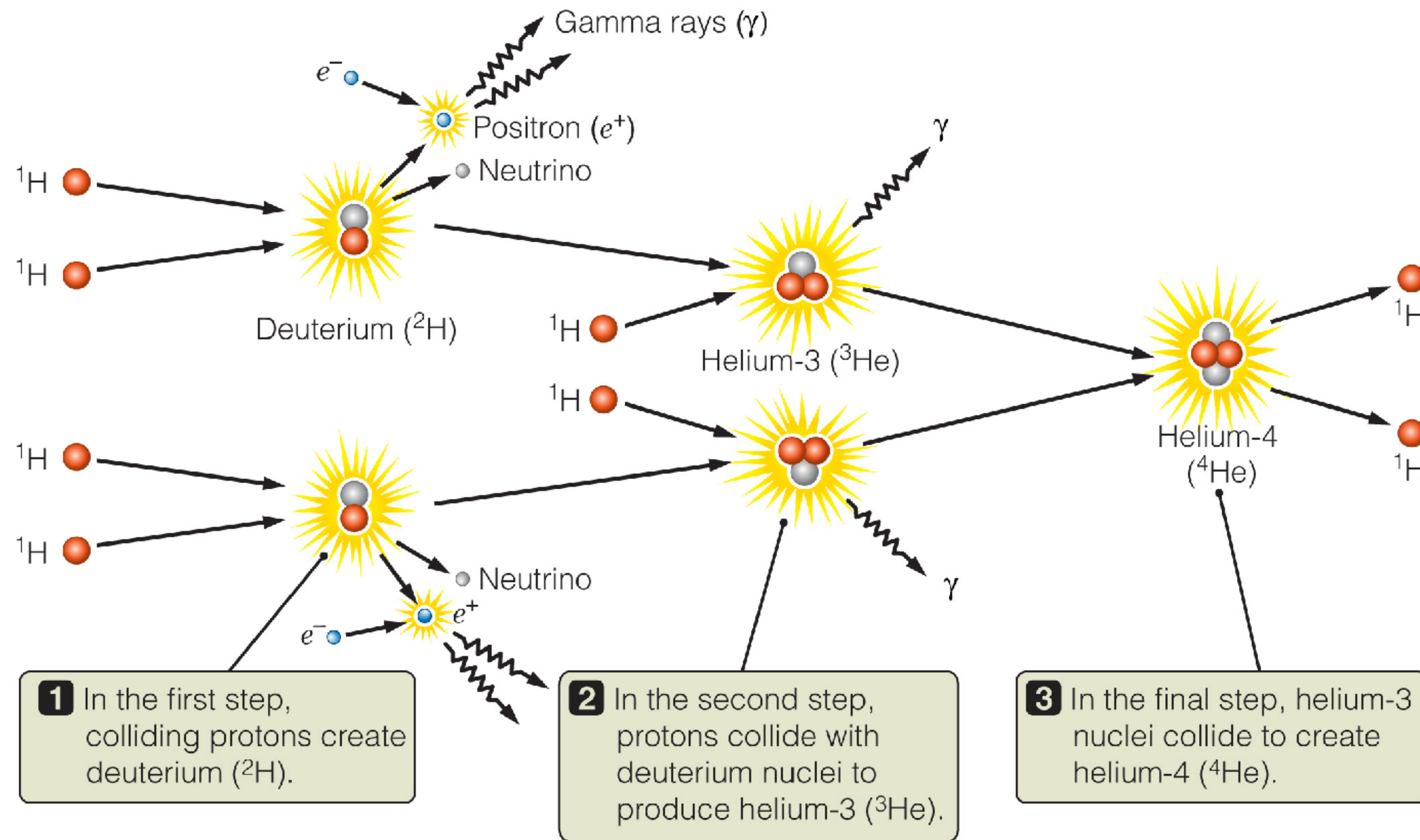
Higher temperature makes collisions possible



Hydrogen burning via the p-p chain



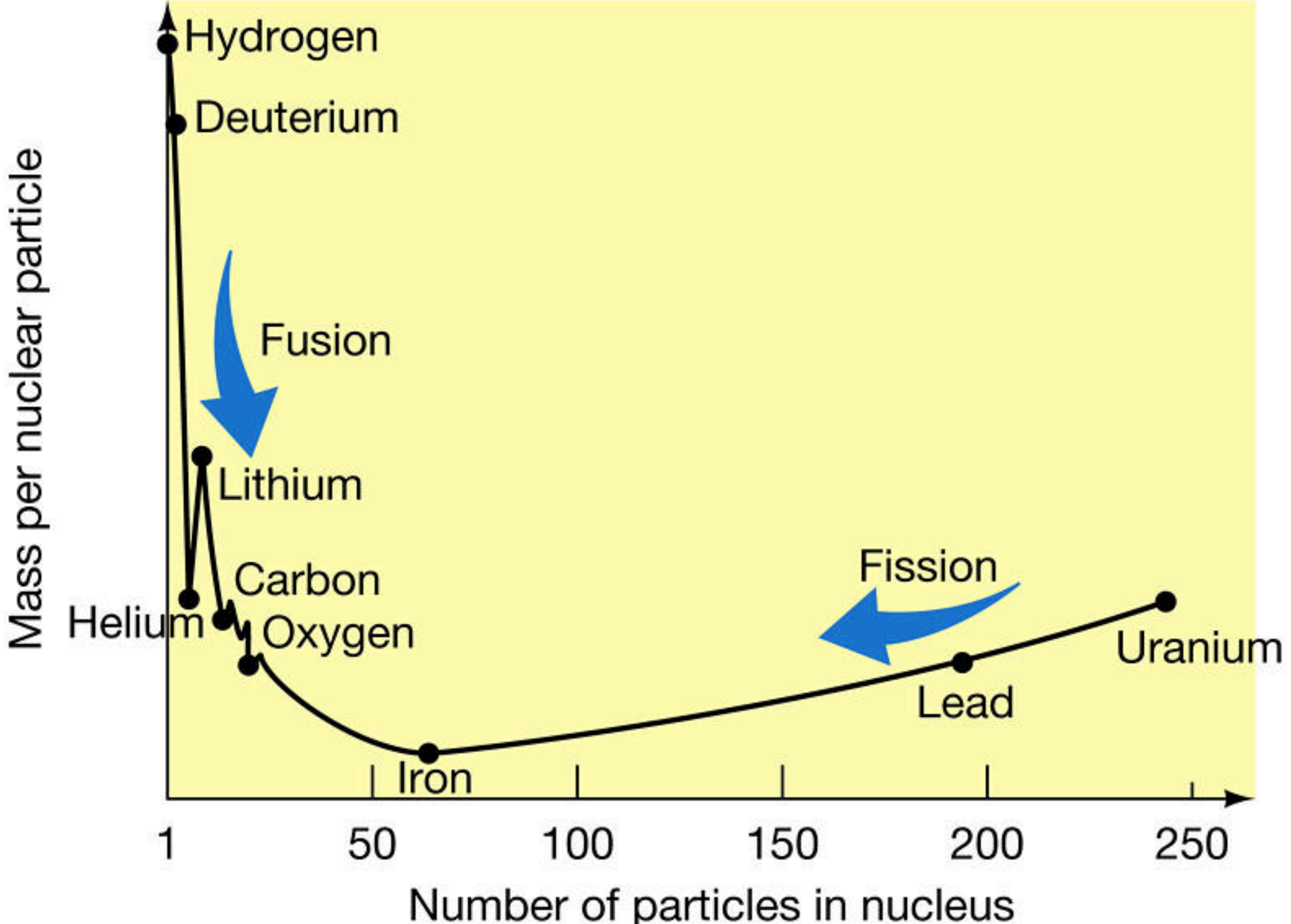
Energy is conserved (and $E=mc^2$), so...



