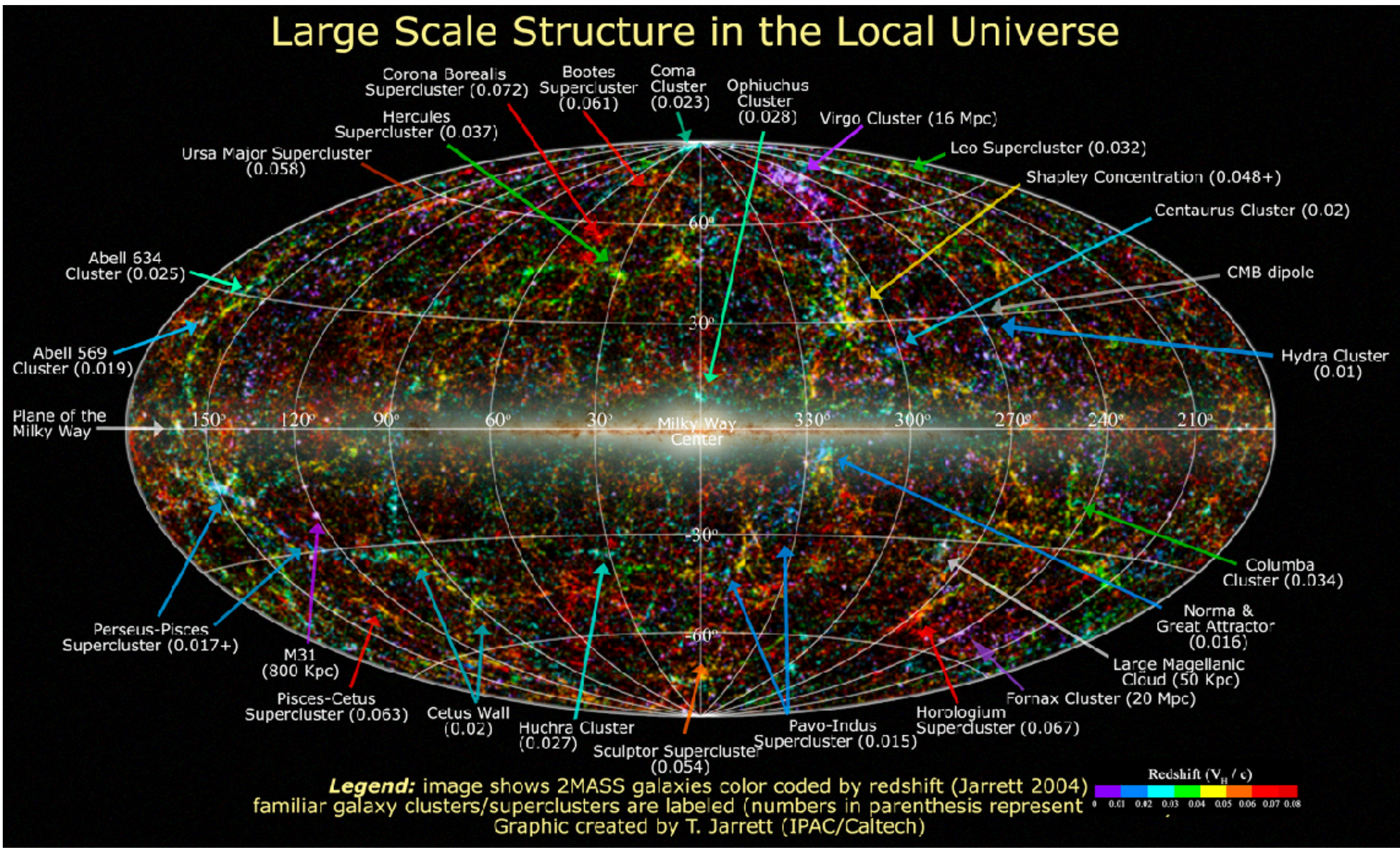
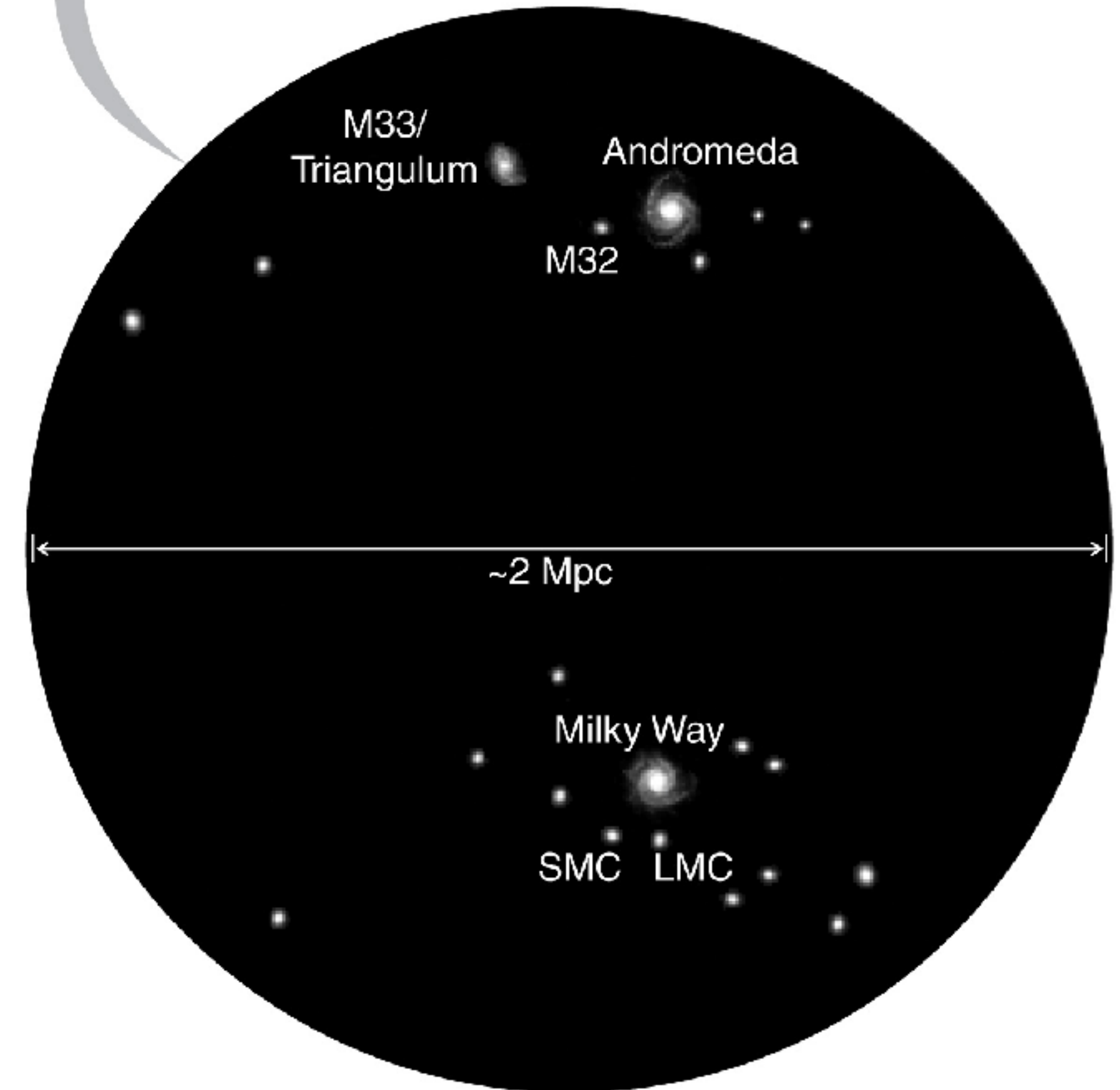
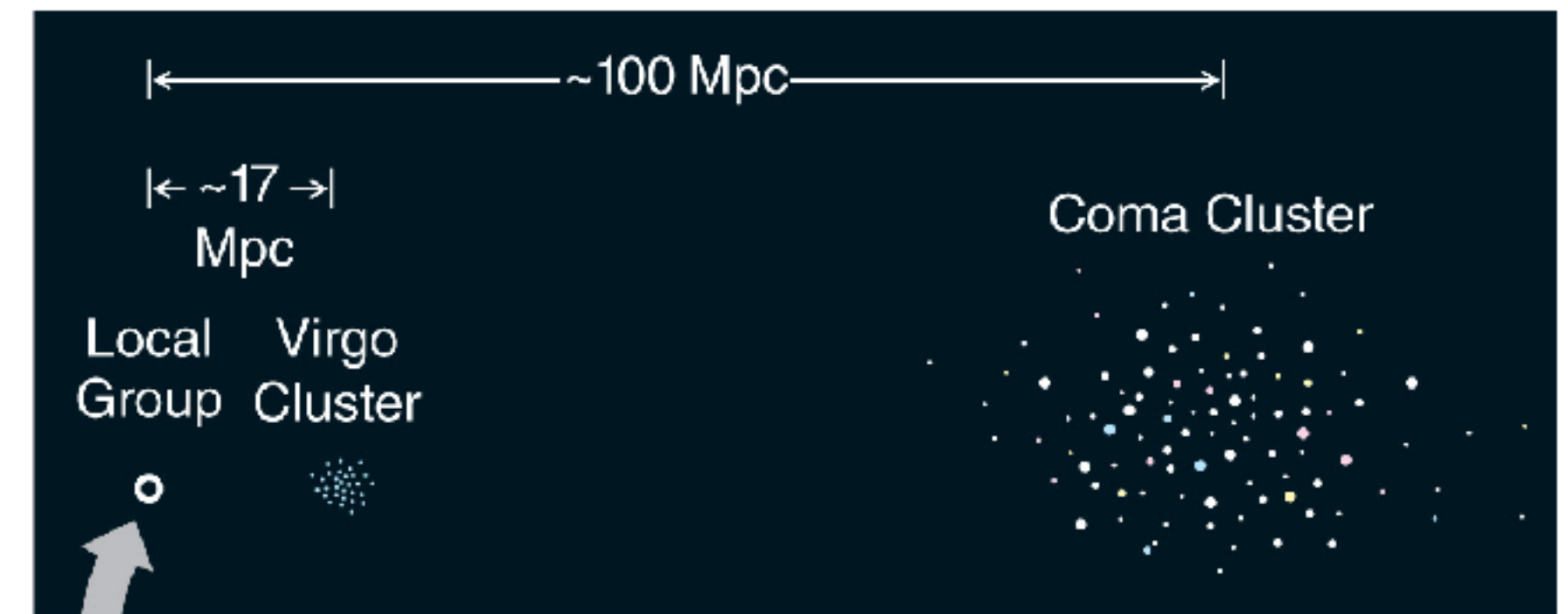
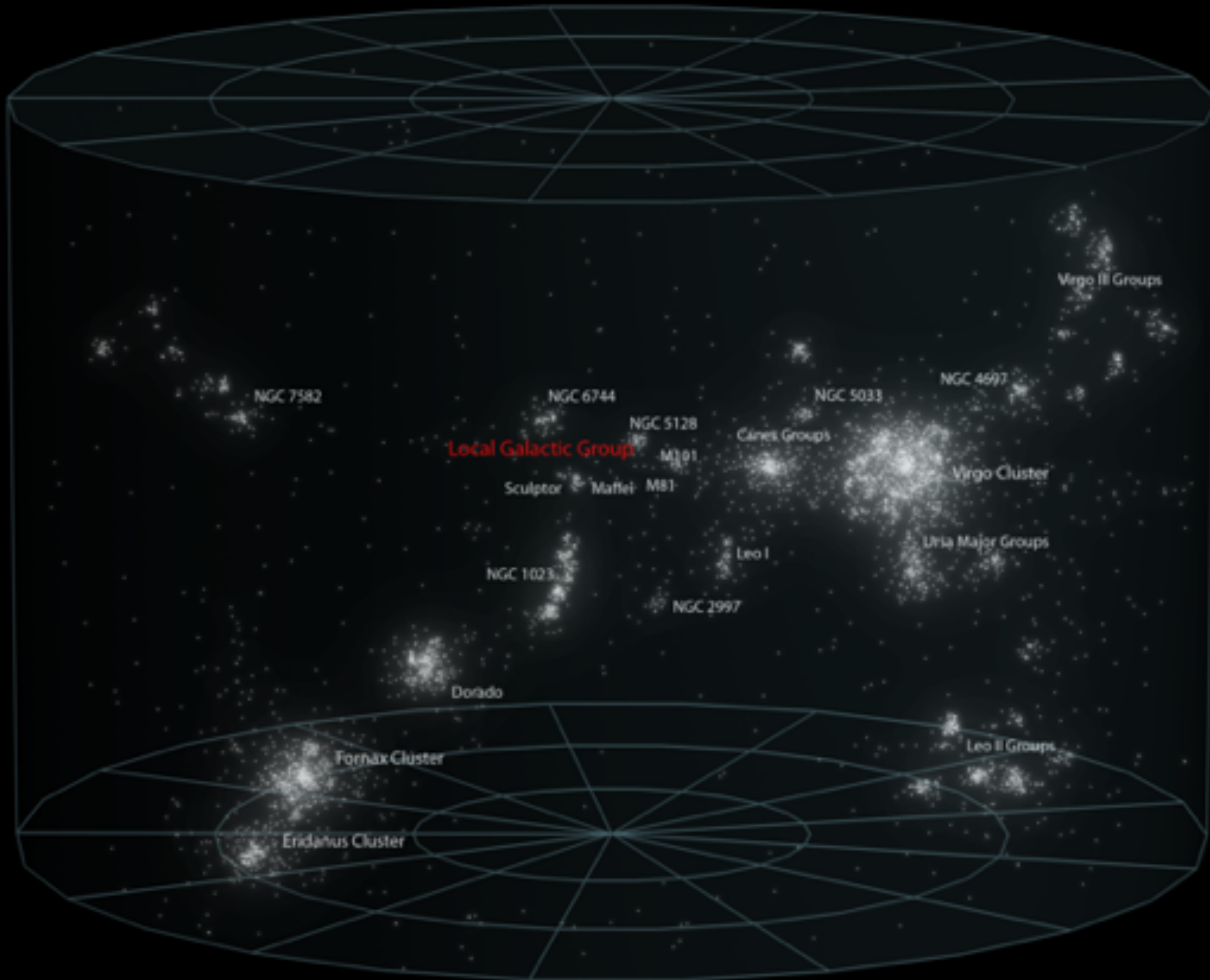


