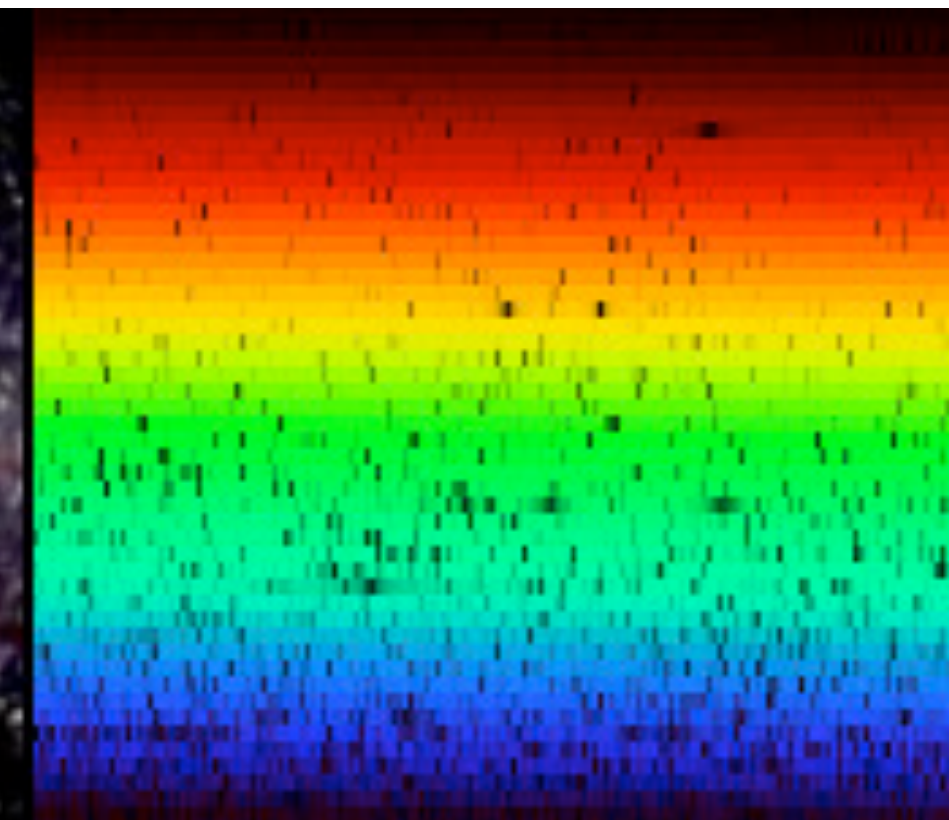




ASTR/PHYS 3070: Foundations Astronomy



Week 15 Tuesday

Today's Agenda

- Milky Way structure
- Evidence for dark matter
- Galaxy types
- Extragalactic distances

Announcements / Reminders

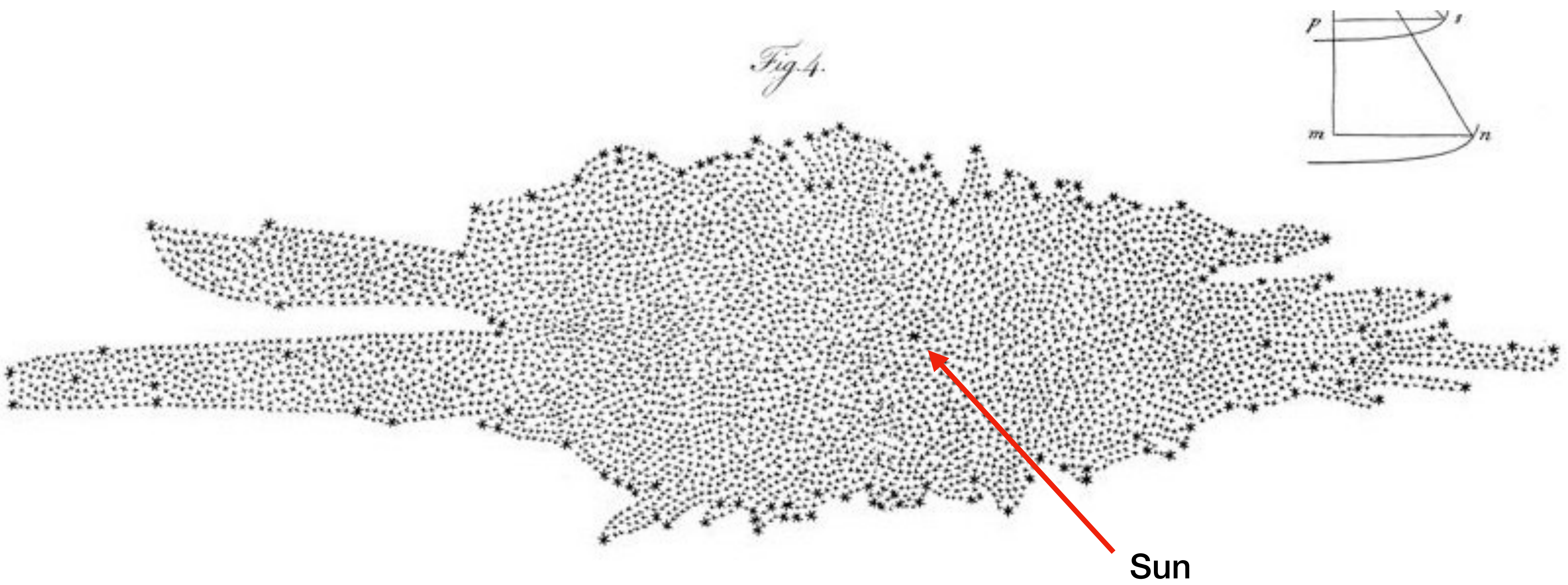
- HW 10 due this Friday @11:59pm
- HW 11 also available, due next week
- Read Chapters 20 & 23.0-3 (& 24)

- HEAP TODAY @ 3pm AND tomorrow @ 8am
 - Neutrino detectors & neutrinos from blazars
- Colloquium Friday @ 2pm
 - Intercellular Communication & Drug Delivery