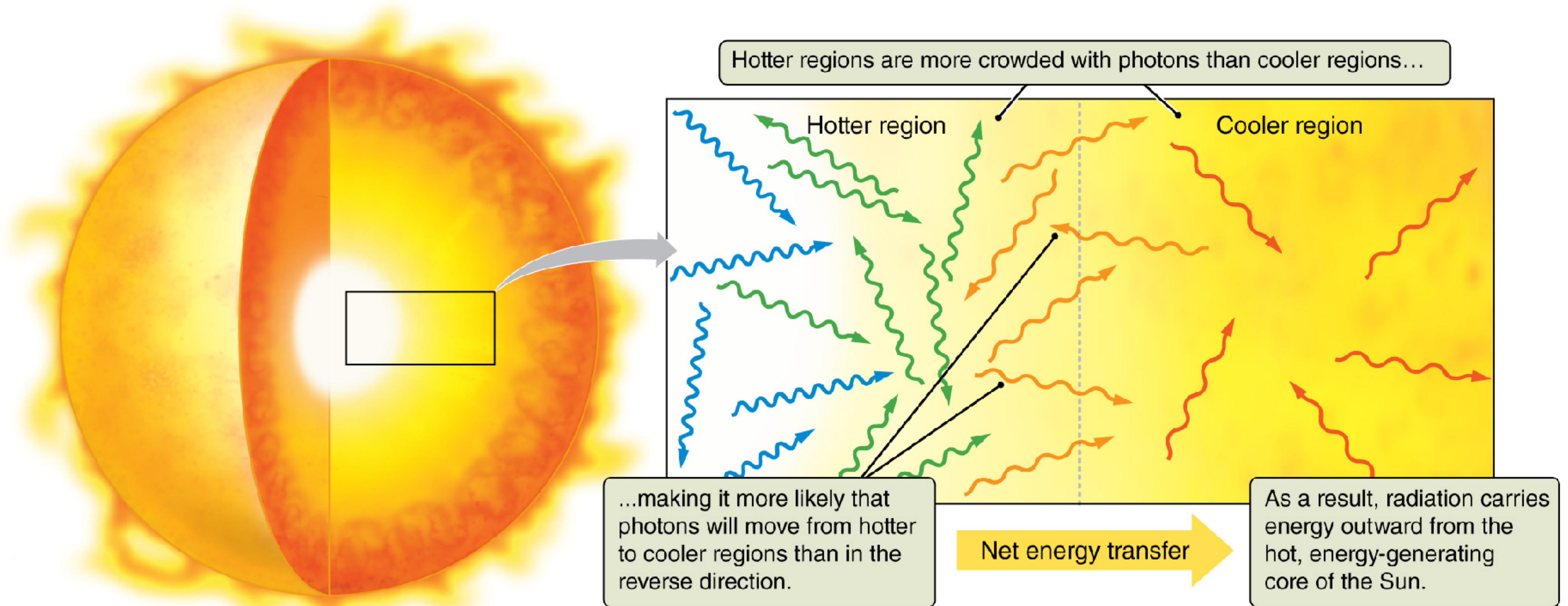
A) then 4 Hydrogen atoms must weigh more than one Helium atom