ASTR/PHYS 1060: The Universe

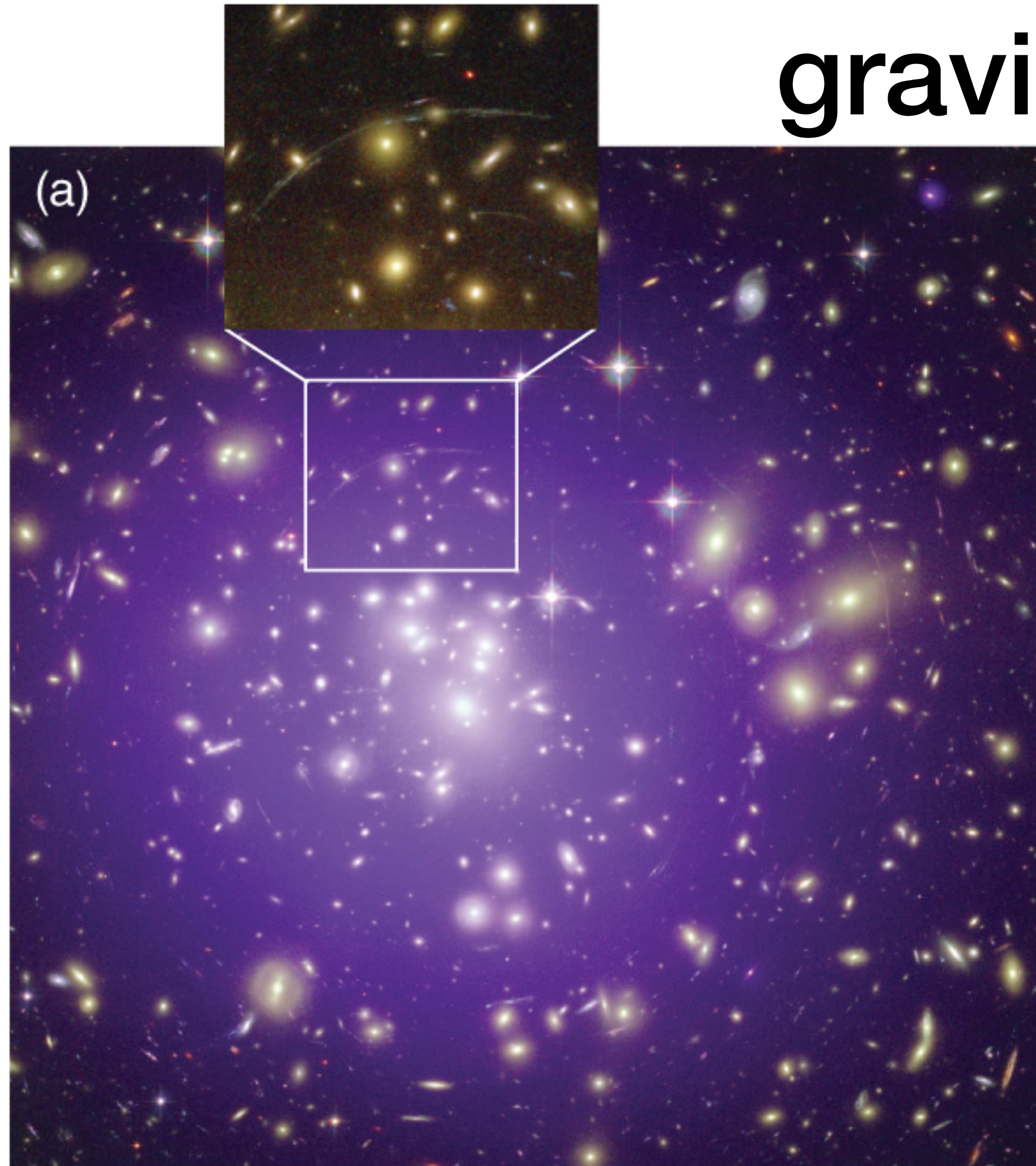
Chapter 17: Large Scale Structure



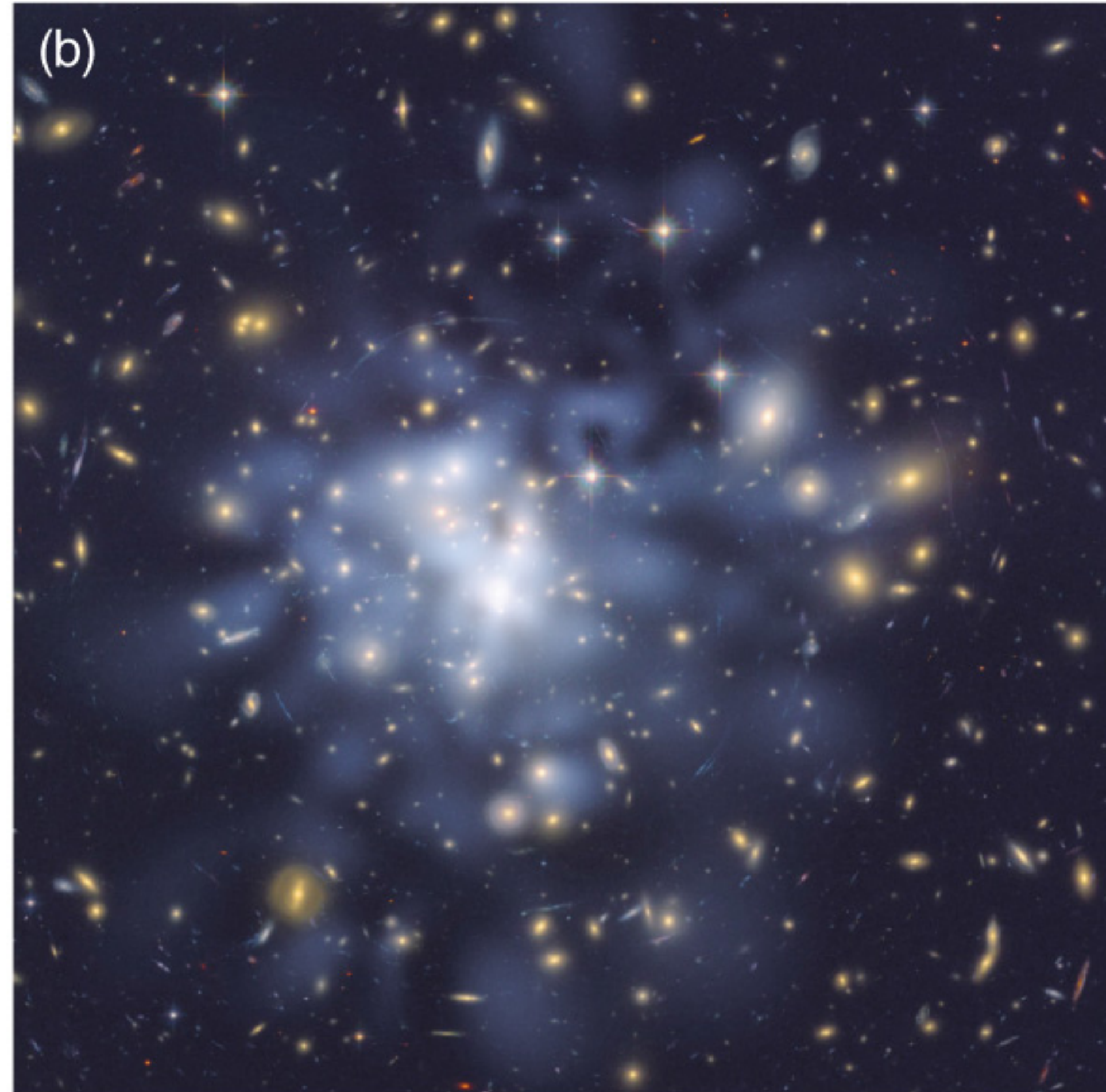
VIRGO SUPERCLUSTER



Galaxy Clusters: the largest gravitationally bound objects

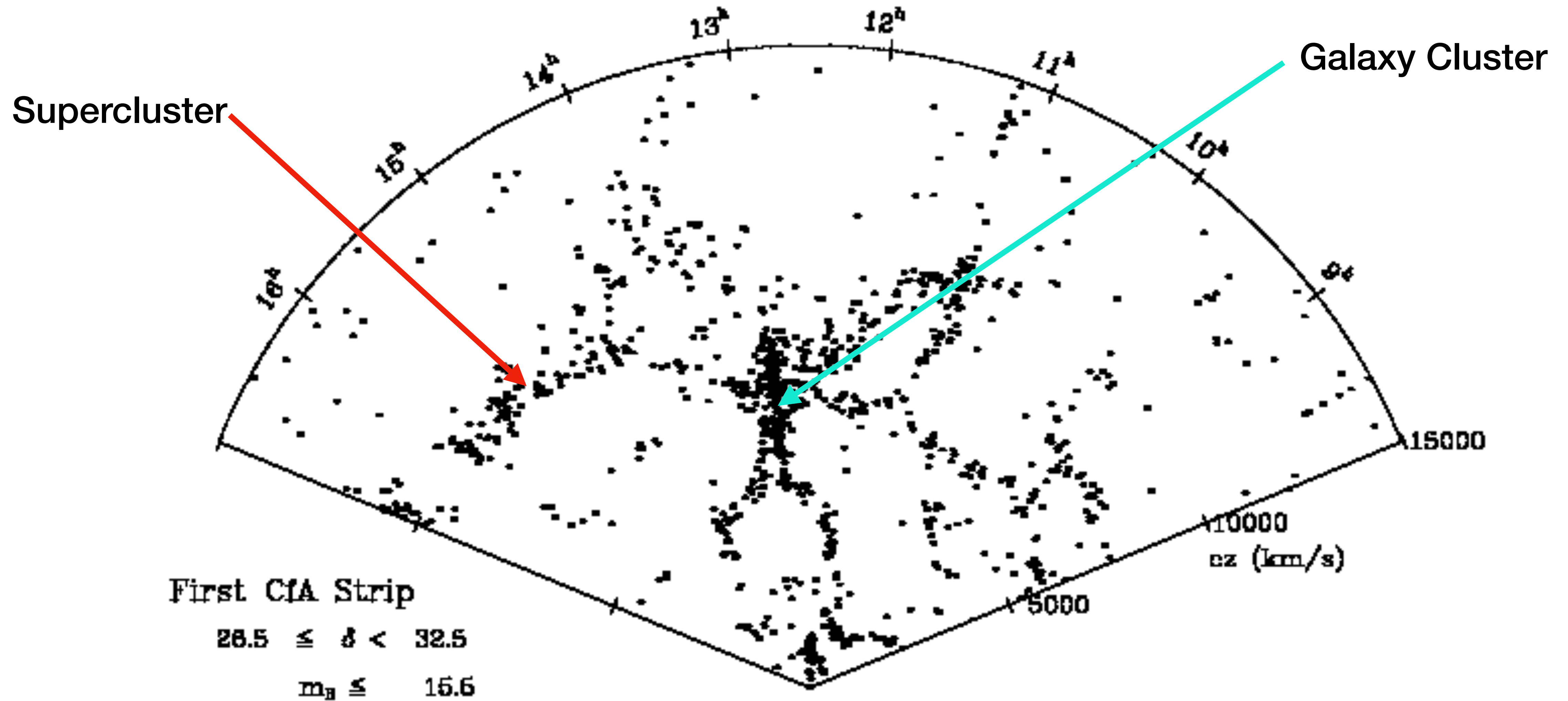


G X U V I R

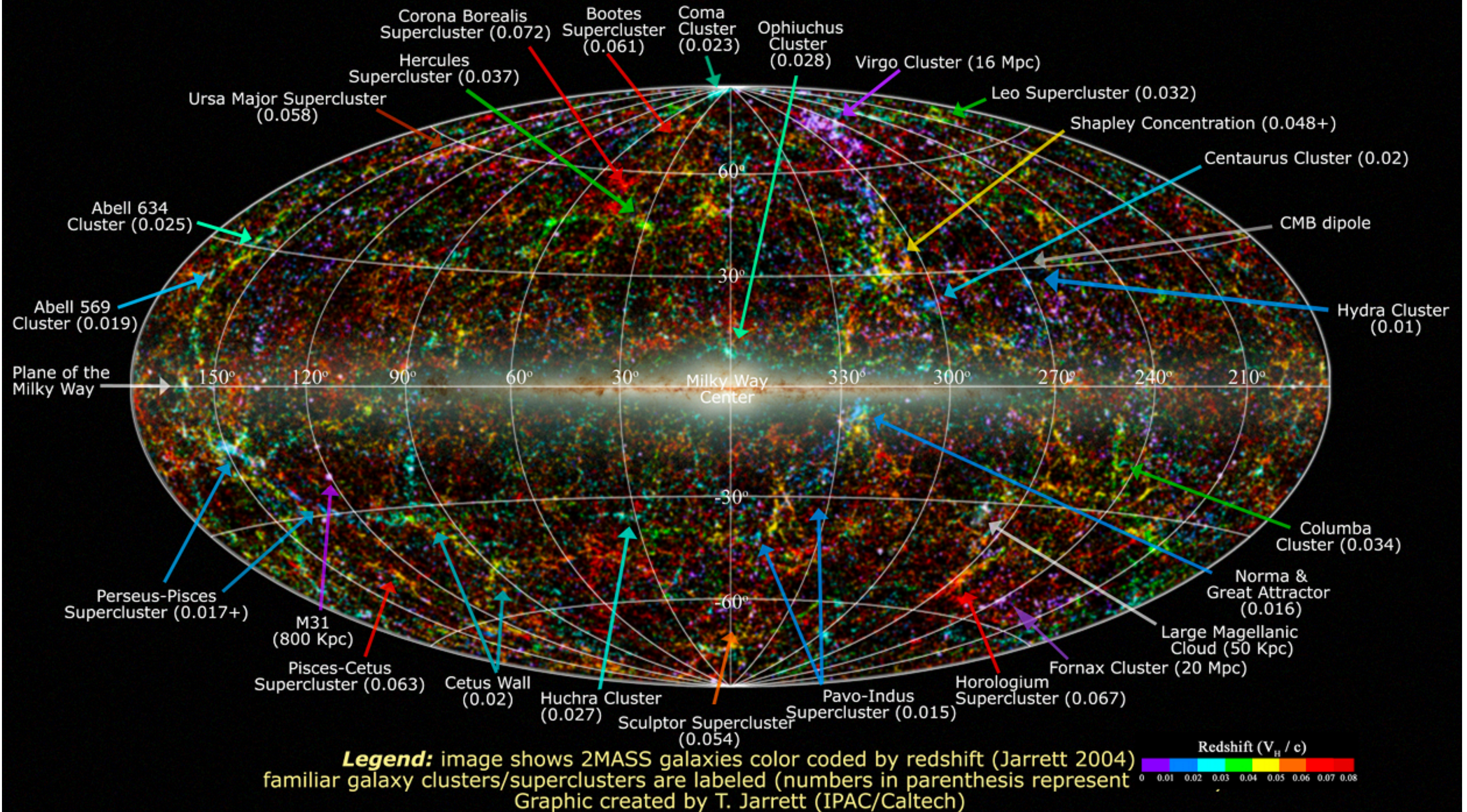


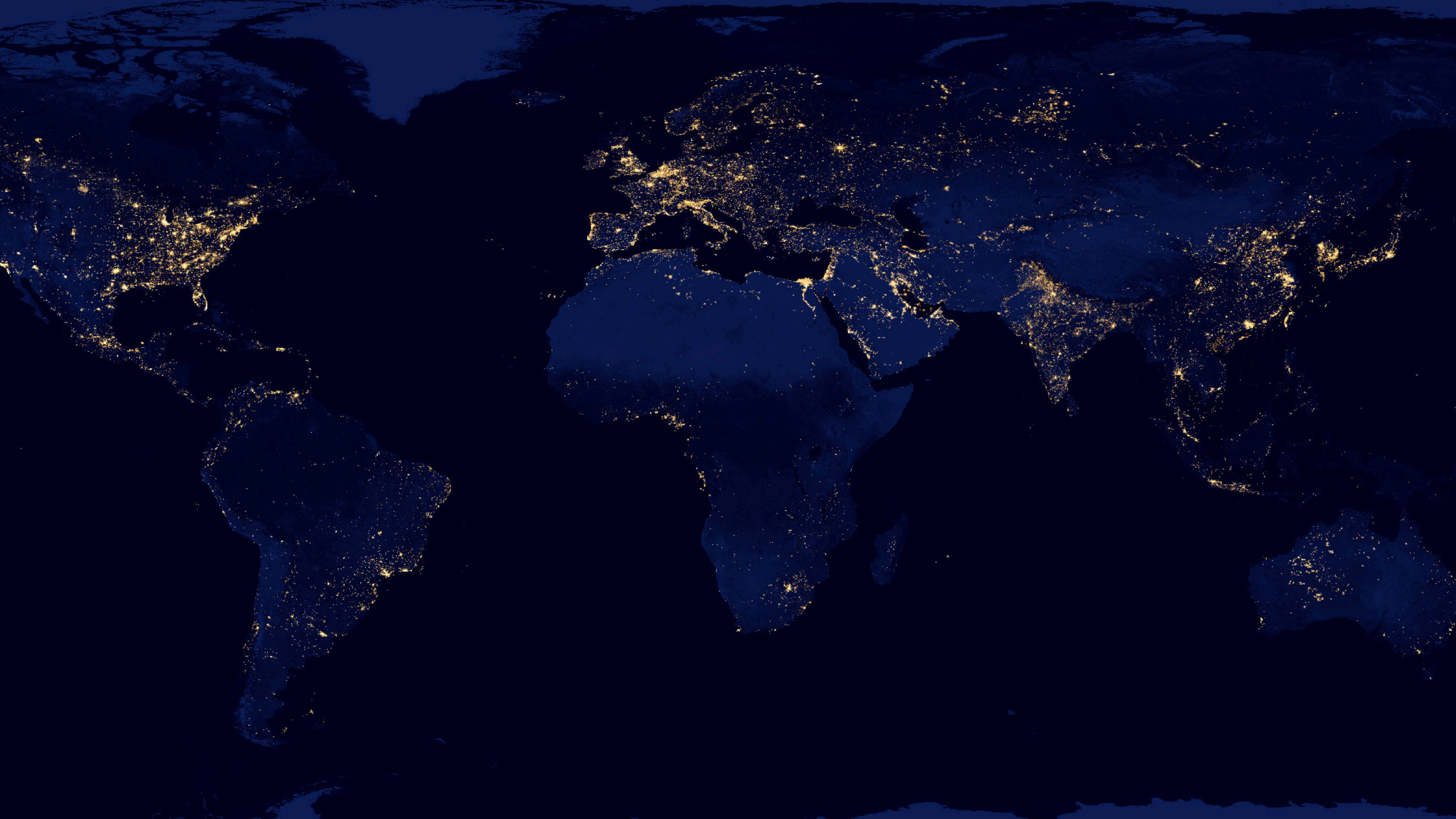
G X U V I R

Finger of God: the Coma Cluster