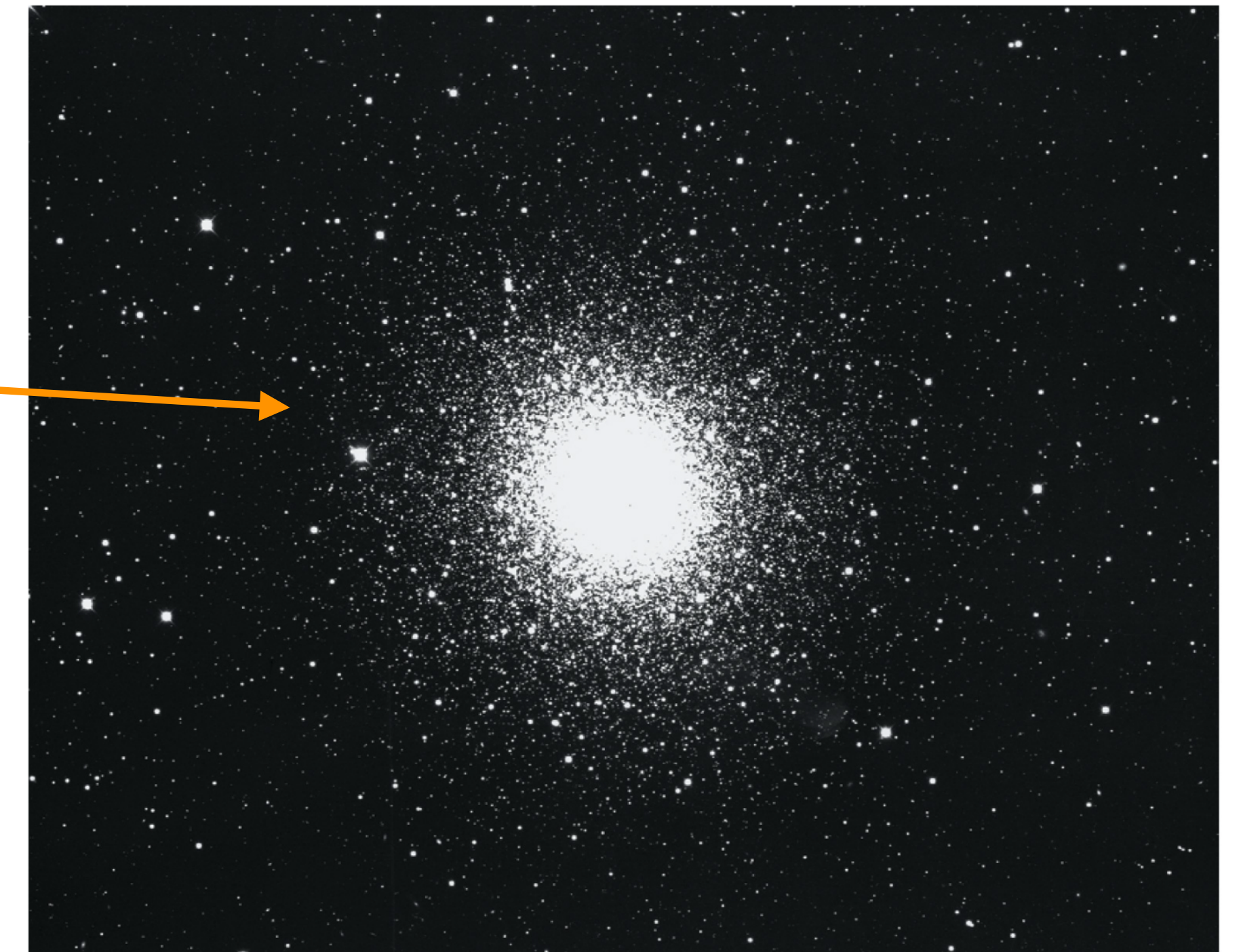
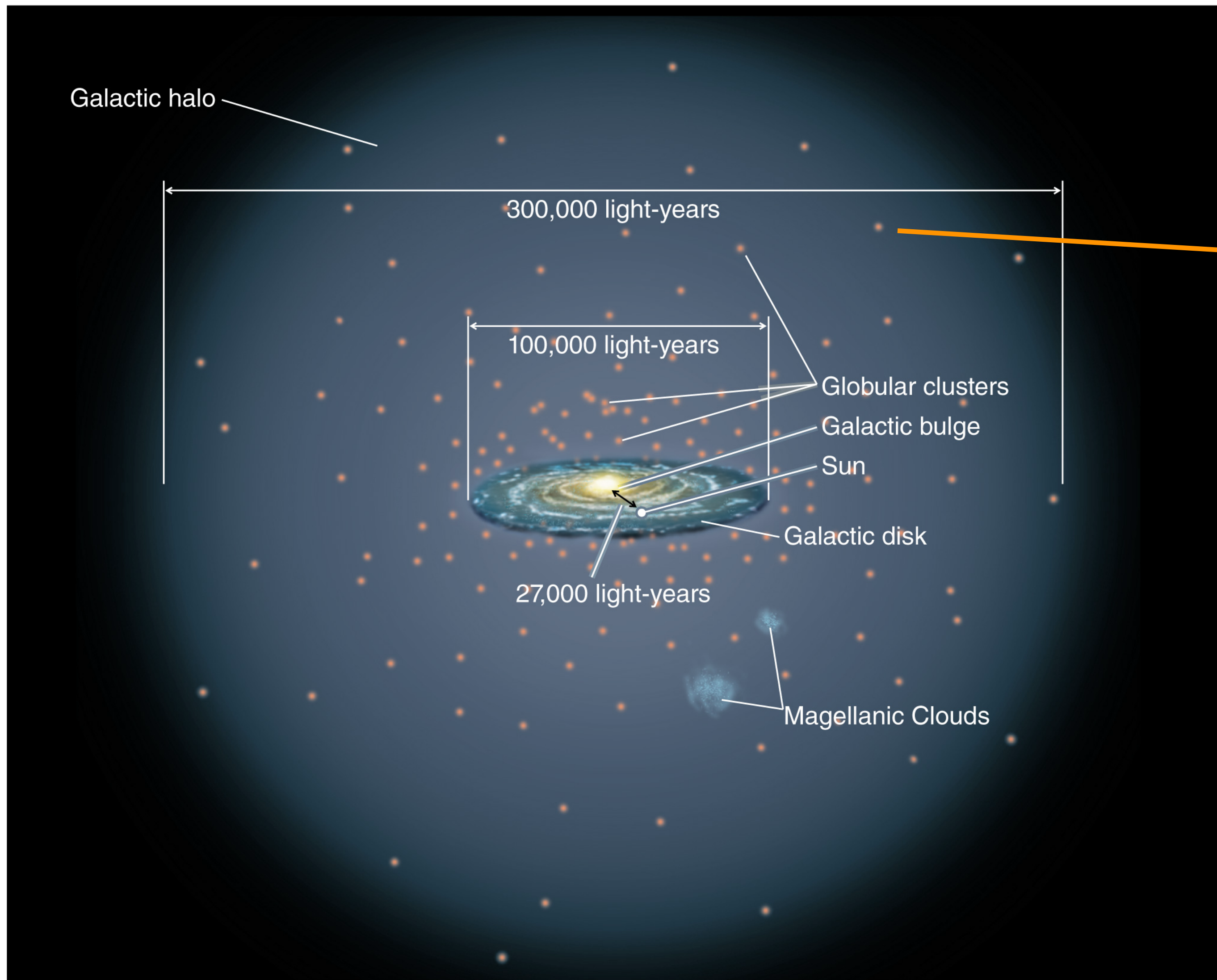
Our Galaxy, the Milky Way



Star counts: William and Caroline Herschel (1785)



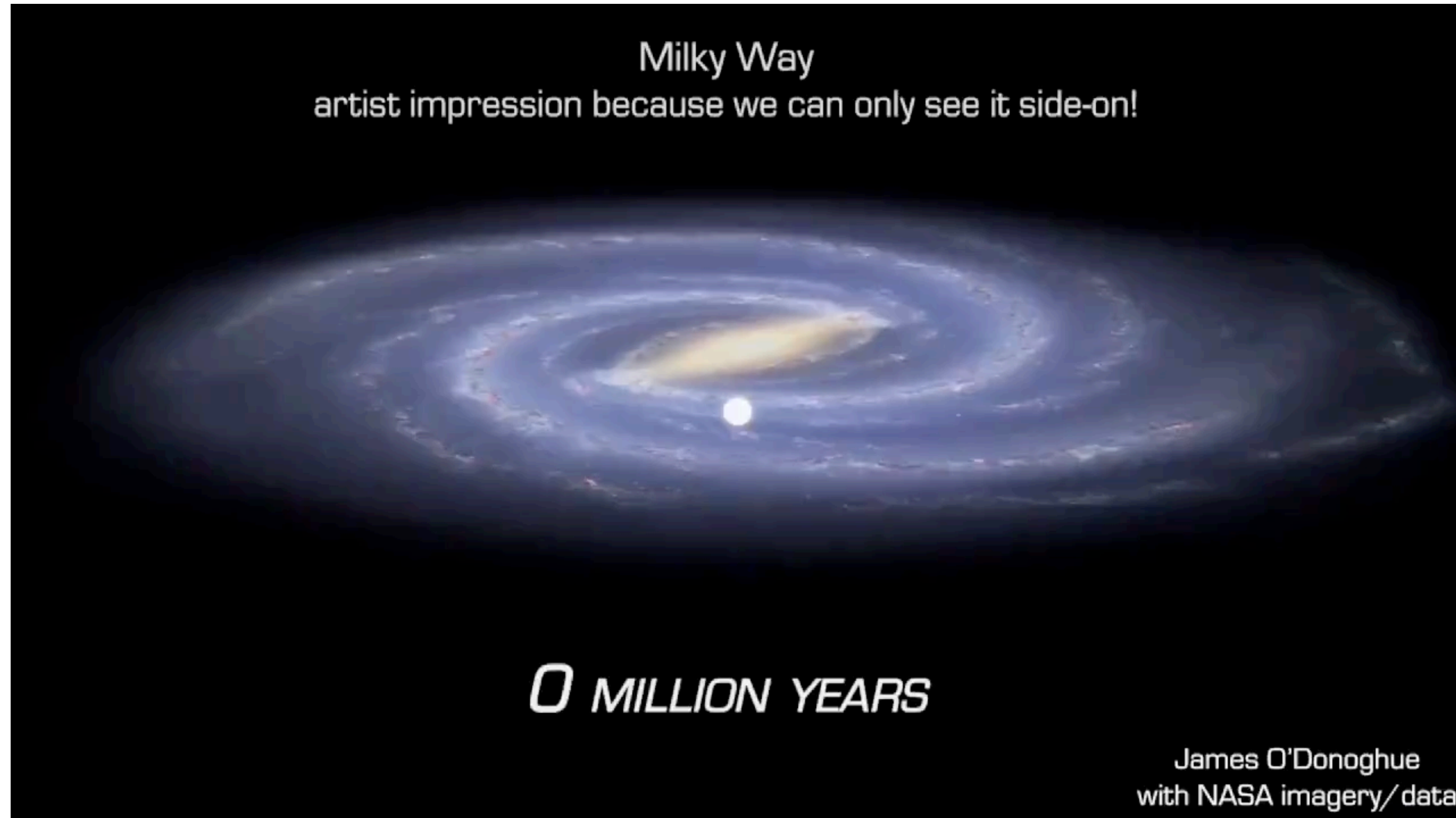
Globular clusters revealed the scale of the MW