B) then 4 Hydrogen atoms must weigh less than one Helium atom

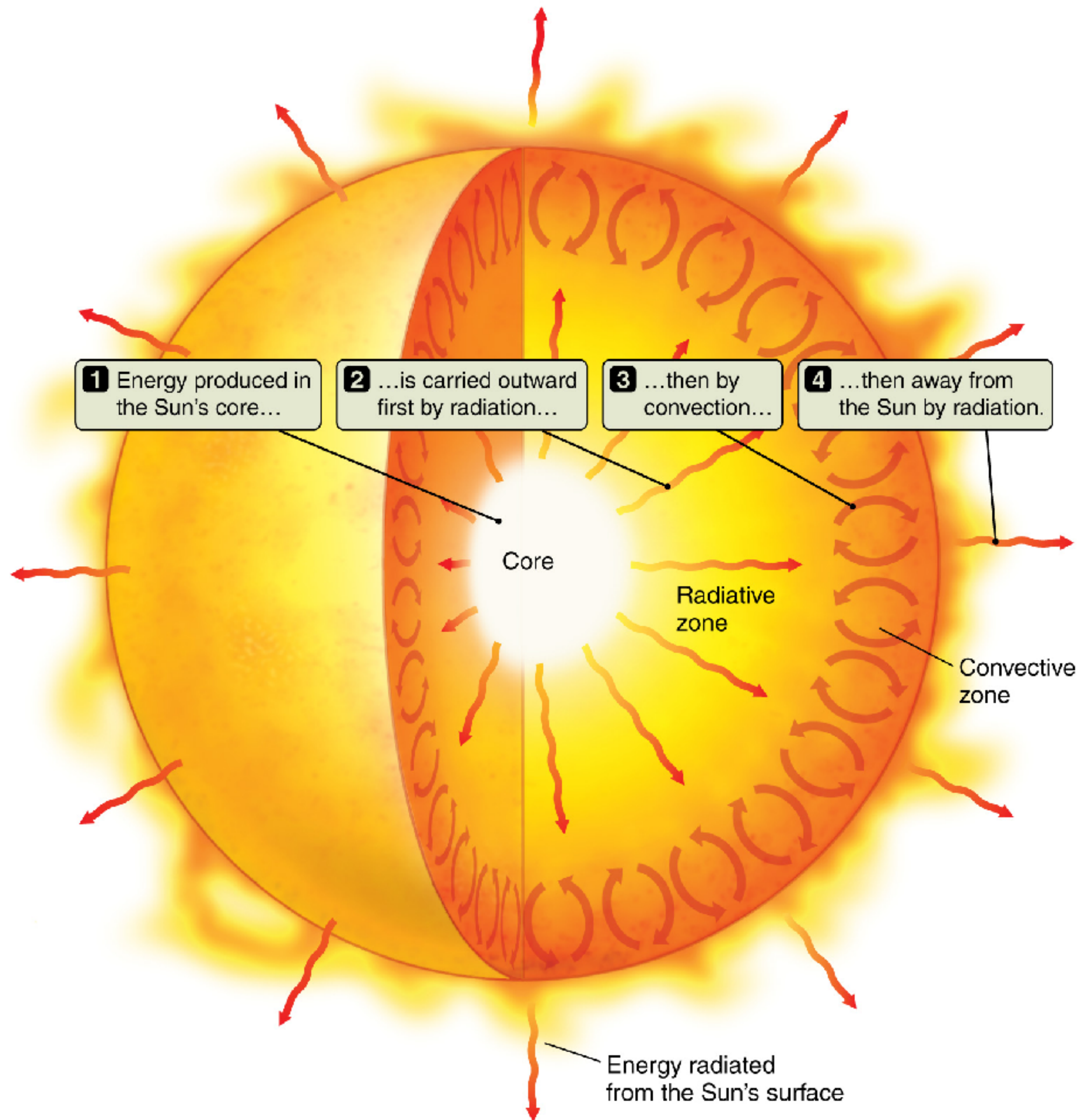
Energy from Mass



Photons are created in the core and move outward to the surface



How light escapes the Sun



A high-energy photon is produced by a nuclear