Large Scale Structure in the Local Universe





ASTR/PHYS 1060: The Universe

Chapter 17: Large Scale Structure

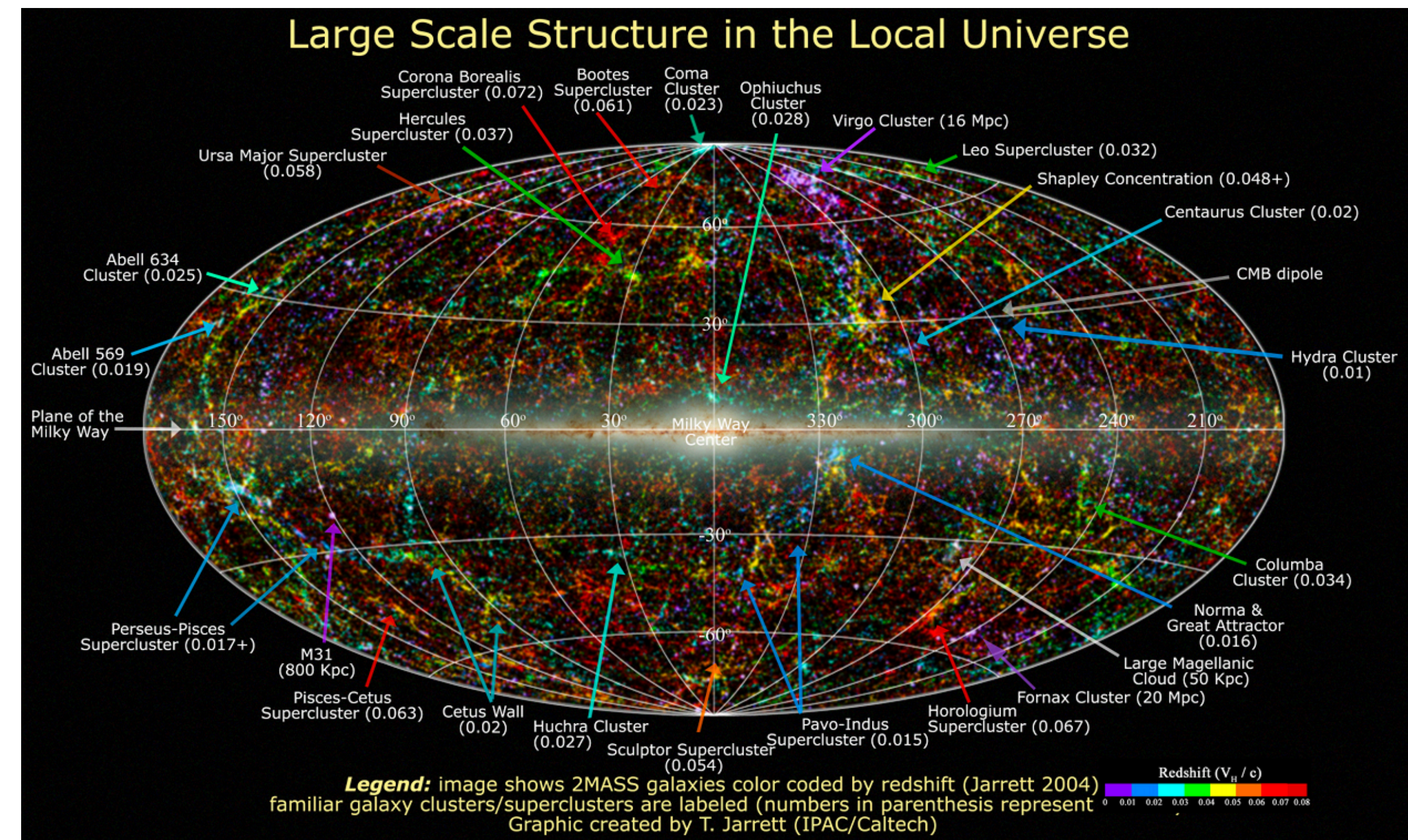
Current grade is ~4pts higher than what's in Canvas

Last Reading Assignment due tomorrow

Midterm 2 up front

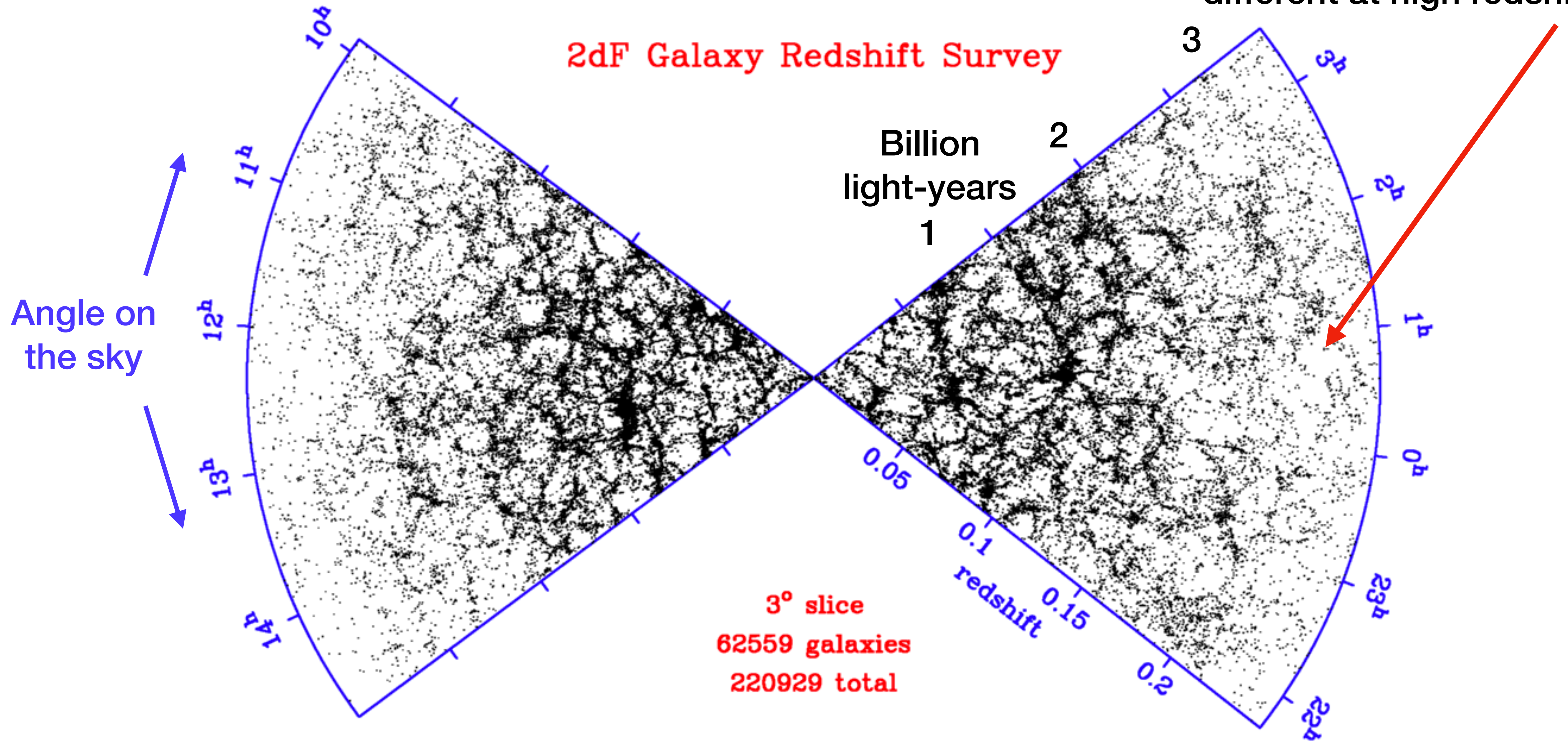
All EC due December 5th

Check your grades in Canvas!



Galaxy Surveys