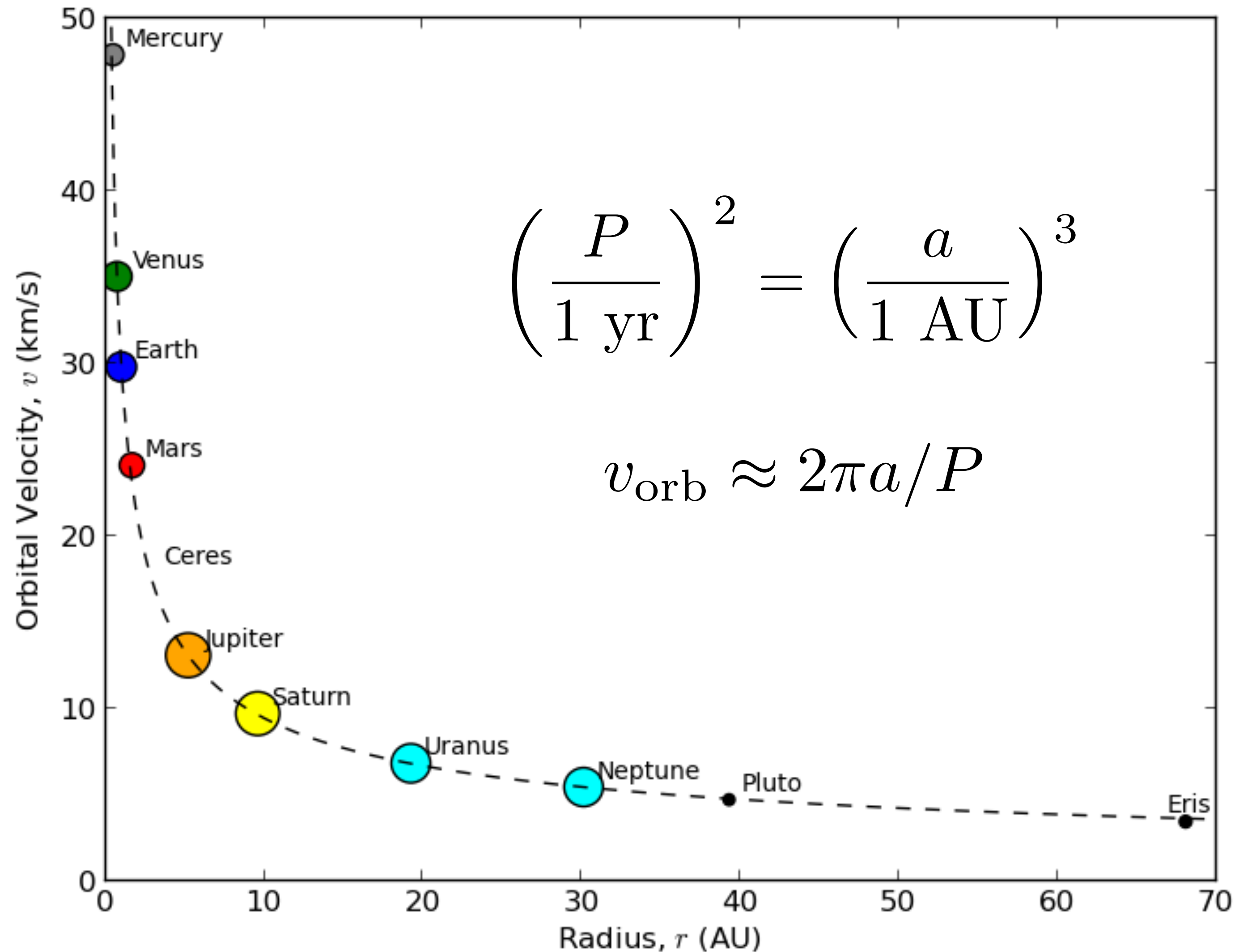
G X U V I R

Variable stars like Cepheids (called RR Lyrae stars) were used to estimate the distance to globular clusters, which were assumed to be distributed uniformly around the center of the MW

Sun's Motion through the Galaxy



How do stars move in the Galaxy?



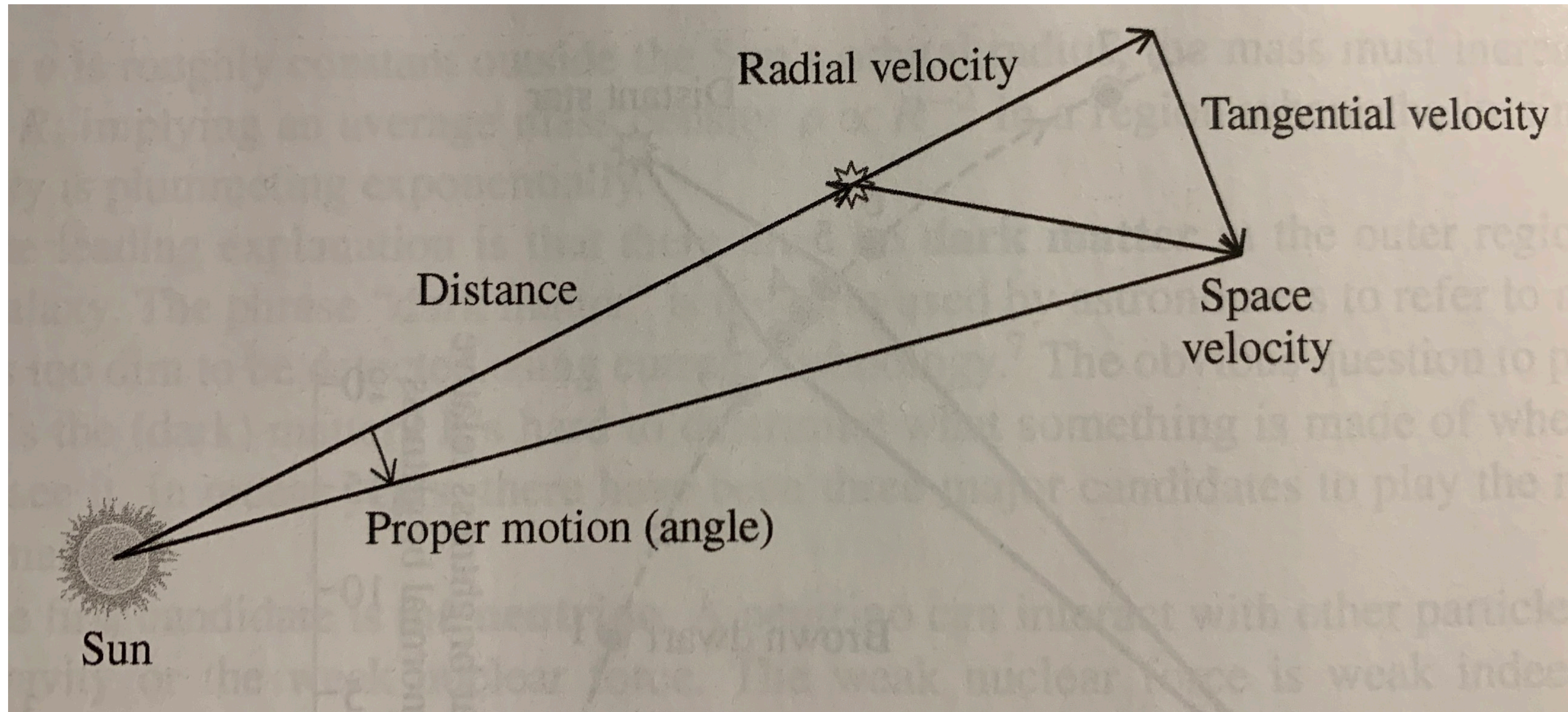
Kepler's 3rd Law in the Galaxy

$$M_{\odot} + M_G(< r) = \frac{(a/1 \text{ AU})^3}{(P/1 \text{ yr})^2}$$

**Mass
inside
Sun's orbit**

How do we get 3D star velocities?