reaction at the center of the Sun. Which diagram best represents how that photon's energy escapes from the Sun's radiative zone?

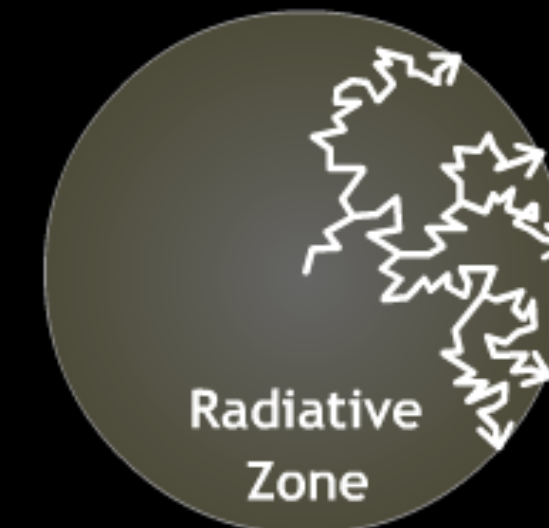
(A)



(B)



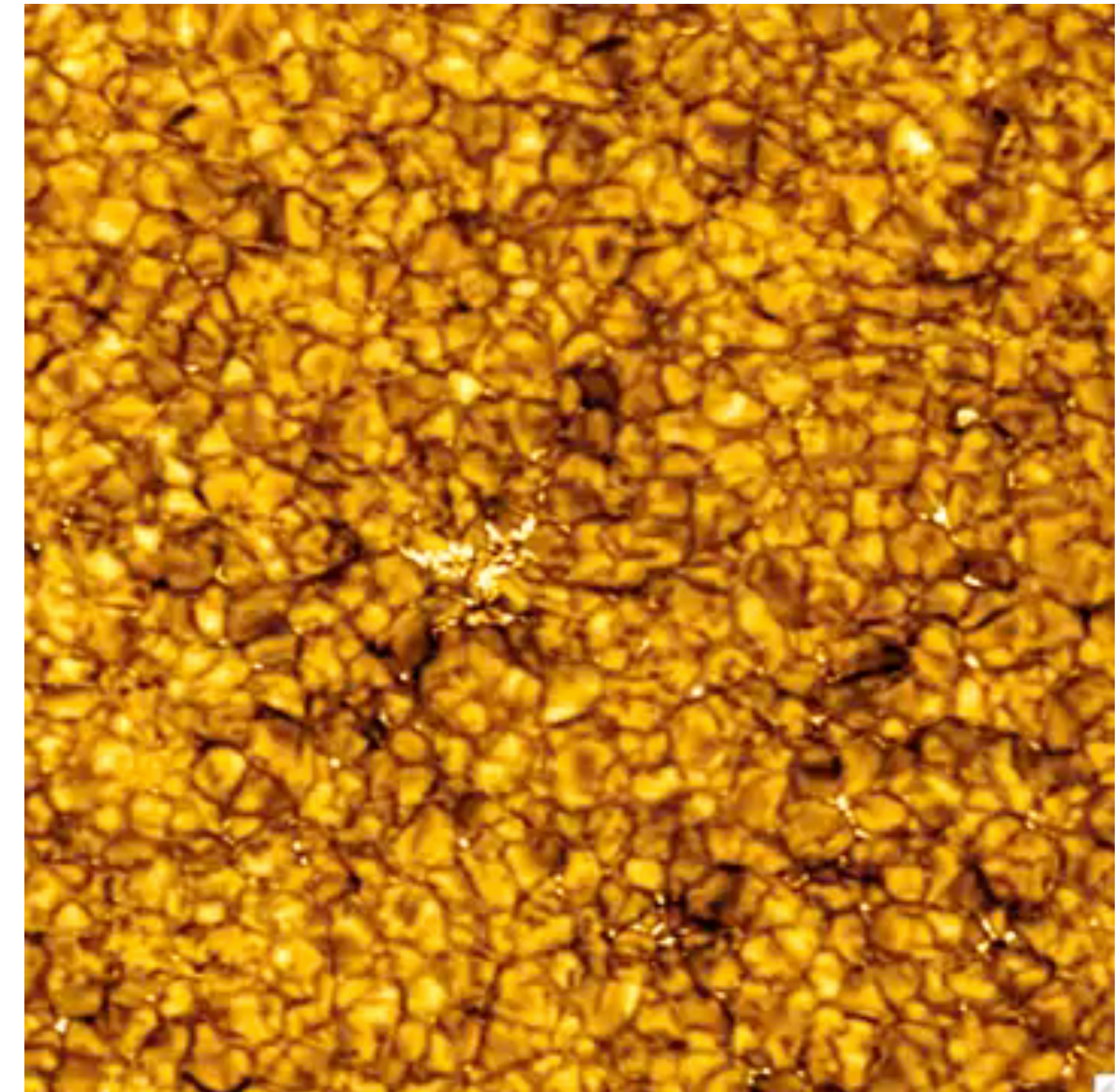
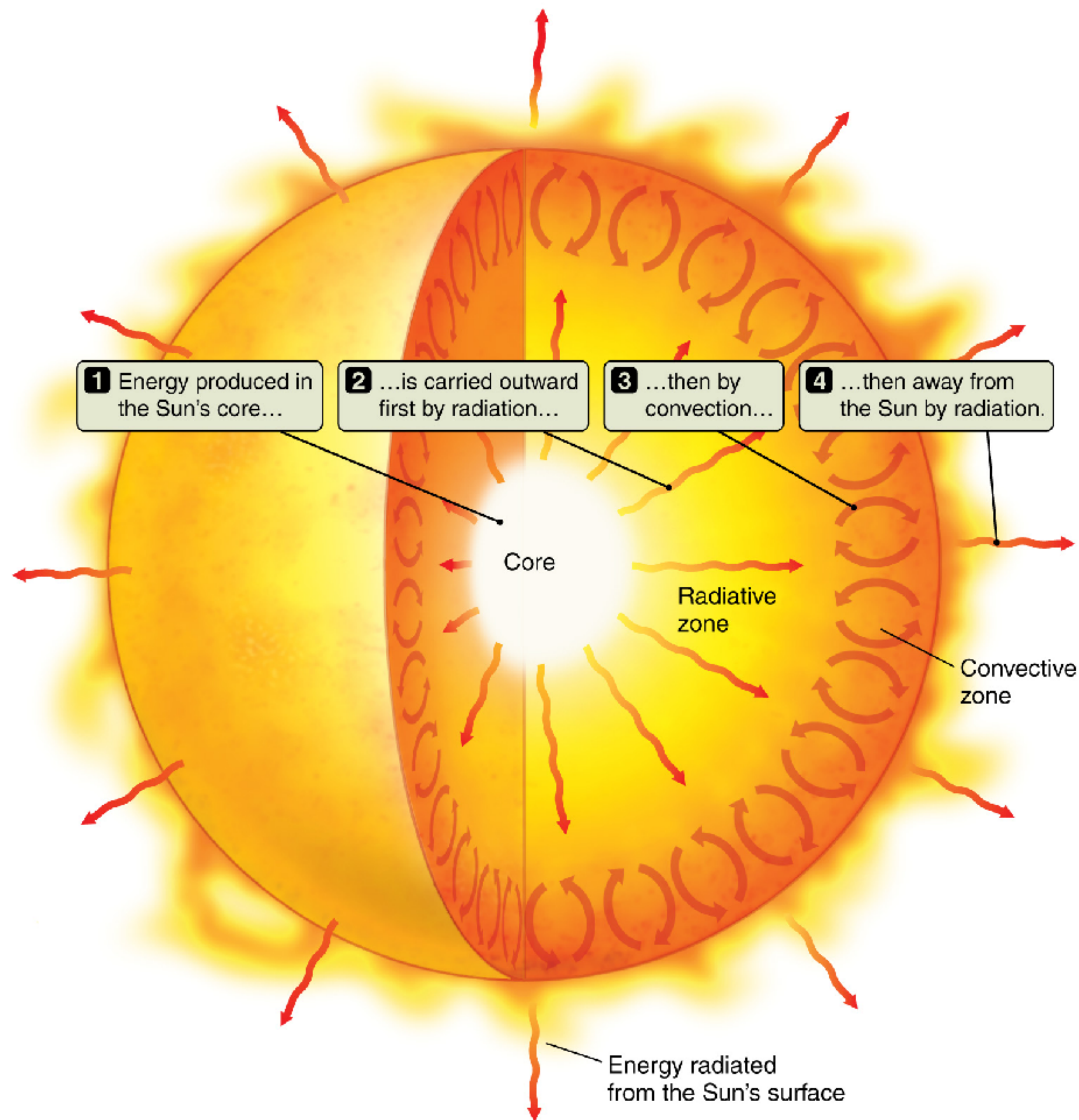
(C)