Why does the pattern look different at high redshifts?



Making the “galaxy seeds” with inflation



ultra-tiny
quantum
fluctuations

Time
→



become...

large lumps seen in cosmic
microwave background

**Process is random
and becomes
“non-linear,” need
to simulate this
growth with
computers**

traveling through “slices” of a
simulated universe

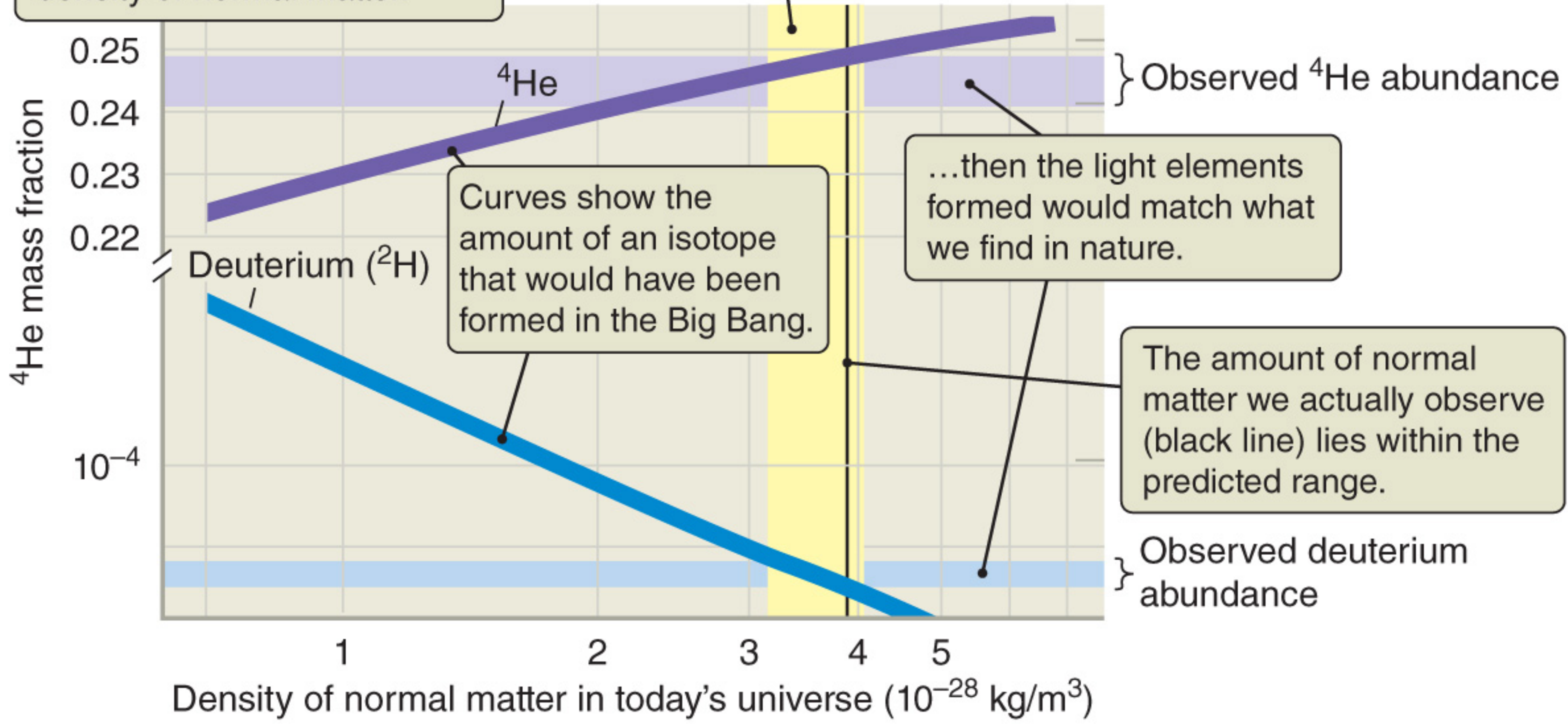
What is dark matter?