Radial Velocity: $v_r = \frac{\Delta\lambda}{\lambda} c$

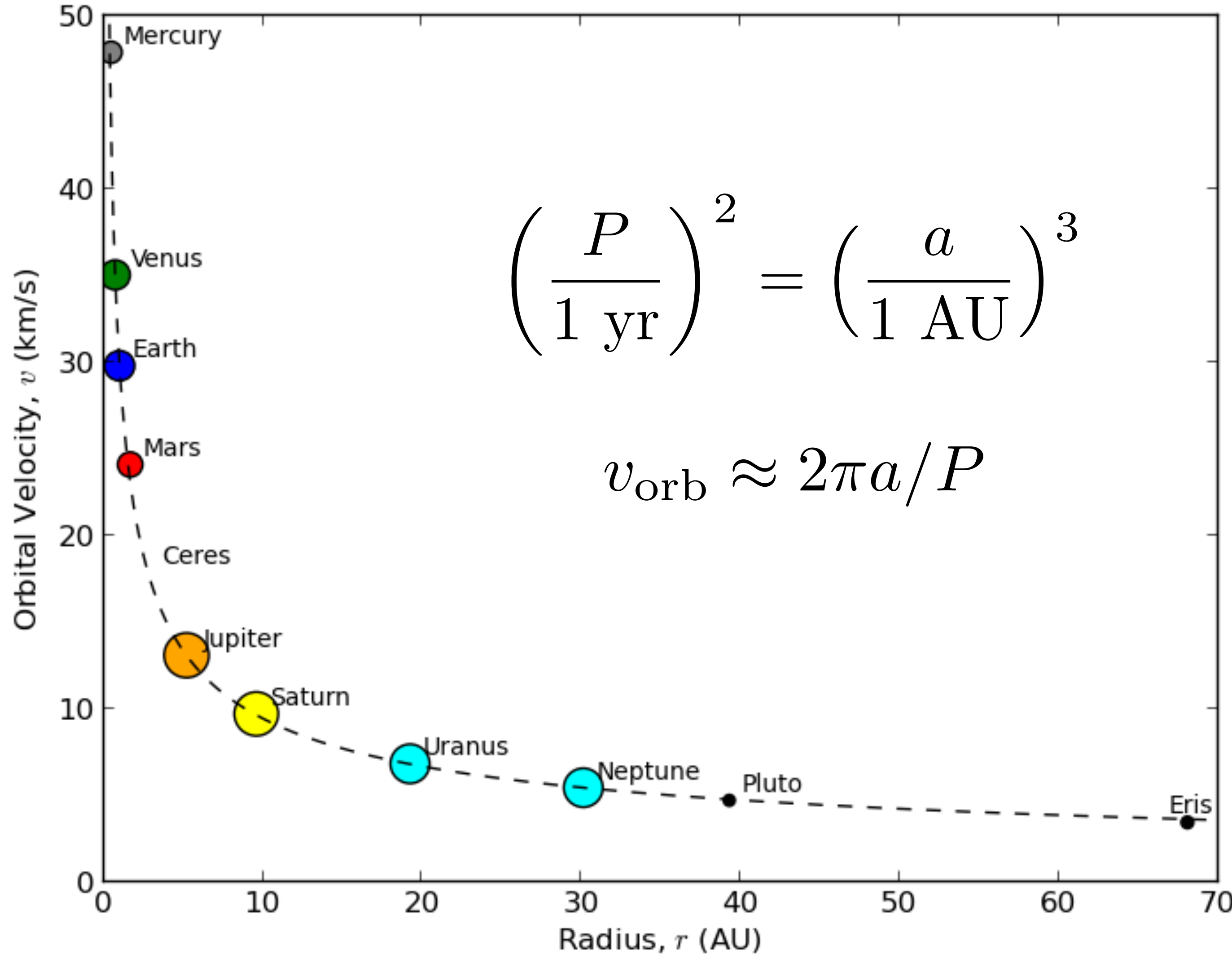


Proper Motion:

$$\mu = \frac{v_t}{d}$$

$$v = \sqrt{v_r^2 + v_t^2}$$

How do stars move in the Galaxy?