(D)

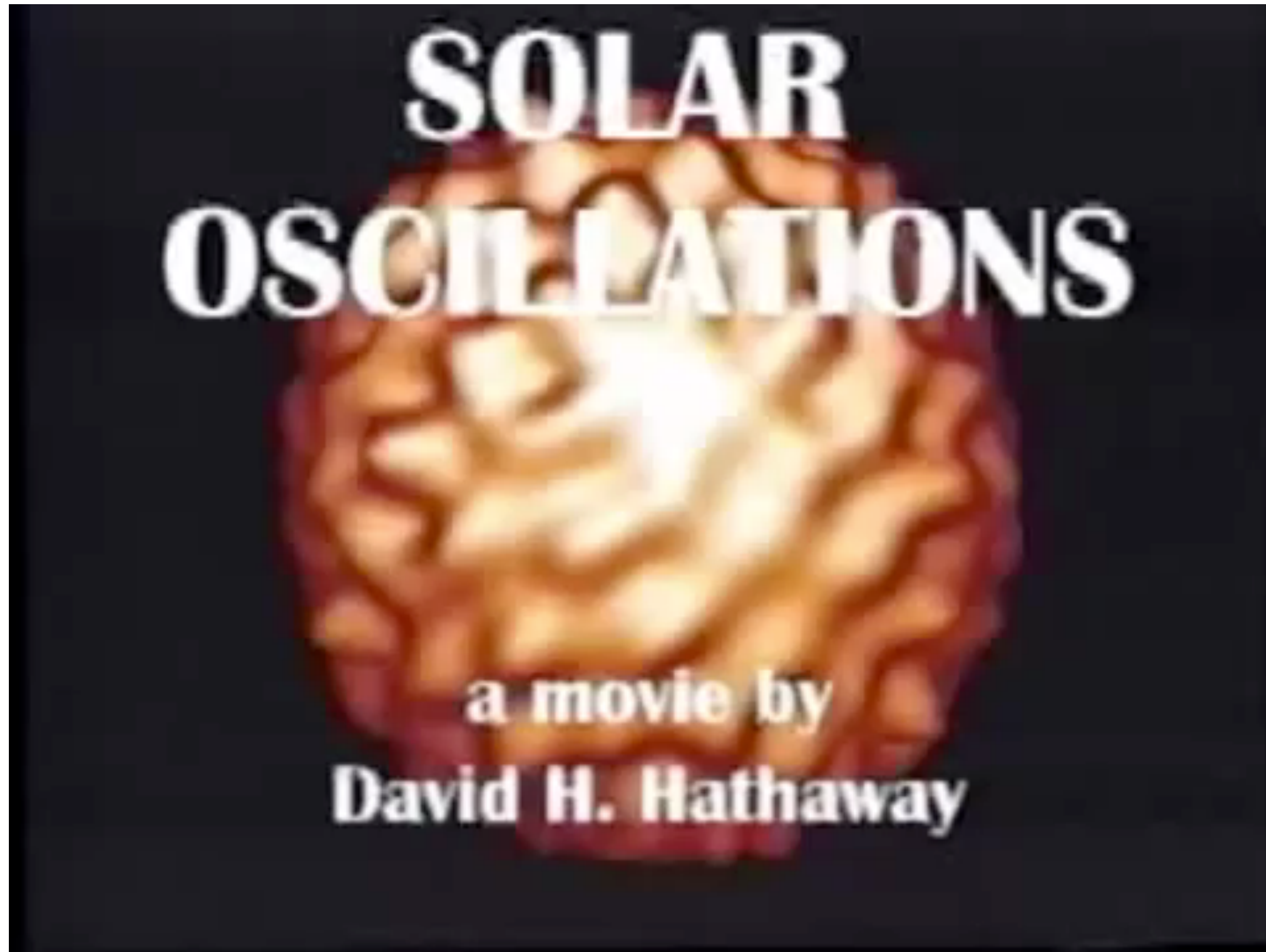


How light escapes the Sun

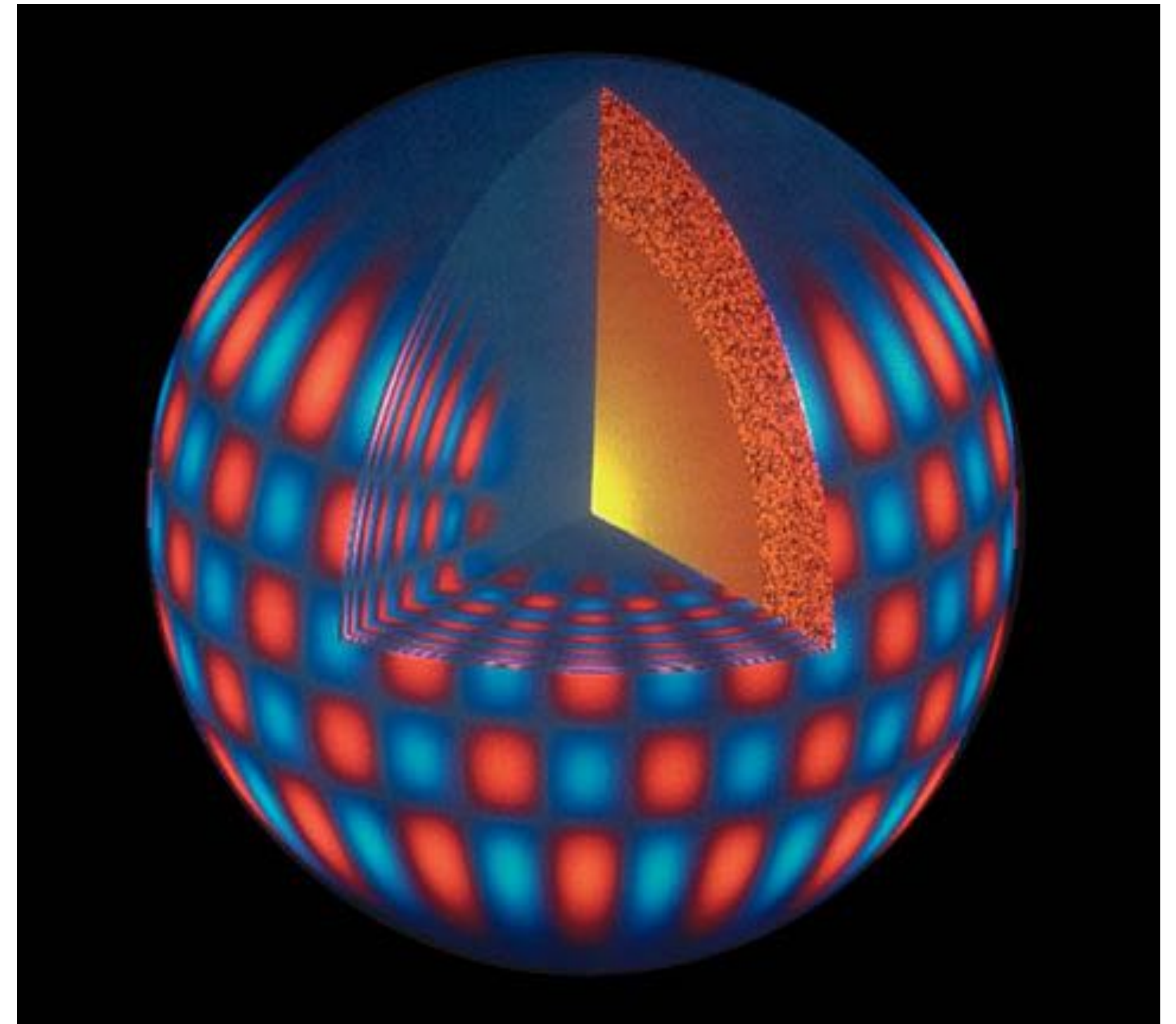


https://www.youtube.com/watch?v=W_Scoj4HqCQ