**Must it be different than the
matter you and I are made of?**

The amounts of different elements produced in the Big Bang were determined by the density of normal matter.

If the density of normal matter in today's universe lies in this range...

amount of normal matter is not enough to explain the gravity we observe



Curves show the amount of an isotope that would have been formed in the Big Bang.

...then the light elements formed would match what we find in nature.

The amount of normal matter we actually observe (black line) lies within the predicted range.

Observed deuterium abundance

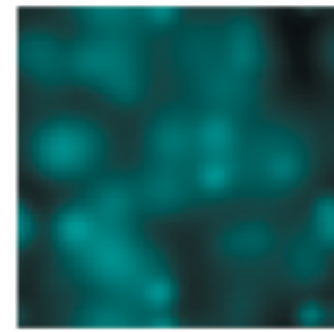
Observed ${}^4\text{He}$ abundance

How do structures grow?

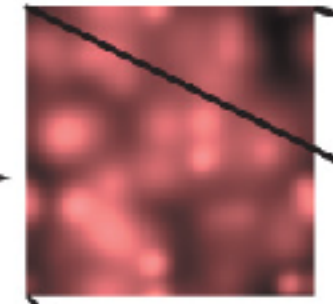
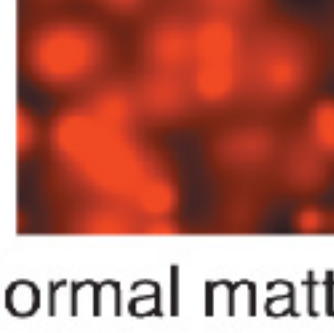
- A) Galaxies form (medium size), then stars (small size), then clusters (large size)
- B) Stars form (small size), then clusters (large size), then galaxies (medium size)
- C) Clusters form (large size), then galaxies (medium size), then stars (small size)
- D) Stars form (small size), then galaxies (medium size), then clusters (large size)

1 Clumps of dark matter, normal matter, and radiation emerge from the very early universe.

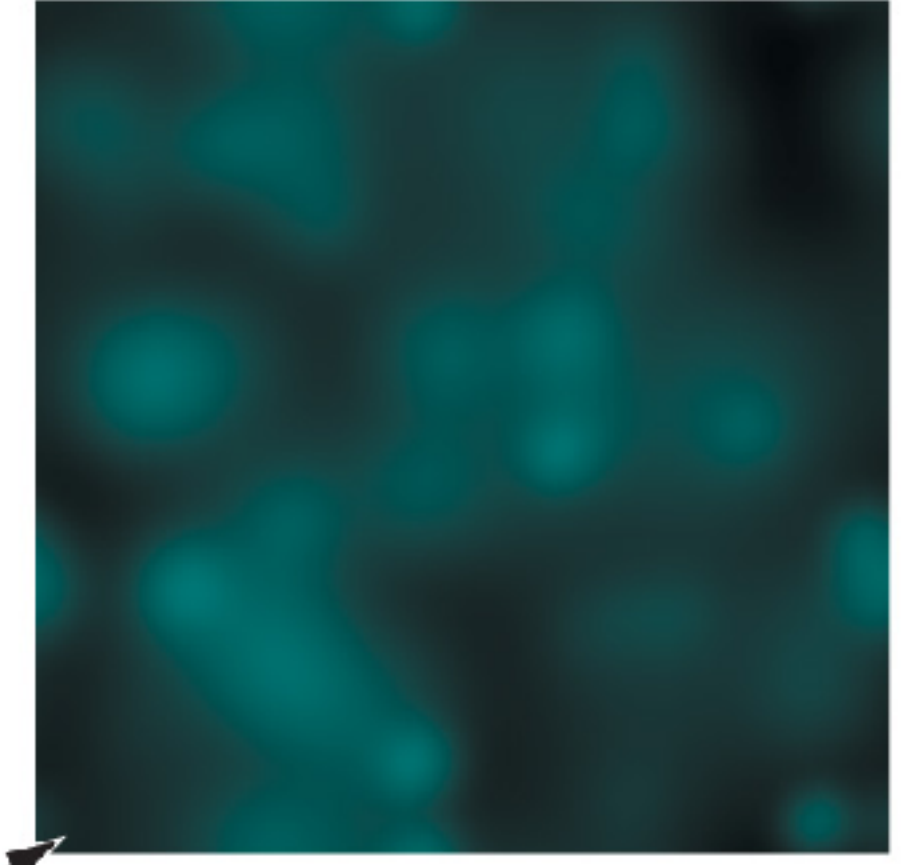
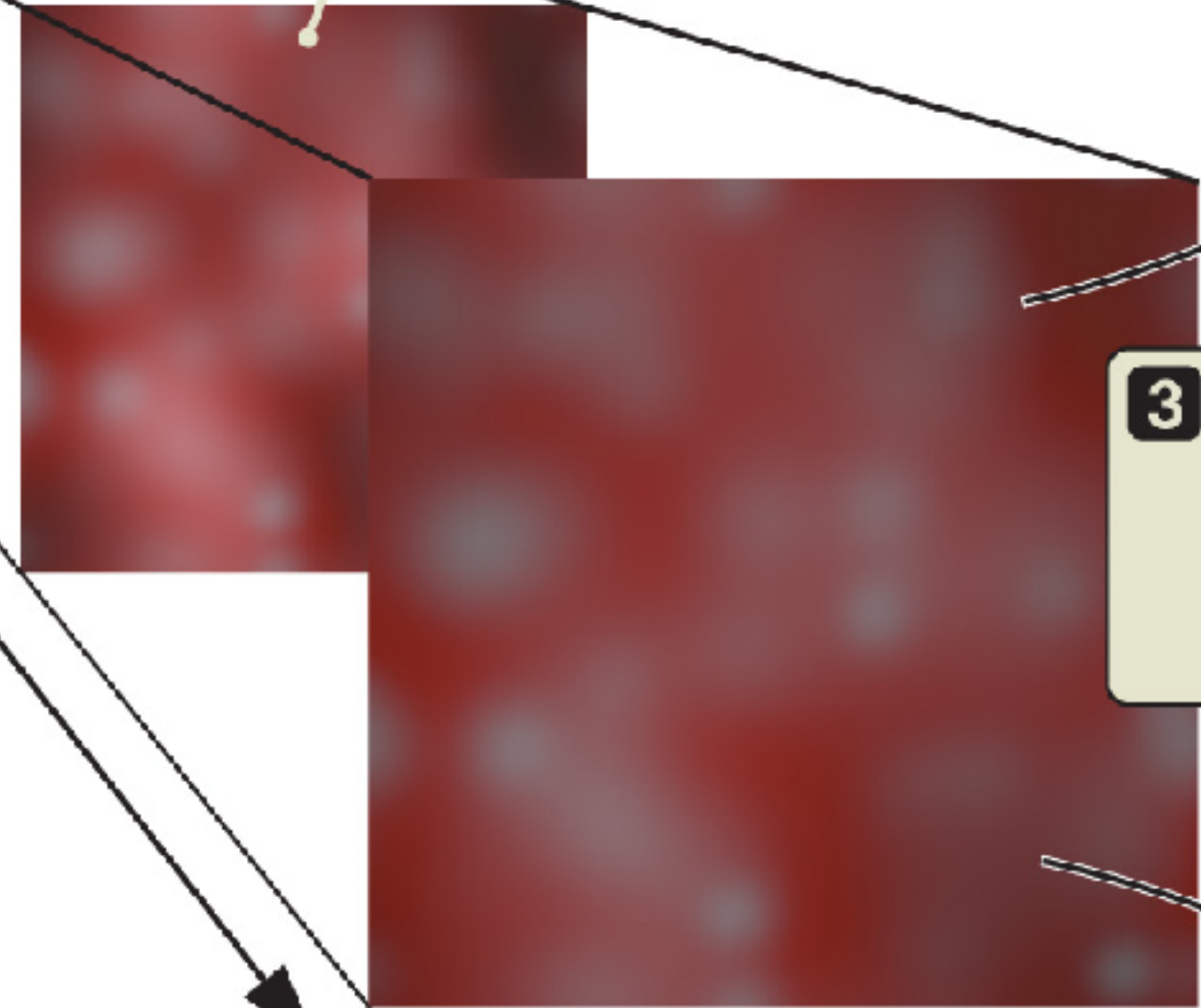
Dark matter



Normal matter and radiation



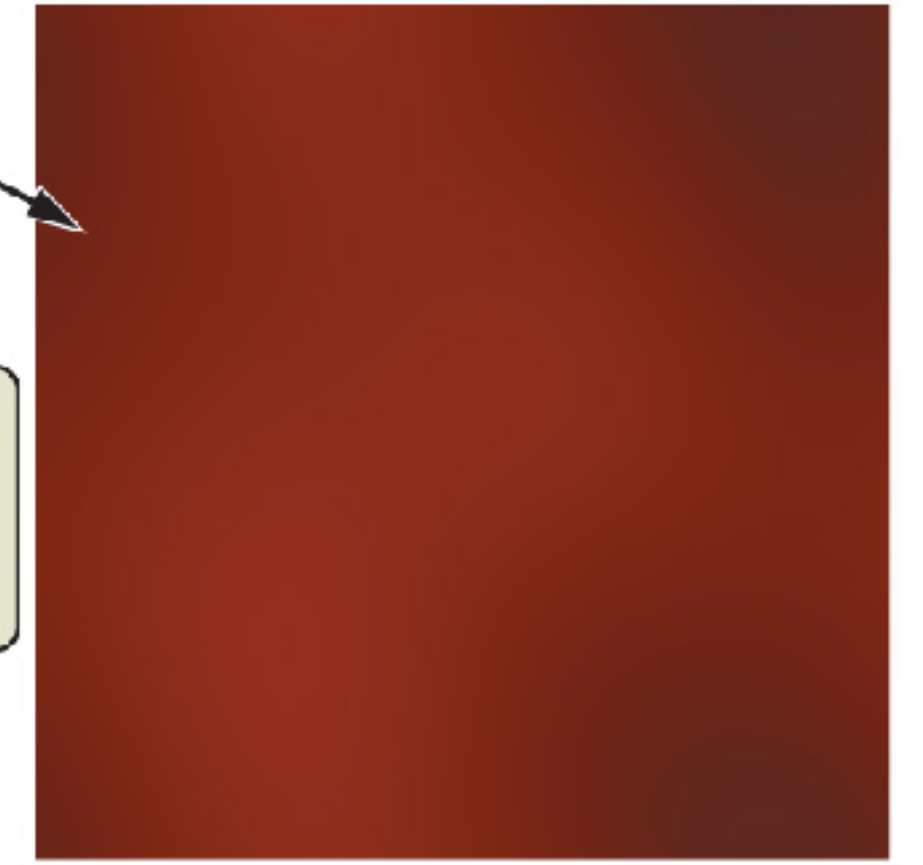
2 Pressure of radiation smooths out normal matter clumps as the universe expands...