Kepler's 3rd Law in the Galaxy

$$M_{\odot} + M_G(< r) = \frac{(a/1 \text{ AU})^3}{(P/1 \text{ yr})^2}$$

**Mass
inside
Sun's orbit**

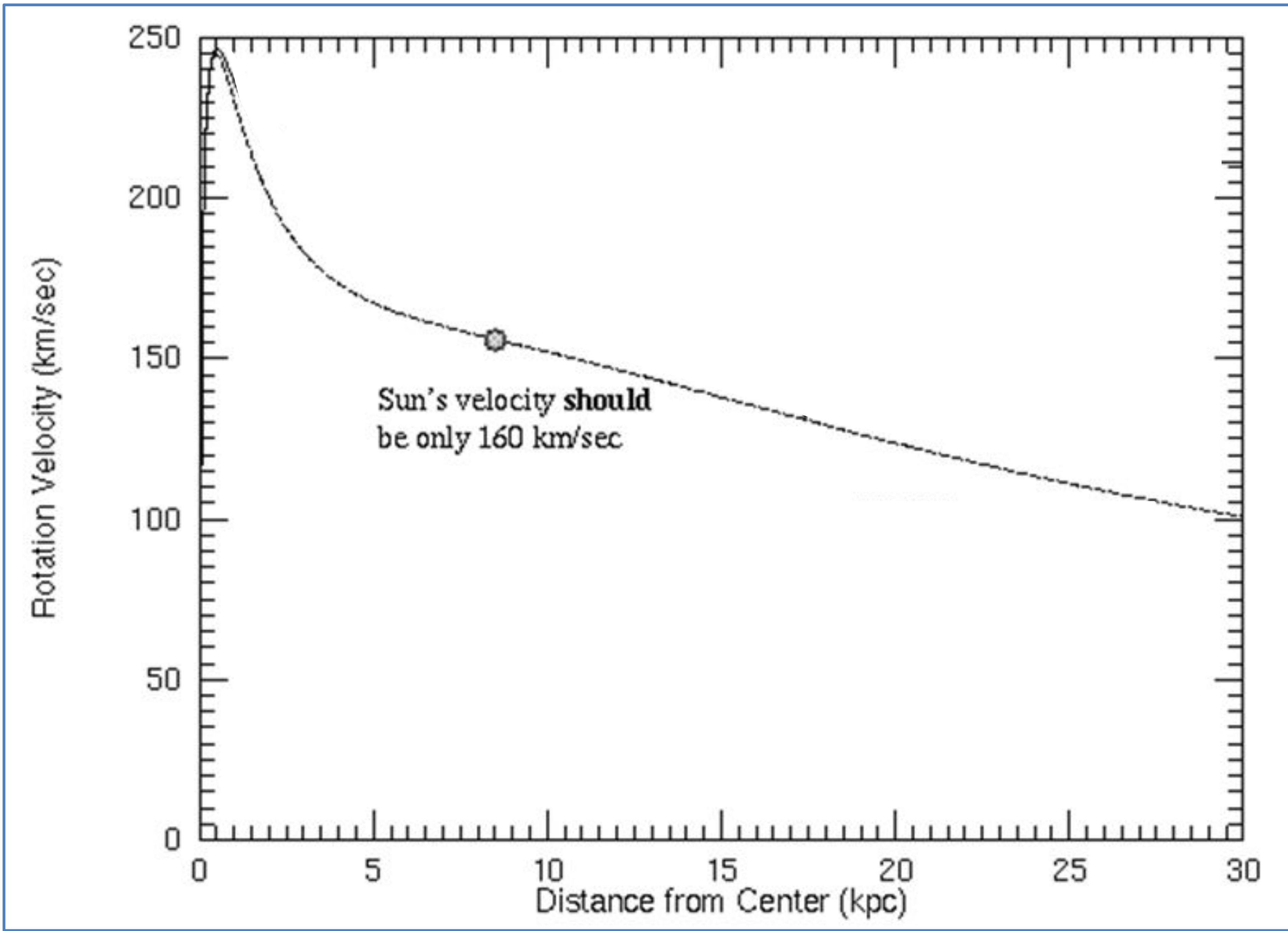
$$a \approx 8 \text{ kpc}$$

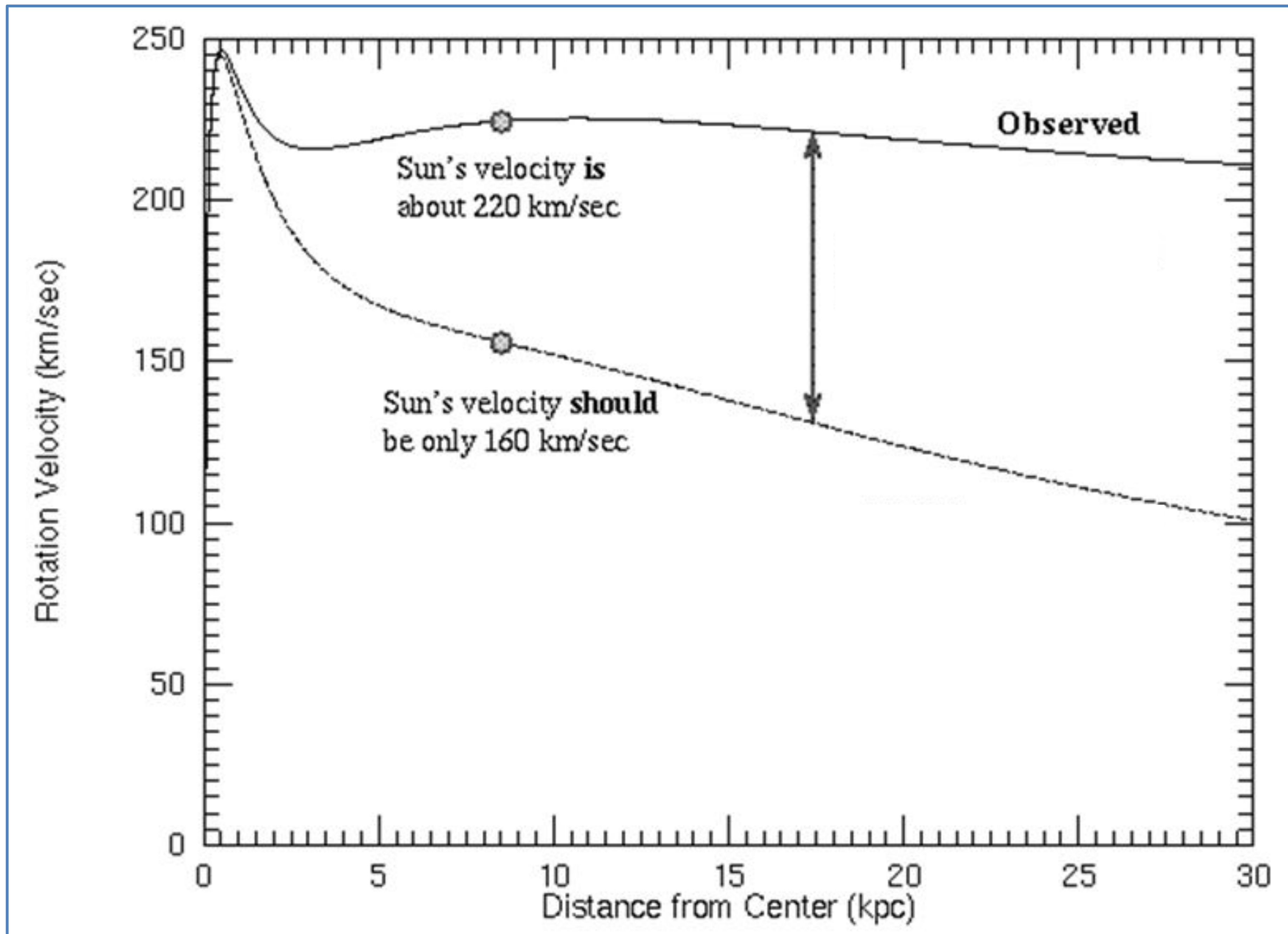
$$P \approx 220 \text{ Myr}$$

$$M_G(< r) \approx 9.3 \times 10^{10} M_{\odot}$$

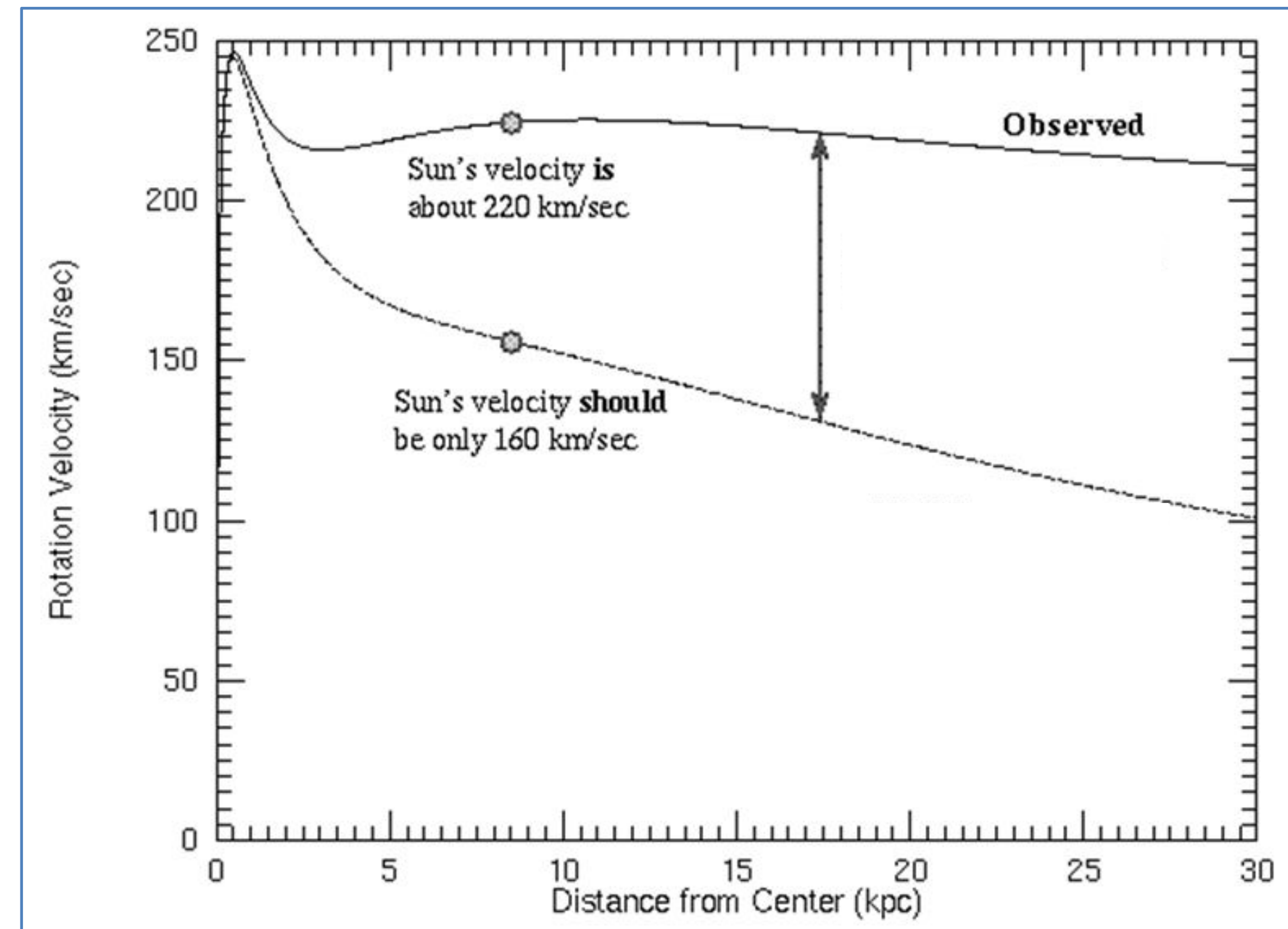