We learn about the Sun's interior from “sunquakes”: helioseismology

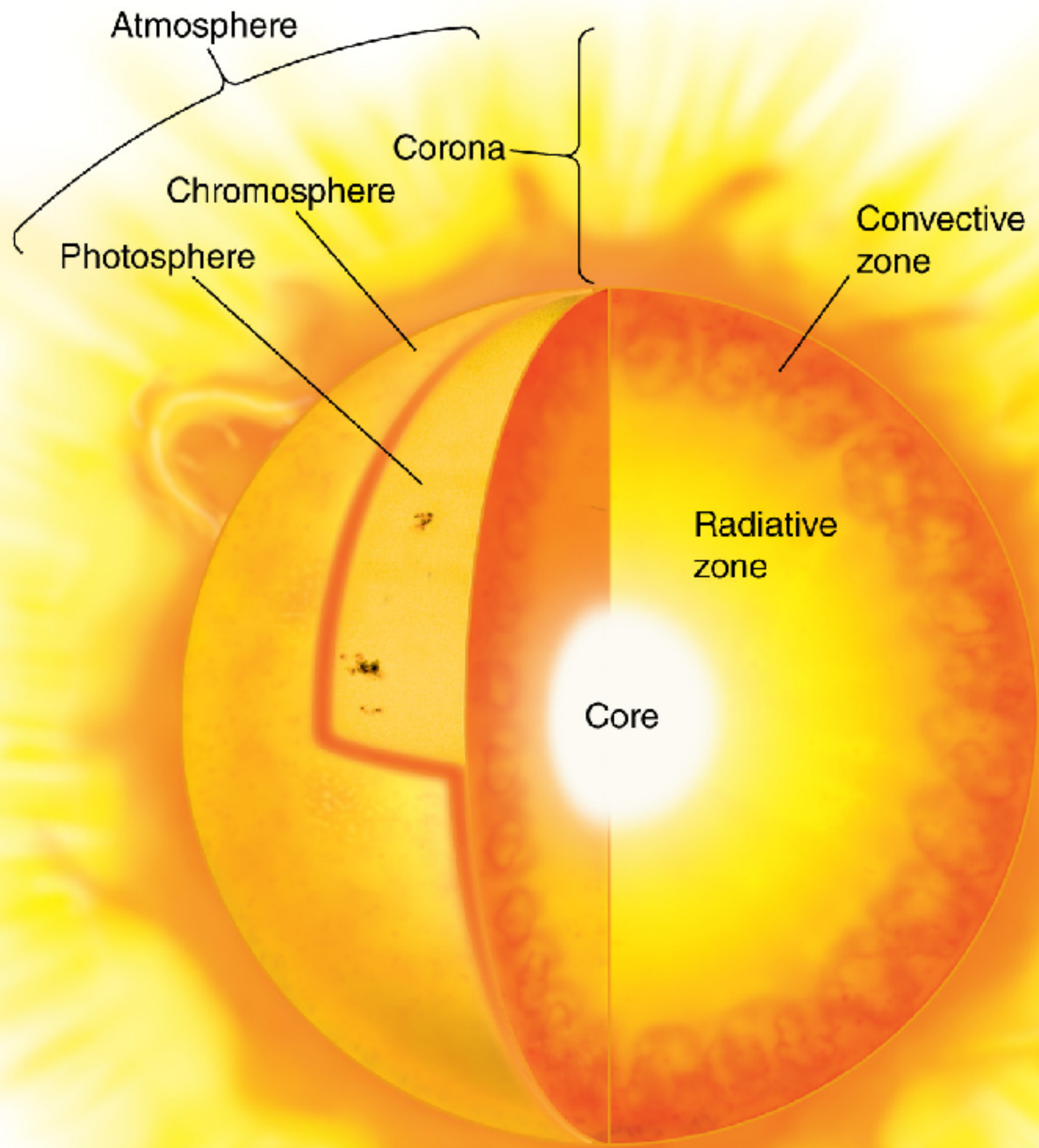
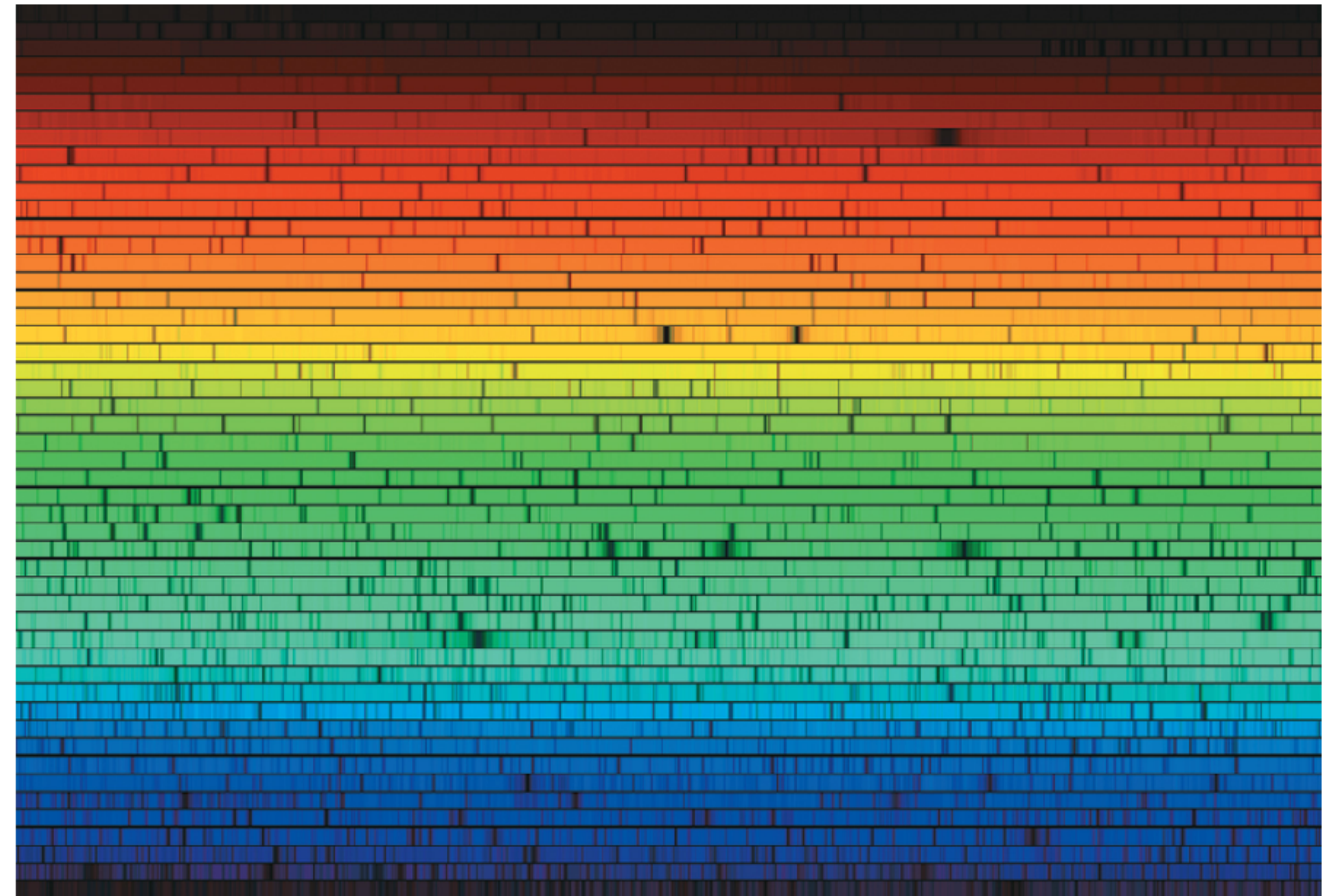