Dark matter

3 ...leaving behind dark matter clumps in a far more uniform background of normal matter and radiation.

4 Very slight irregularities remain in the cosmic background radiation that we see.

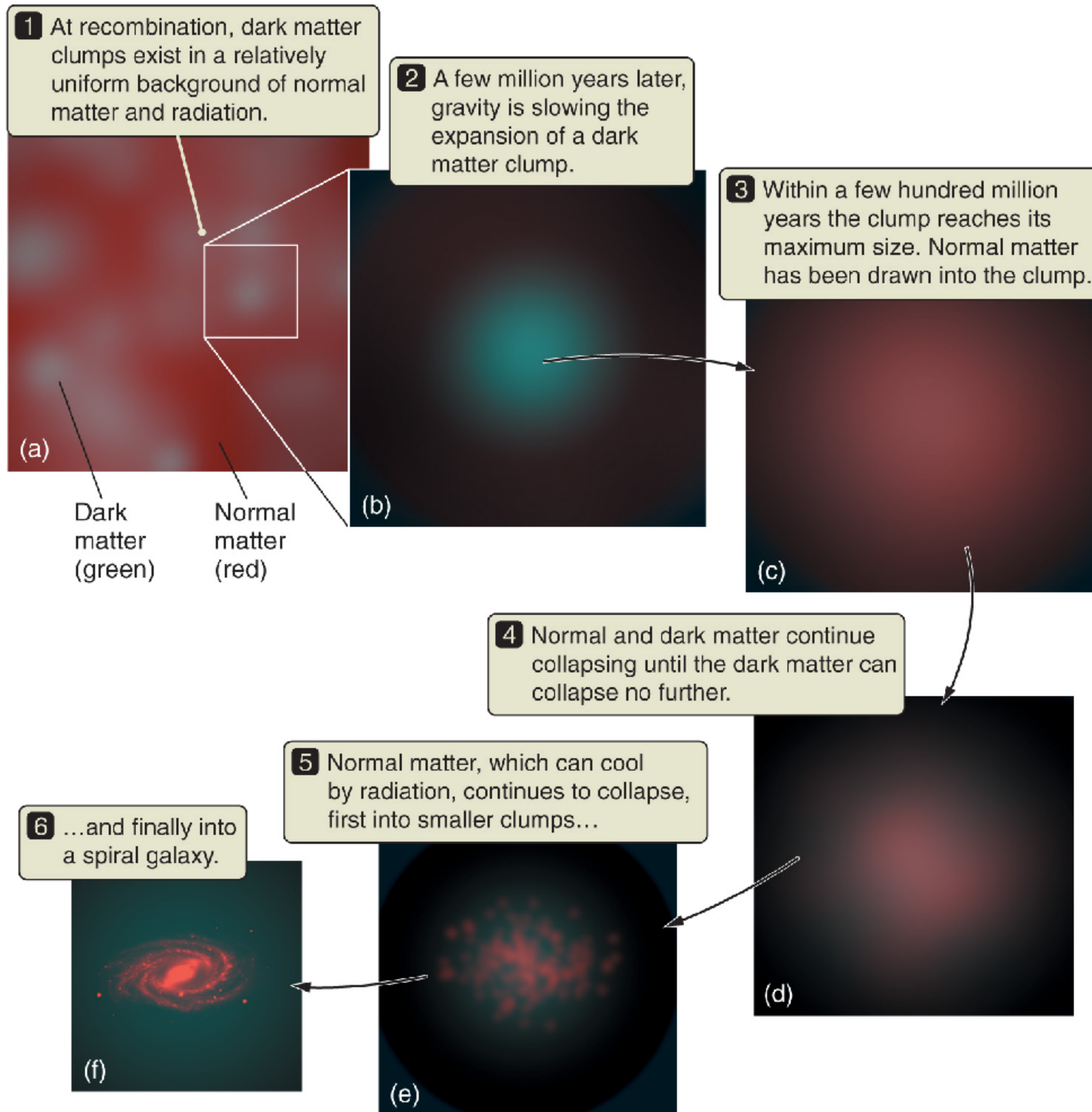


Normal matter and radiation

Time

Dark matter drives the formation of structures

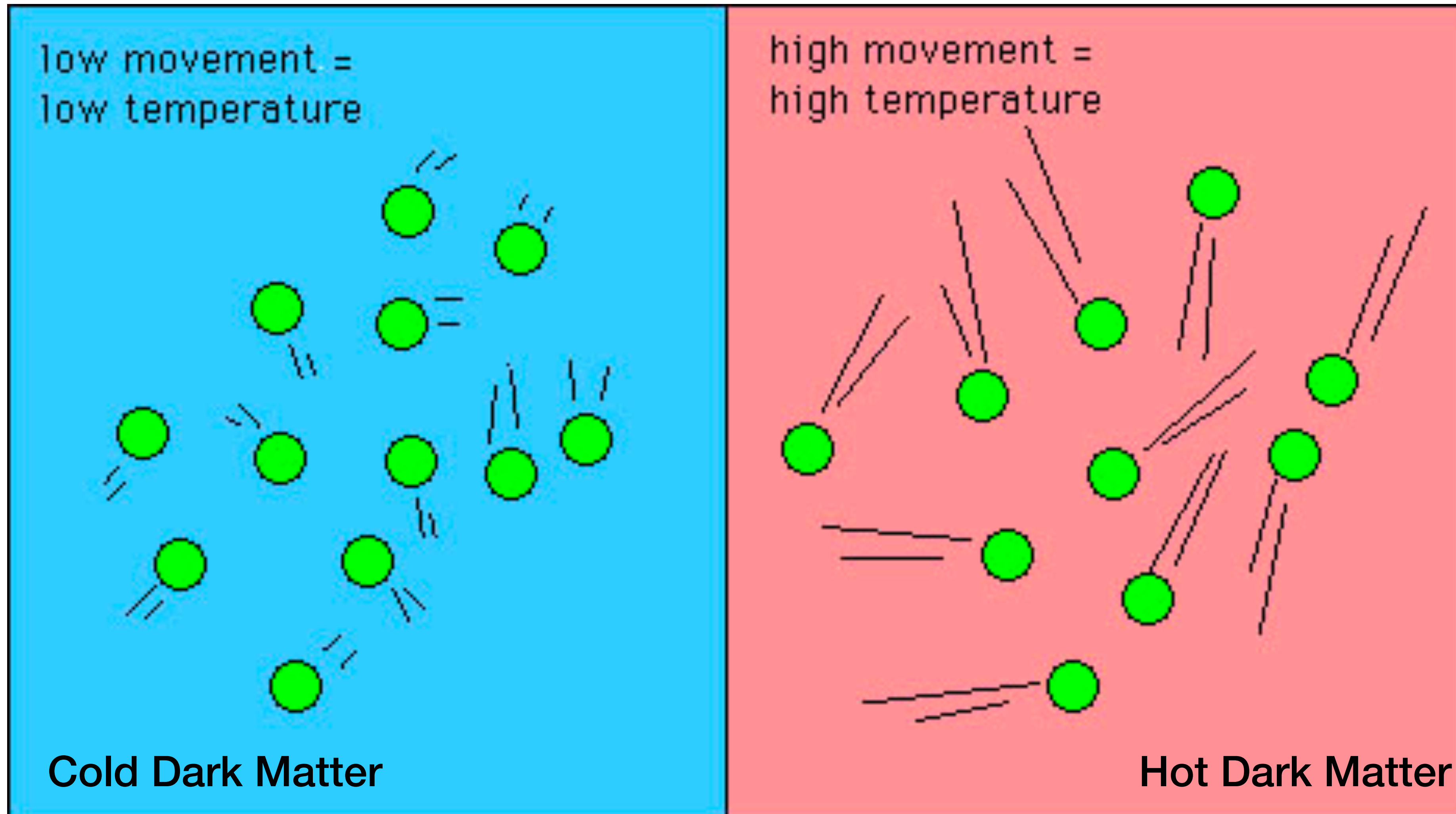
(the type of dark matter matters!)



Galaxies form inside dark matter halos, becoming more compact in centers than the dark matter

So, the distribution of galaxies is related to the distribution of dark matter that was able to collapse and form halos

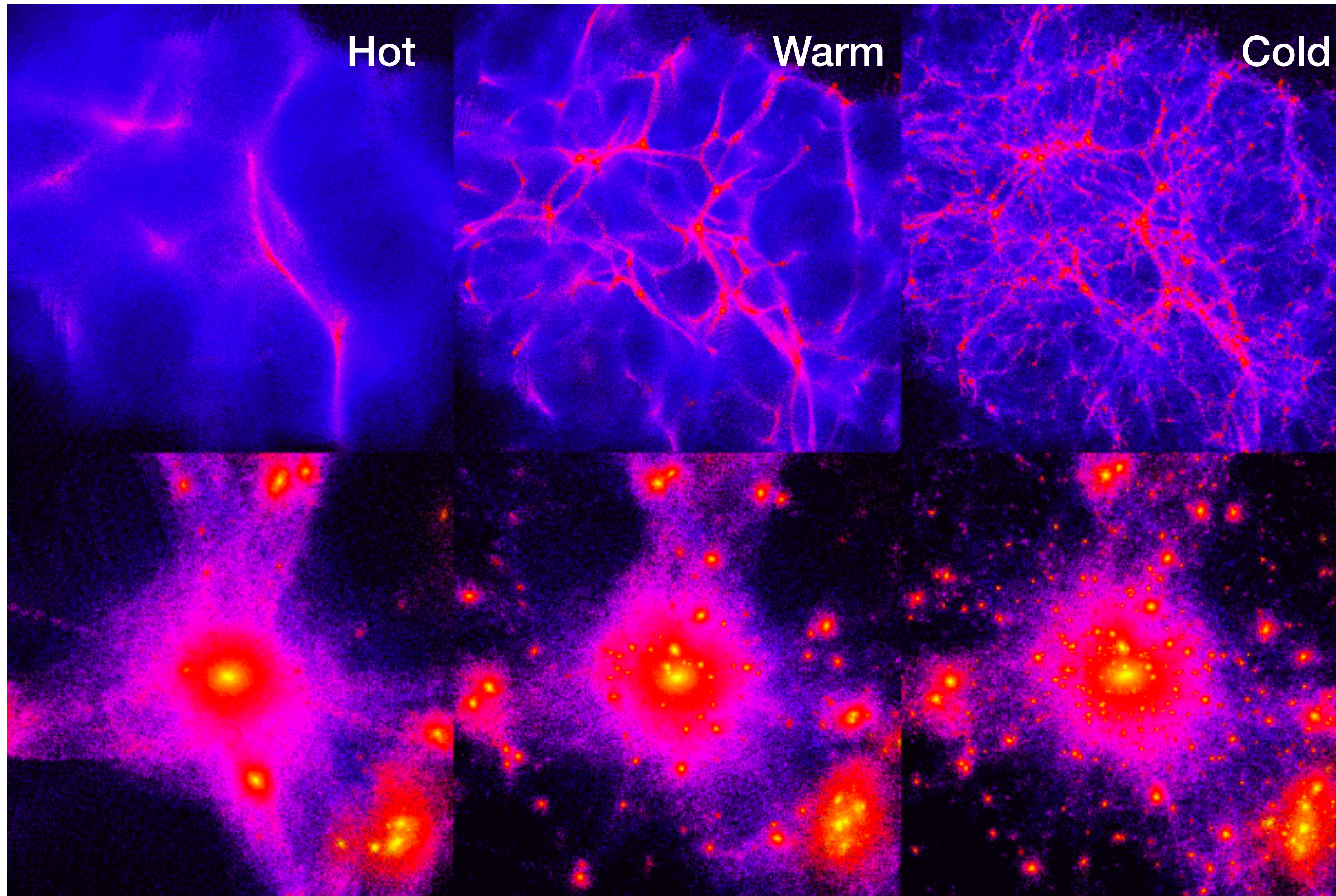
What is temperature?