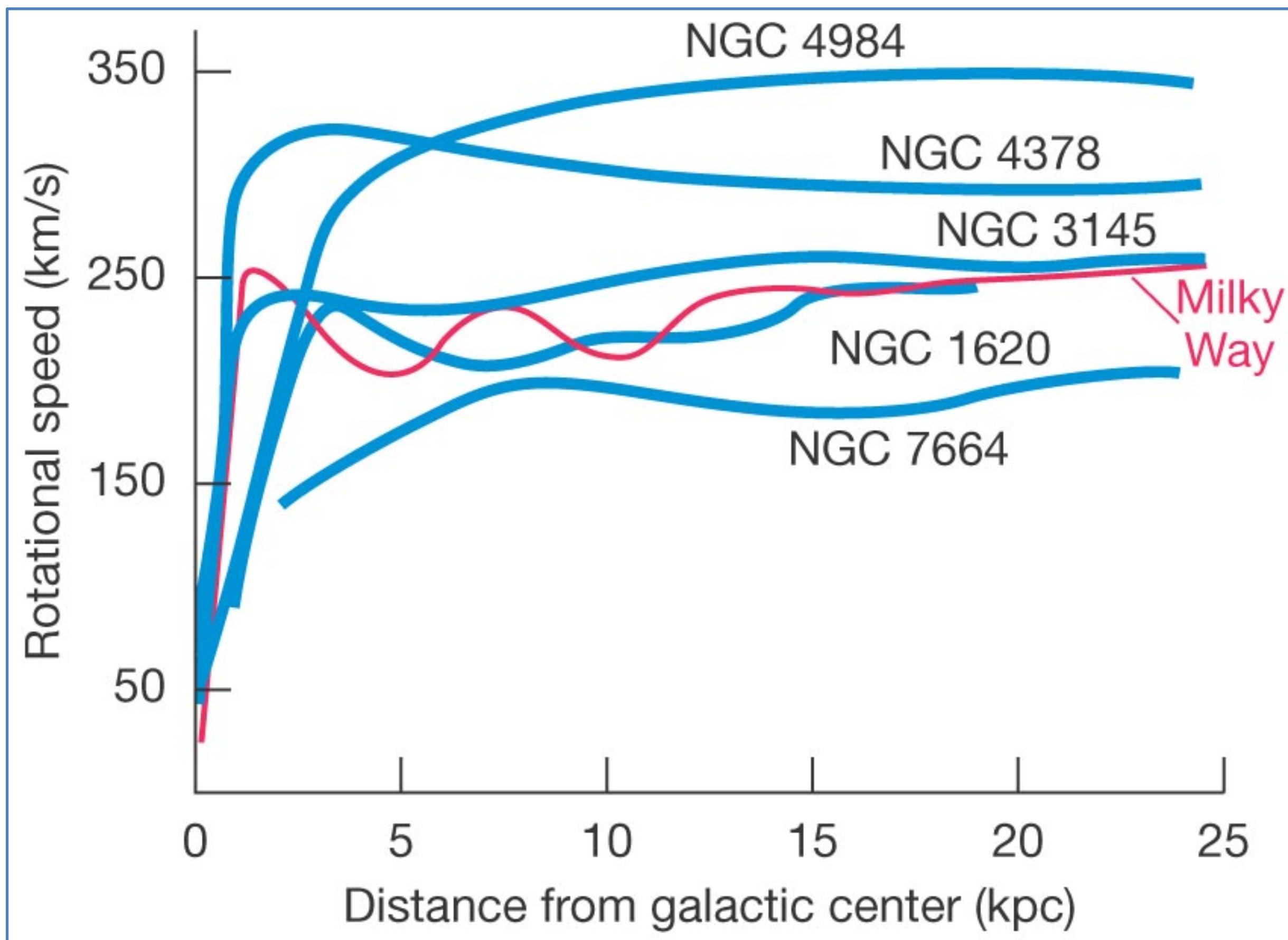
$$\frac{v(R)^2}{R} = \frac{GM_G(< R)}{R^2}$$

$$v(r) \propto \left(\frac{M_G(< R)}{R}\right)^{1/2} \propto R^{-1/2}$$

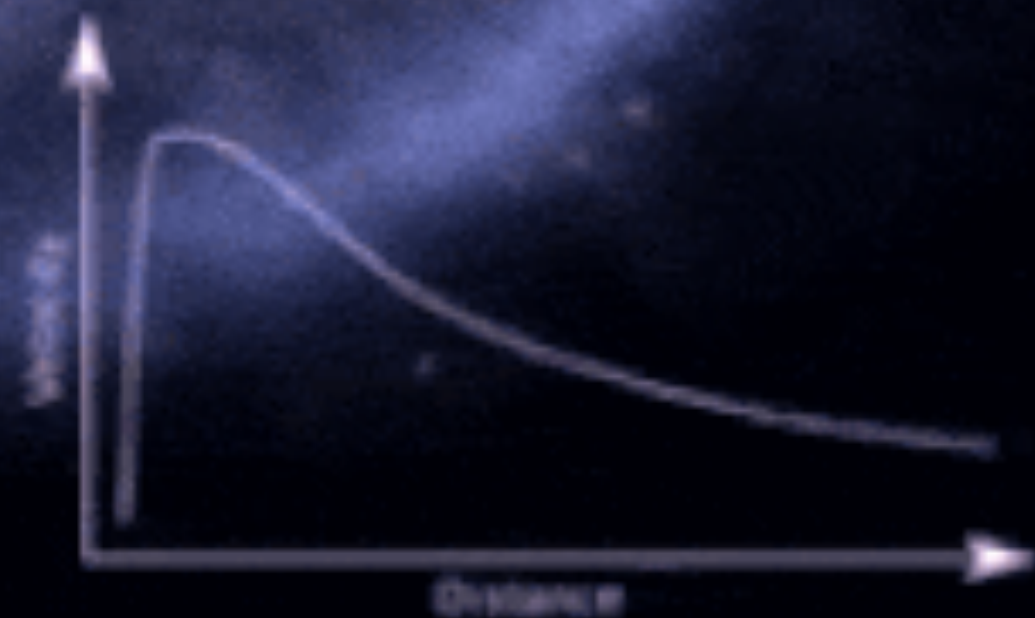




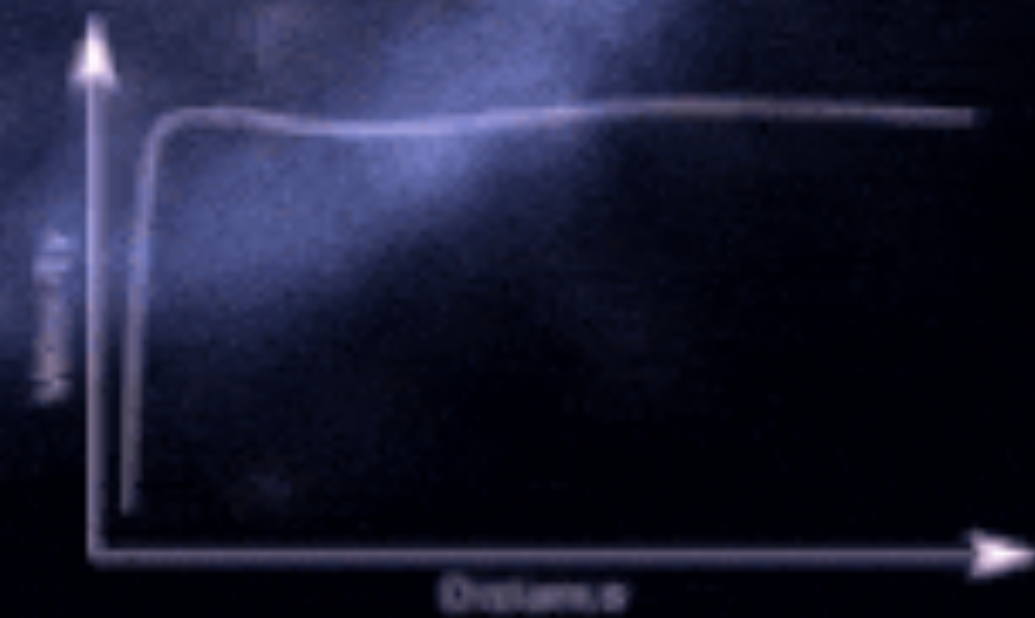
Milky Way is not alone – there is extra, non-luminous matter in galaxies: “dark matter”