<http://www.youtube.com/watch?v=YxUsr4vp3yM>

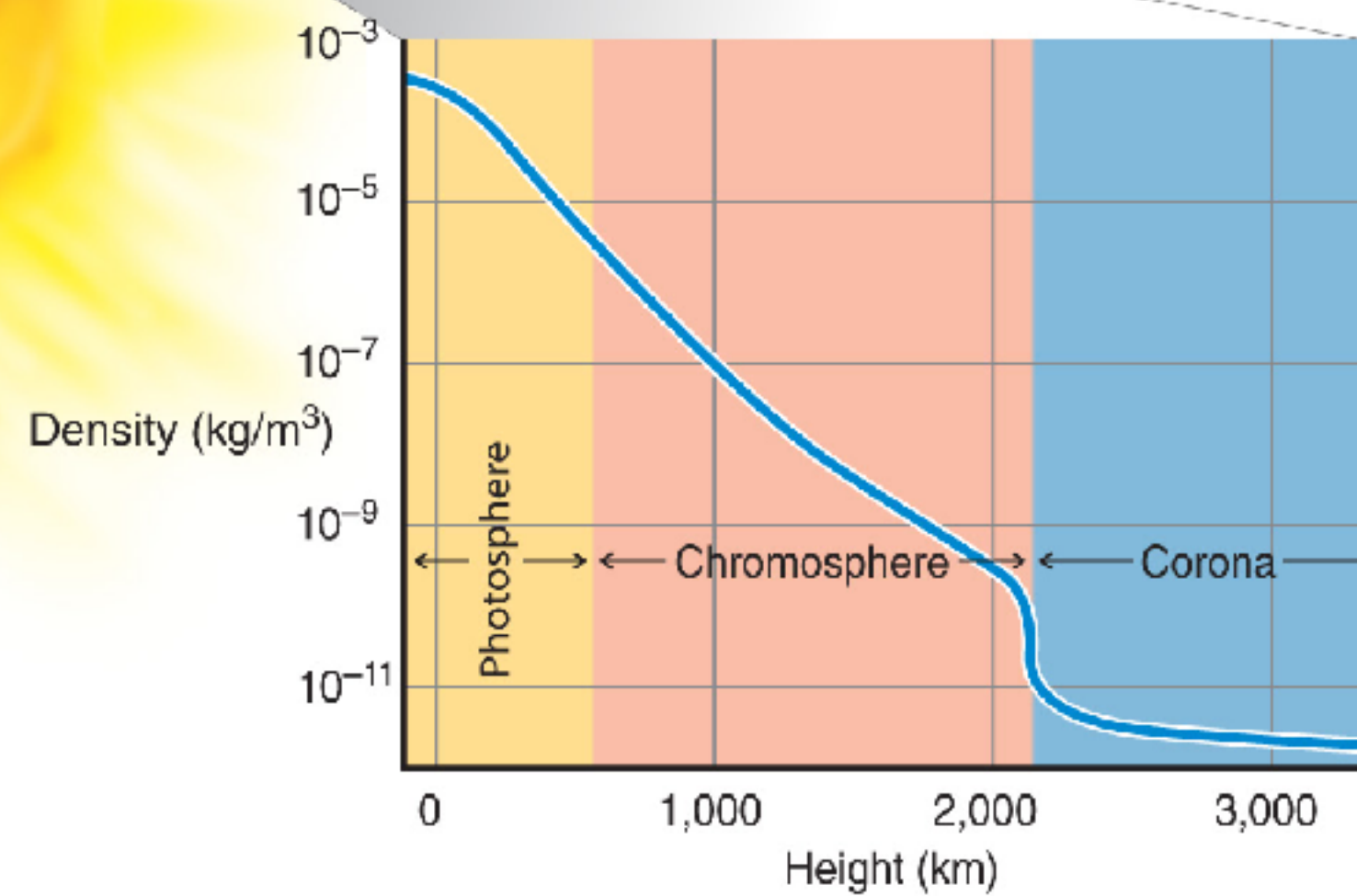
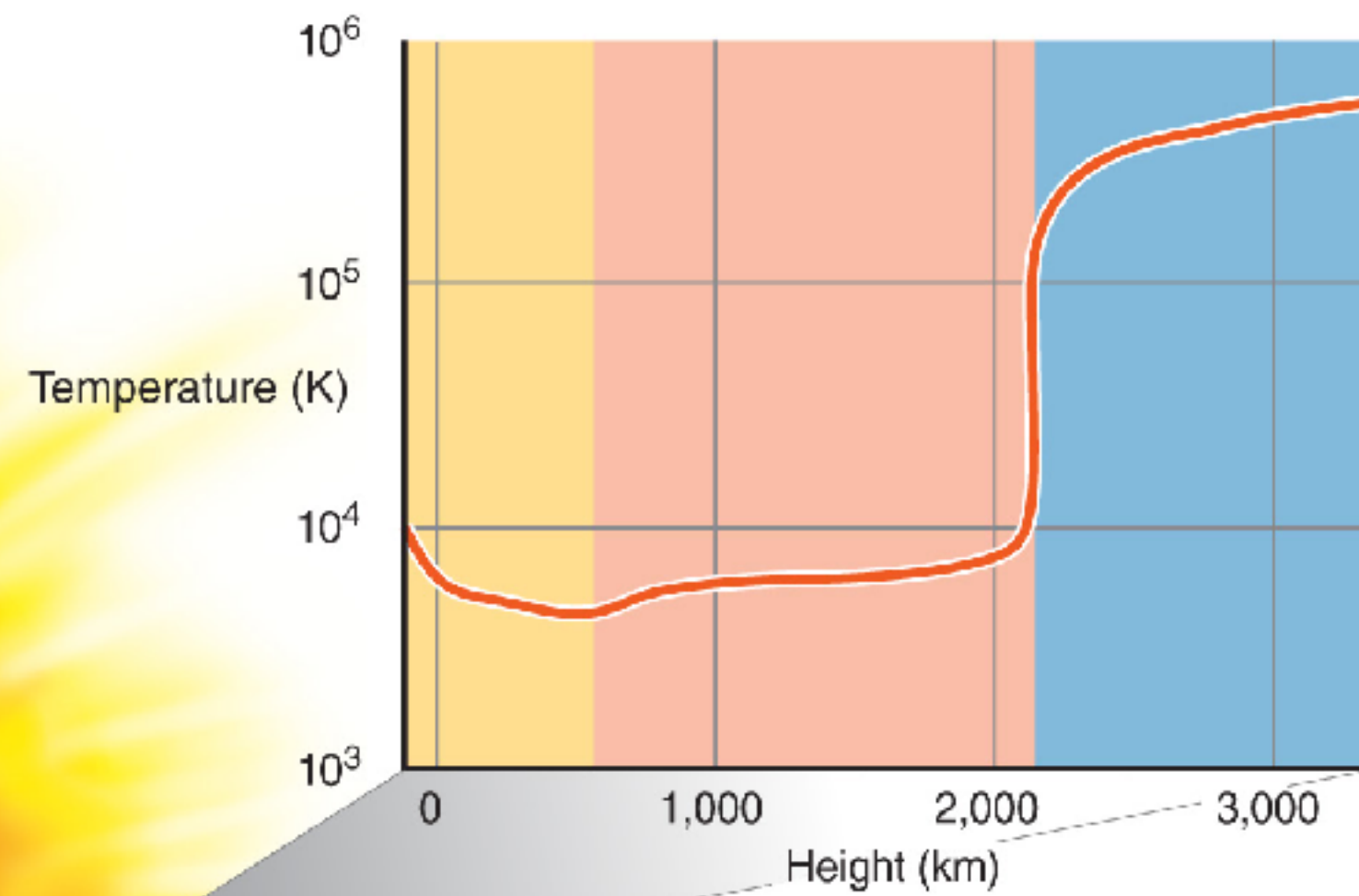
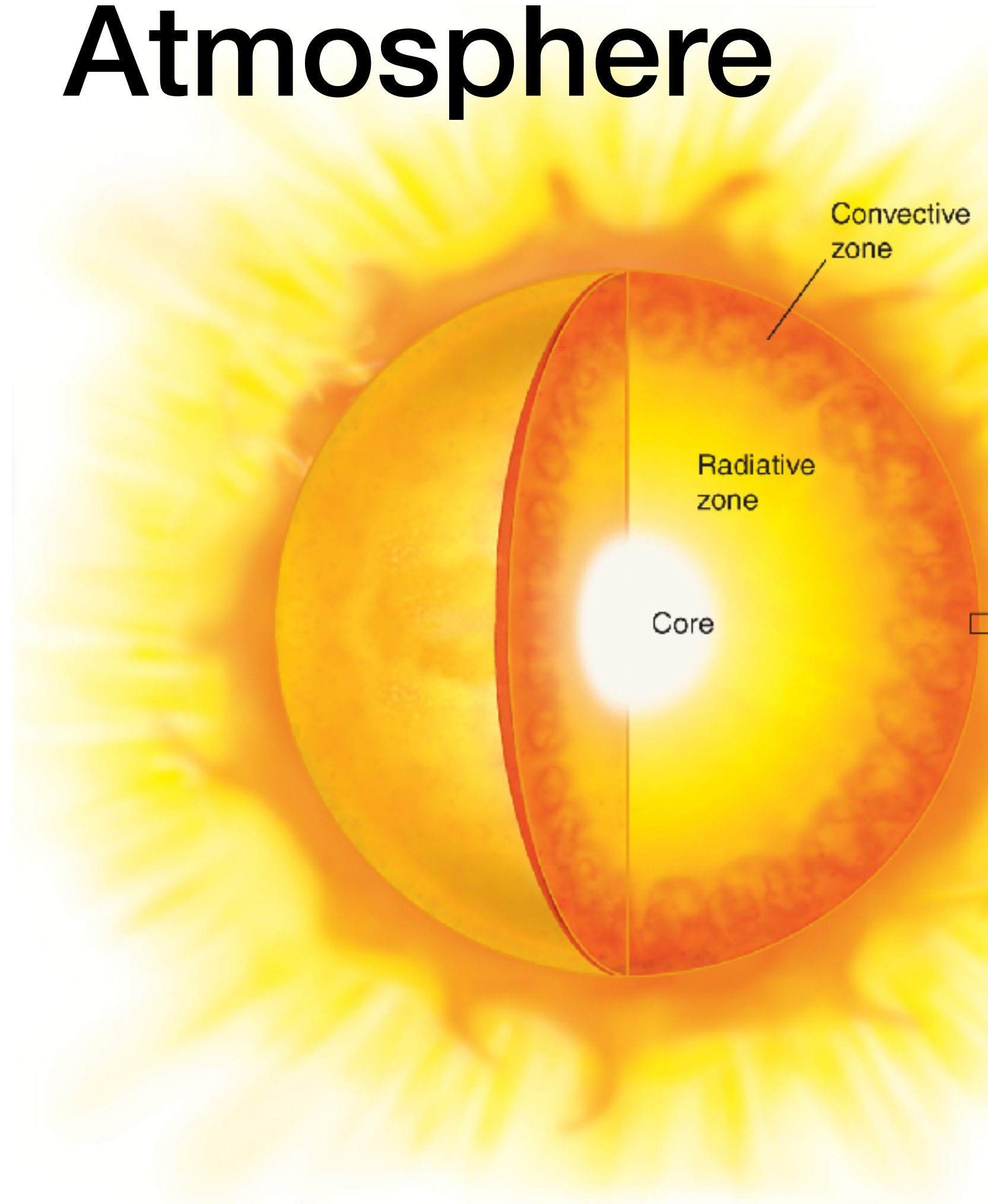