How does the temperature of dark matter affect how it grows?

- A) Hot dark matter grows small structures first
- B) Cold dark matter grows small structures first
- C) Temperature doesn't affect the size scale that grows first

Temperature of the Dark Matter



velocity of particles
compared to the speed of
light

relativistic at time of collapse
(like neutrinos): hot

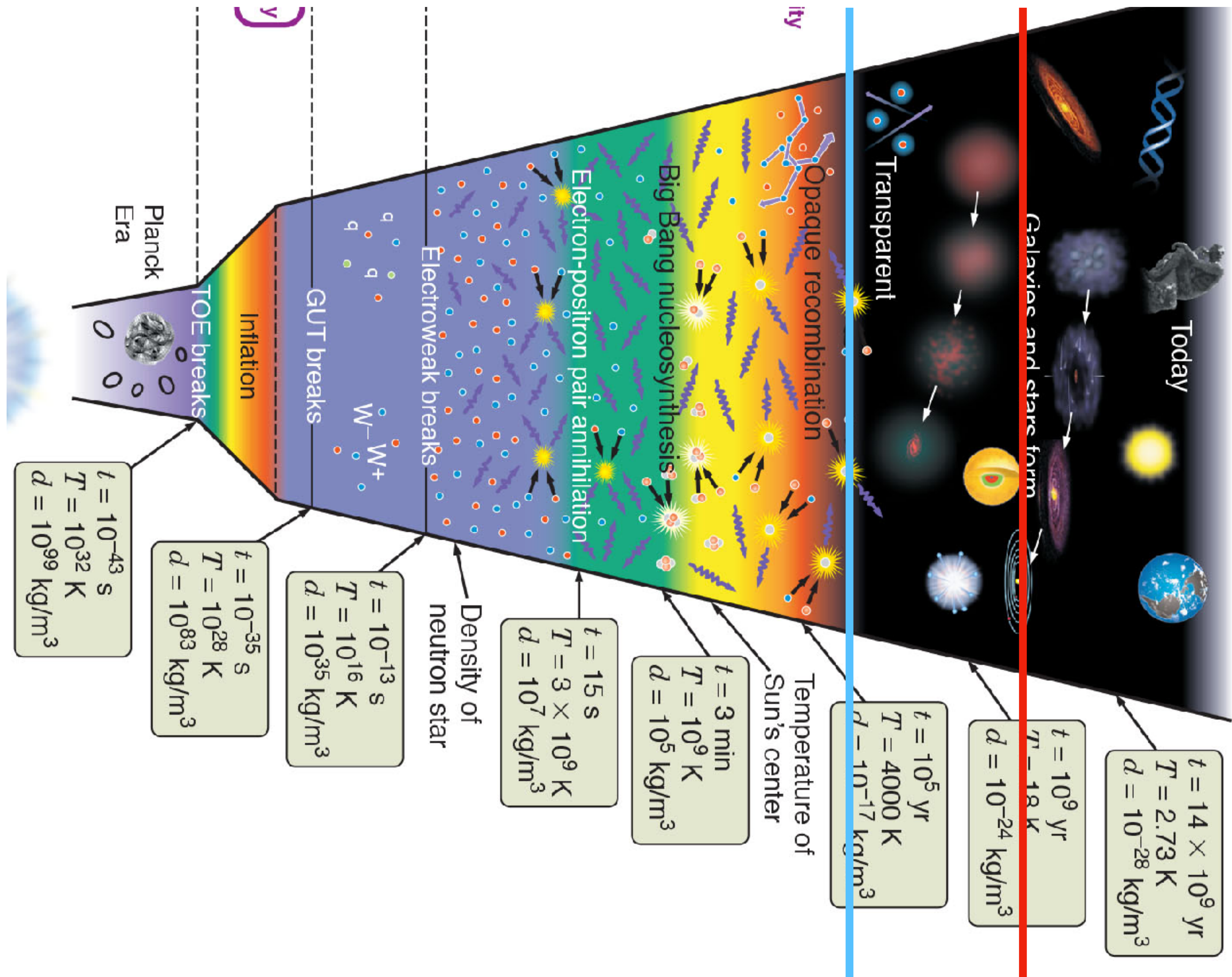
non-relativistic at time of
collapse (like WIMPs): cold

fast motions wipe out initial
overdensities on small
scales: “free-streaming”

Universe is
opaque

Electrons &
ions combine
(recombination)

Universe gets
ionized again
(reionization)



If the universe got reionized, why can we see through it?

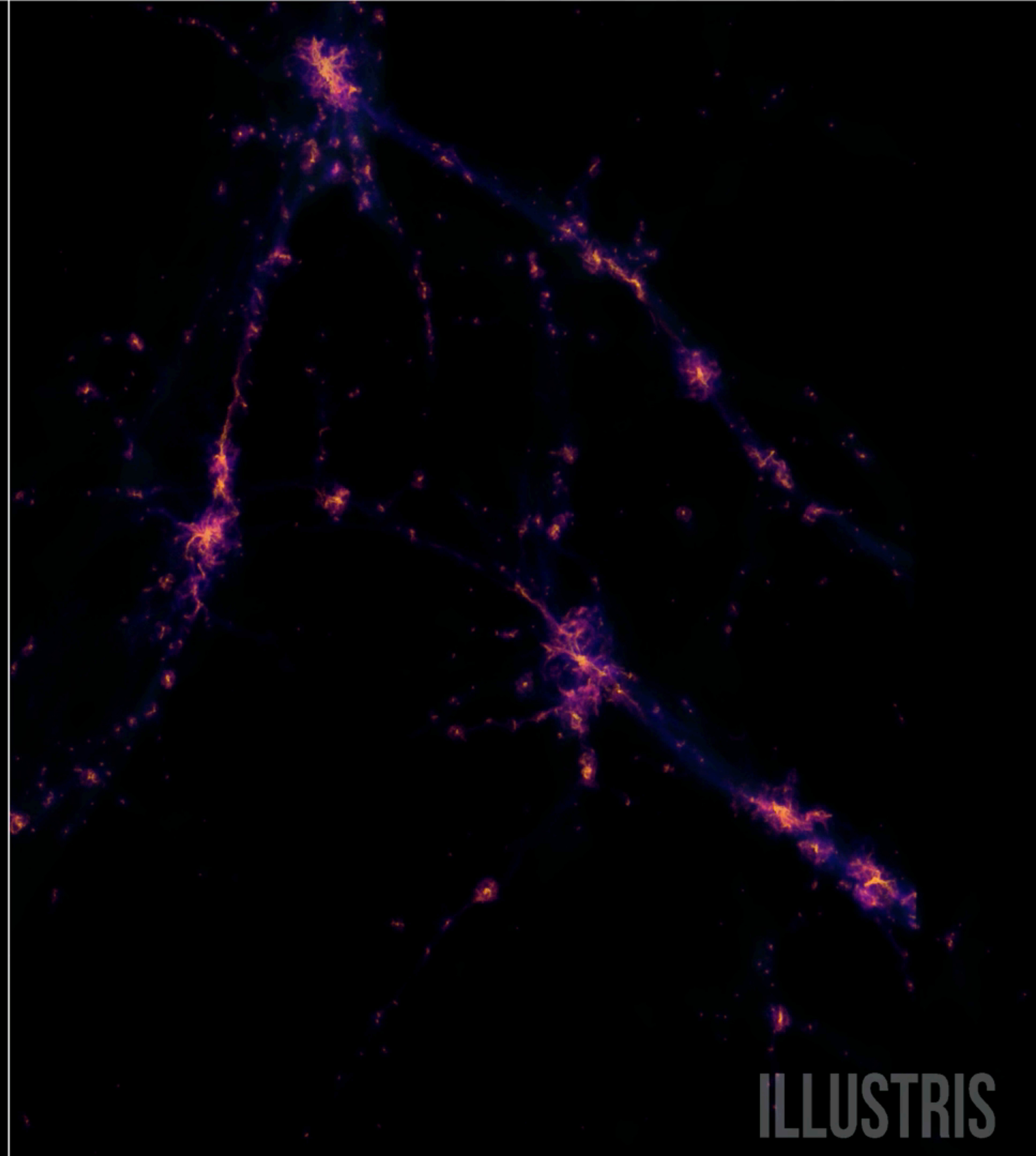
- A) Density of electrons is too low to scatter light
- B) Electrons and ions recombined again after the rate of star formation declined
- C) Electrons and ions recombined again after the rate of supermassive black hole accretion declined