Expected rotation

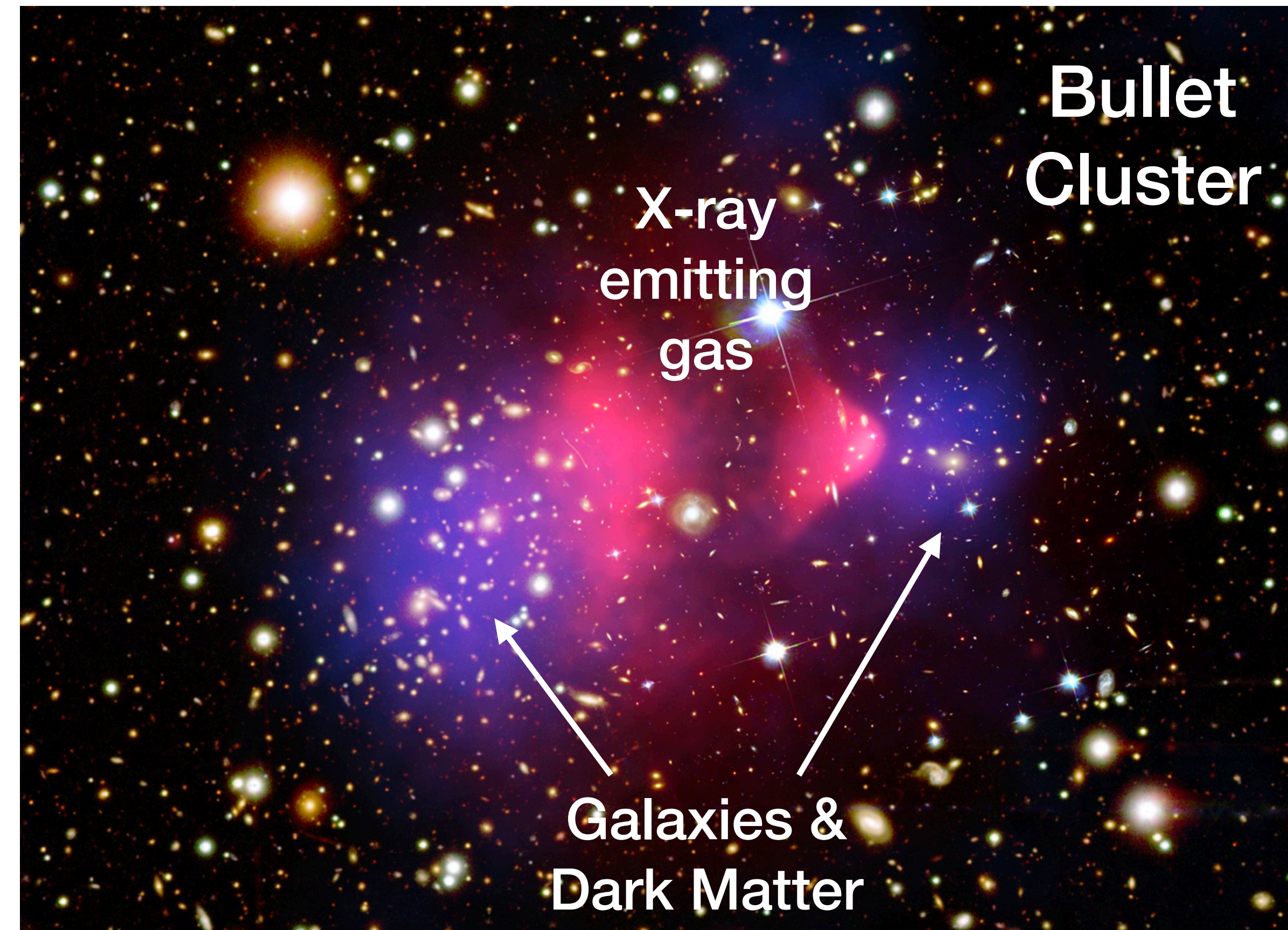


Observed rotation

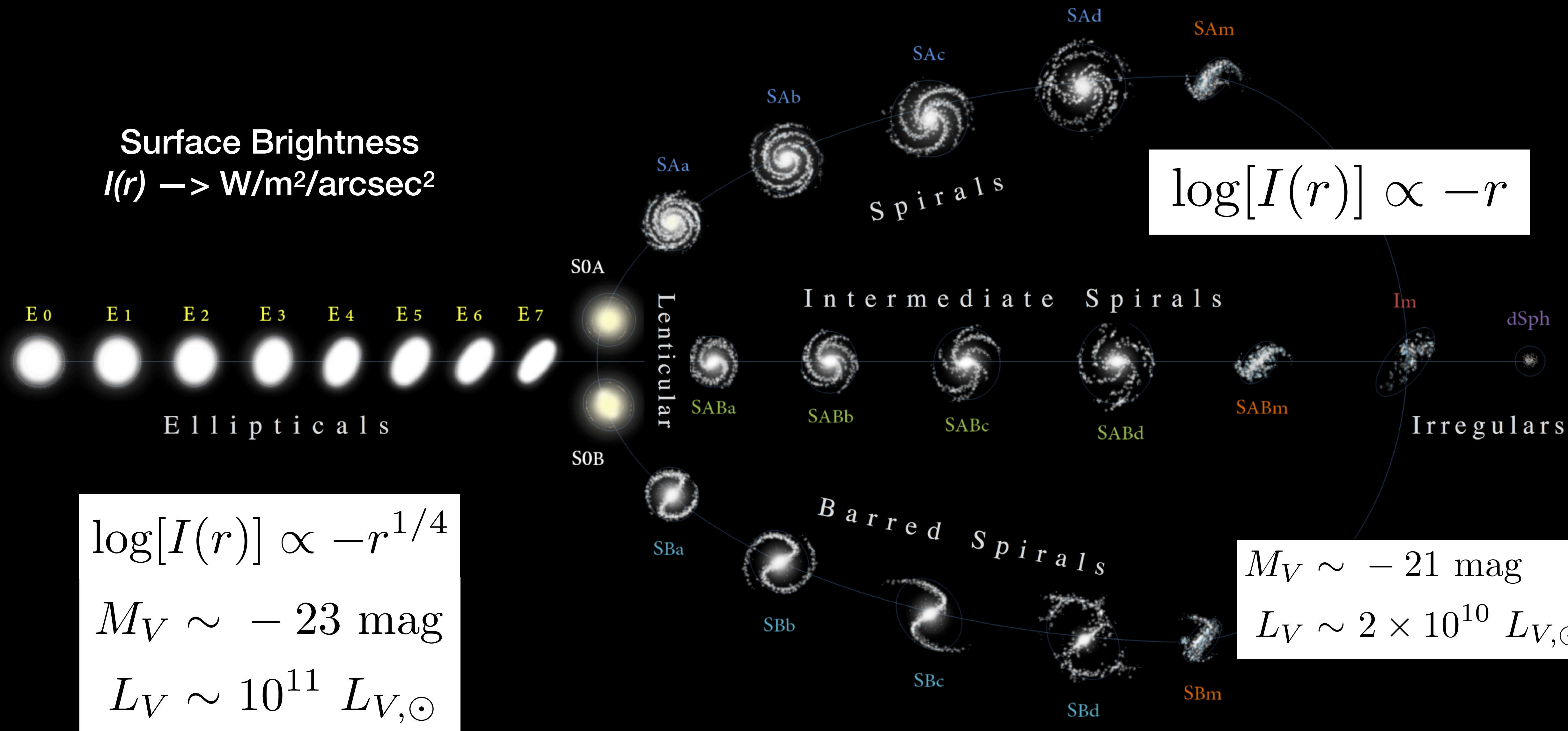


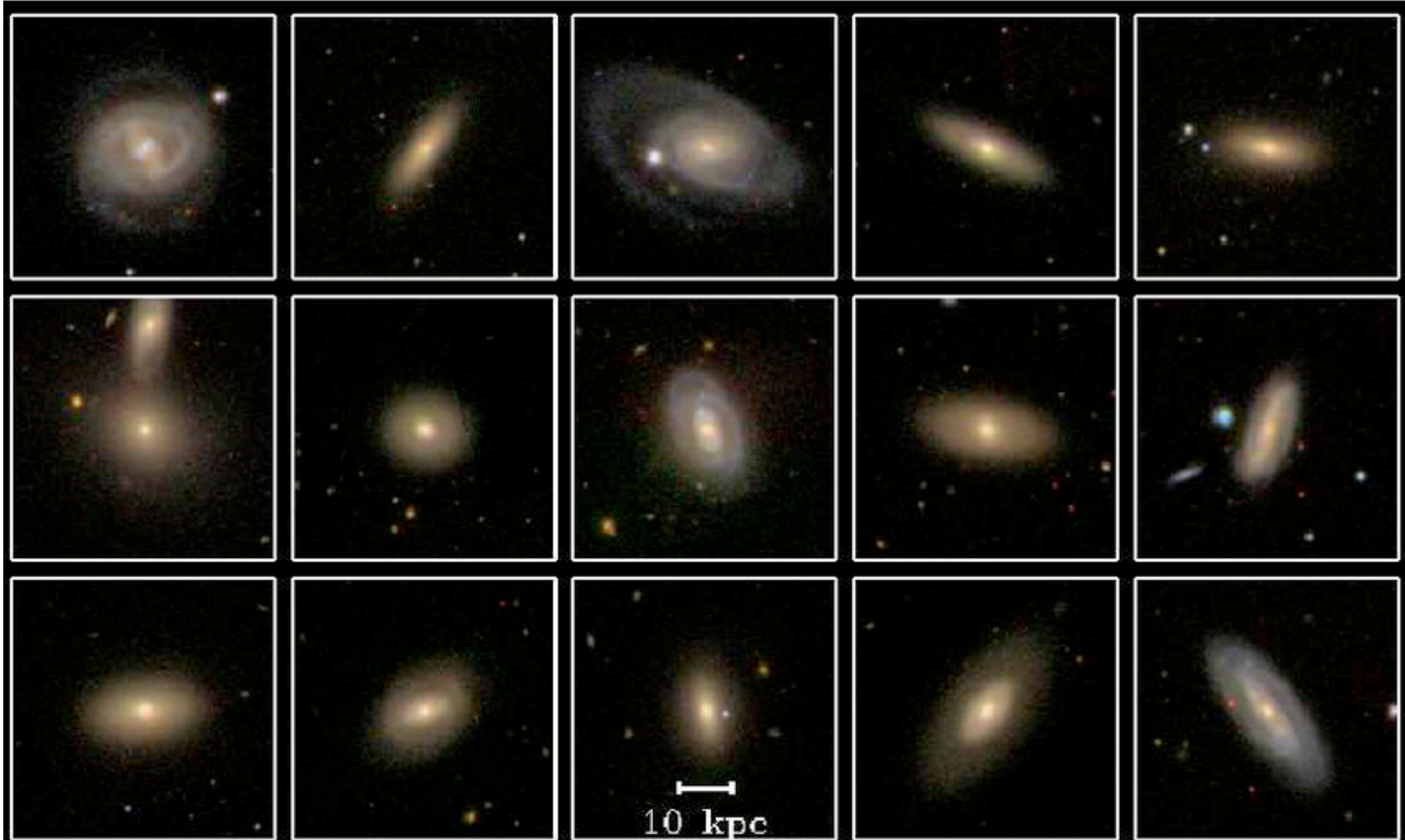
Dark Matter: what is it?

- Neutrinos (like those produced in fusion)
 - Have mass, but not enough
 - New kind? Sterile Neutrino
- WIMP (Weakly Interacting Massive Particle)