Zones of the Sun

High Resolution Solar Spectrum

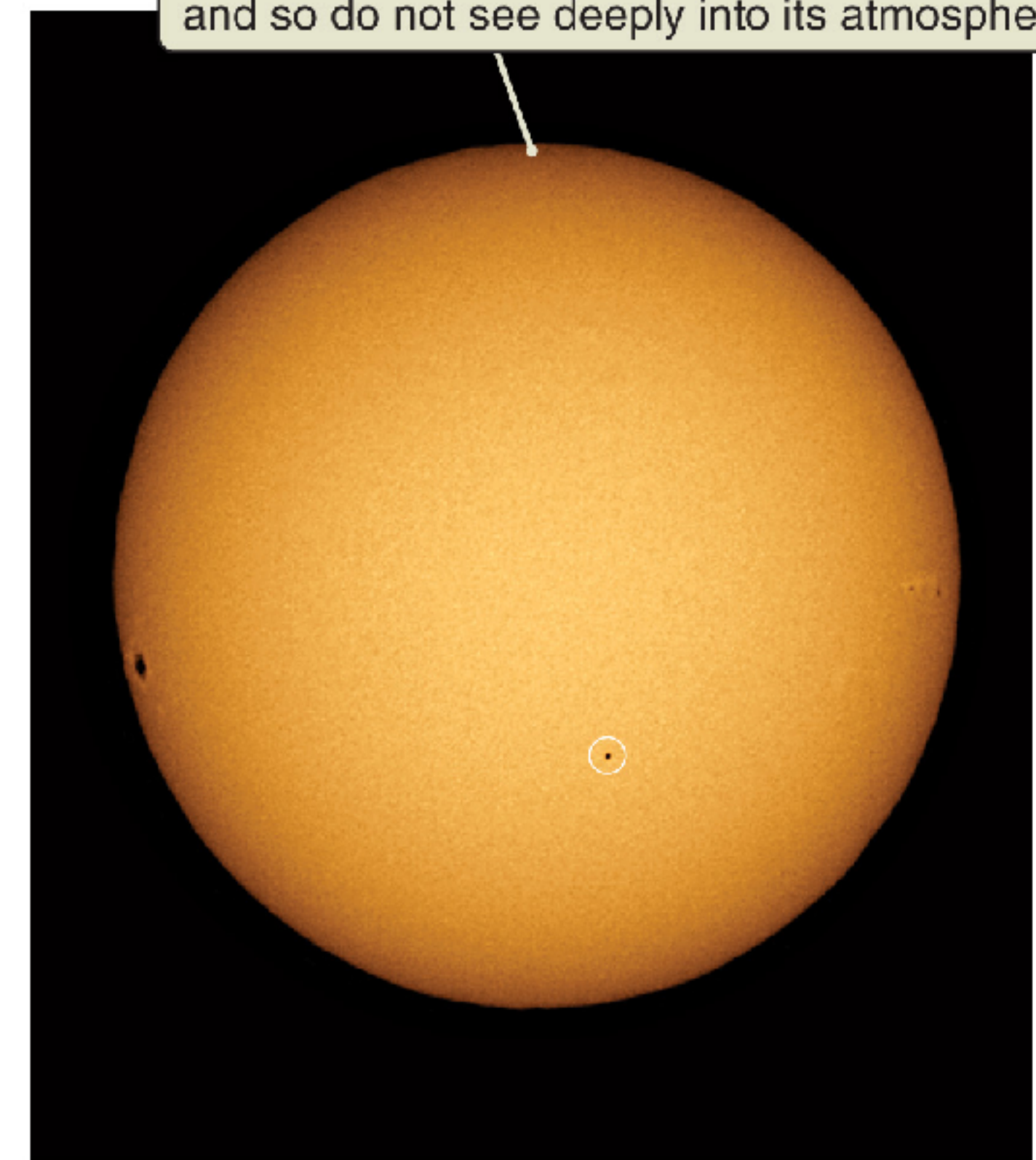


Sun's Atmosphere



Limb Darkening

The Sun is "limb darkened." It is dimmer near its edge because near its edge we see the Sun at a steep angle and so do not see deeply into its atmosphere.



G X U V I R