$z=4.00$

$\log_{10}(M_*)=10.4$

SFR=80.0

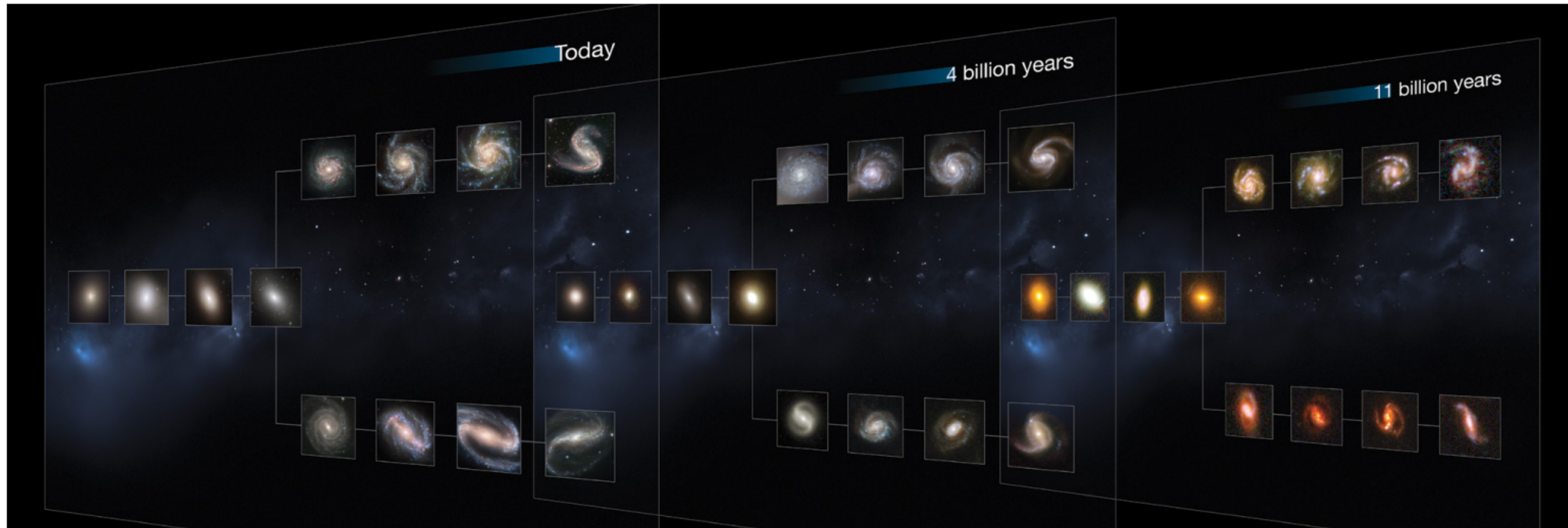
sSFR=3.07Gyr⁻¹



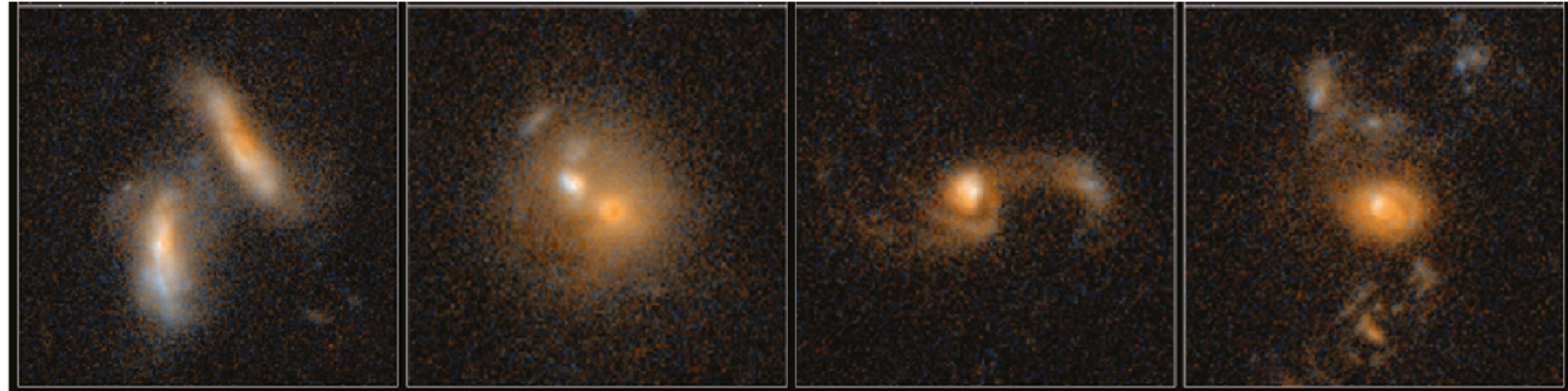
ILLUSTRIS

Spirals take time to form

galaxies get smaller the farther back in time you look



Galaxies grow in size through mergers

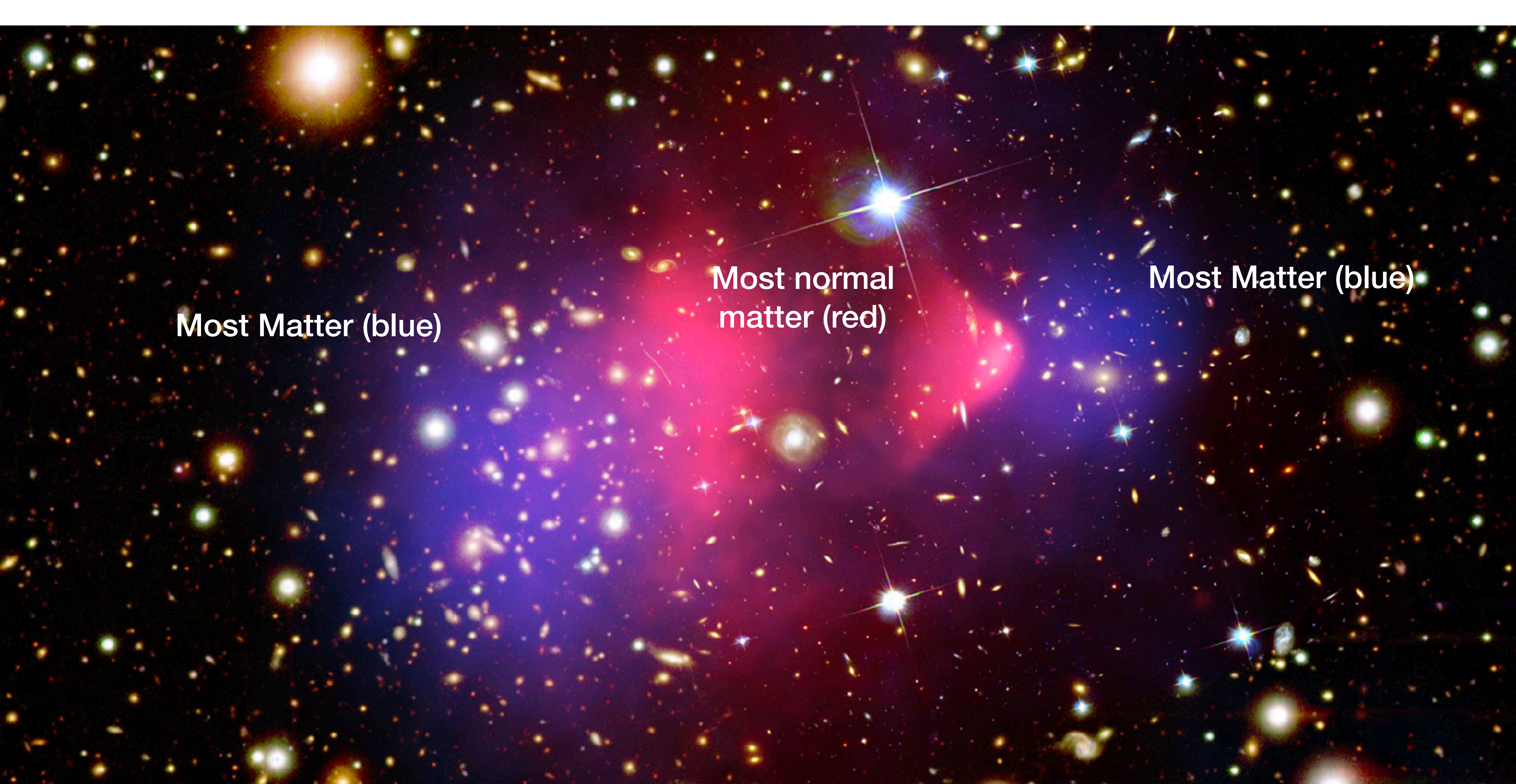


Distant



More Nearby



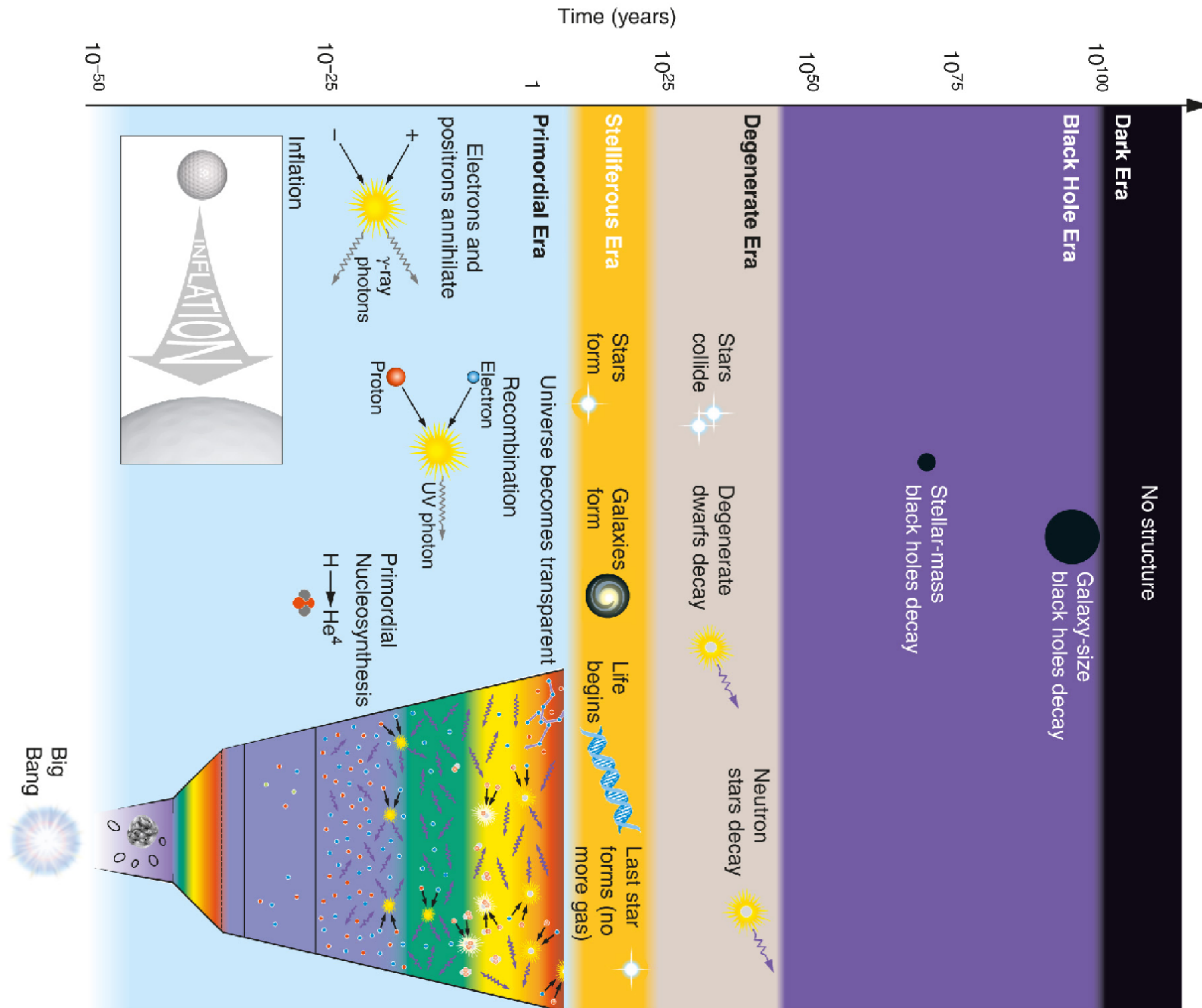


Most Matter (blue)

Most normal matter (red)

Most Matter (blue)

The Deep Future (maybe?)



Primordial Era	10^5 yr
Stelliferous Era	10^{14} yr
Degenerate Era	10^{39} yr
Black Hole Era	10^{100} yr
Dark Era	infinity?