

## Homework 9

Due **November 5 by 10:45am via Canvas upload**

Please show all work, writing solutions/explanations clearly, or no credit will be given. You are encouraged to work together, but everyone must turn in independent solutions; do not copy from others or from any other sources.

1.
  - (a) If the Sun collapsed down to the size of a white dwarf ( $R_{\text{WD}} \approx 6000$  km) without losing any angular momentum, what would its rotation period be? (Assume the Sun rotates as rigid body, so that its moment of inertia  $I = \frac{2}{5}MR^2$ .)
  - (b) How fast can this white dwarf-sized Sun rotate before it breaks up? (Break-up speed is reached when the centrifugal acceleration at its equator equals the surface gravity  $g$ . Assume the star's shape remains spherical, although this will definitely not be the case in reality.)
  - (c) Based on your answer, could a white dwarf be the object behind what we observe as pulsars? Why or why not?
  
2. Consider a neutron star with a mass of  $1.5 M_{\odot}$ .
  - (a) What is its radius, expressed as a fraction of its Schwarzschild radius?
  - (b) What is its mean density? How does it compare to the mean density of a carbon nucleus, which has a radius  $r \approx 3 \times 10^{-15}$  m.
  - (c) If this neutron star collapsed to form a black hole, how close could you get to it before being ripped apart? Assume the human body can withstand  $1.4 \times 10^5$  N of extending force before tearing apart. Explain the difference between this “ripping radius” and the event horizon of a black hole.
  
3. Following the timeline given on the class page describing your Communicating Science Projects, submit a paragraph (at least 4–5 sentences) describing the final project: what presentation method was chosen, why, how it will be executed, and a timeline of steps towards completion. You may receive feedback from me and/or a request for more information on this step as well.