 - Direct detection searches have failed
 - “WIMP miracle” not miraculous
- MACHO (MAssive Compact Halo Object)
 - WDs, NSs, BHs roaming around
 - Can detect via gravitational lensing - ruled out
- Theorists are clever - can invent other options!
- Modified Gravity (explains galaxy rotation, but...)

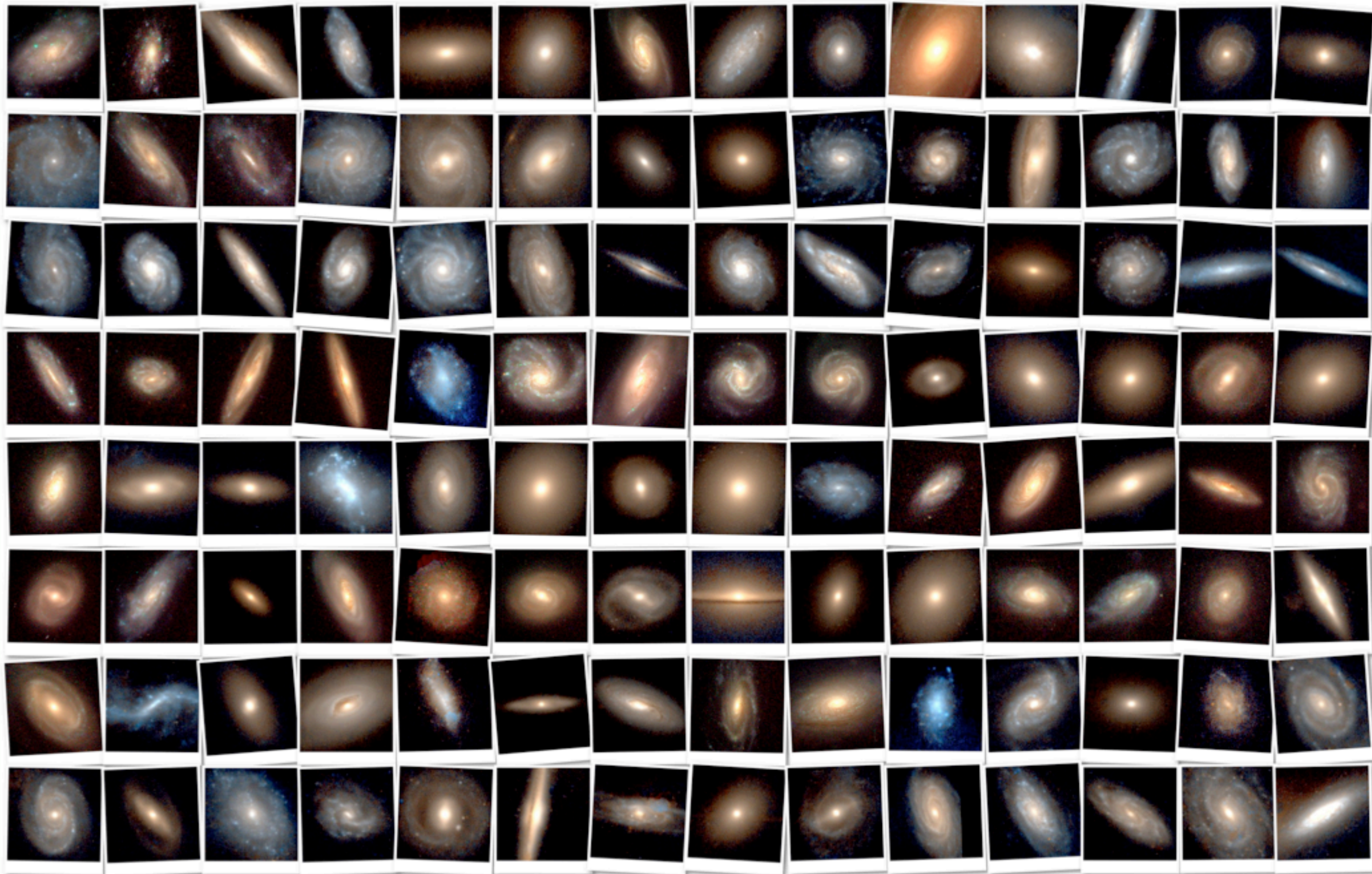


HUBBLE-DE VAUCOULEURS DIAGRAM



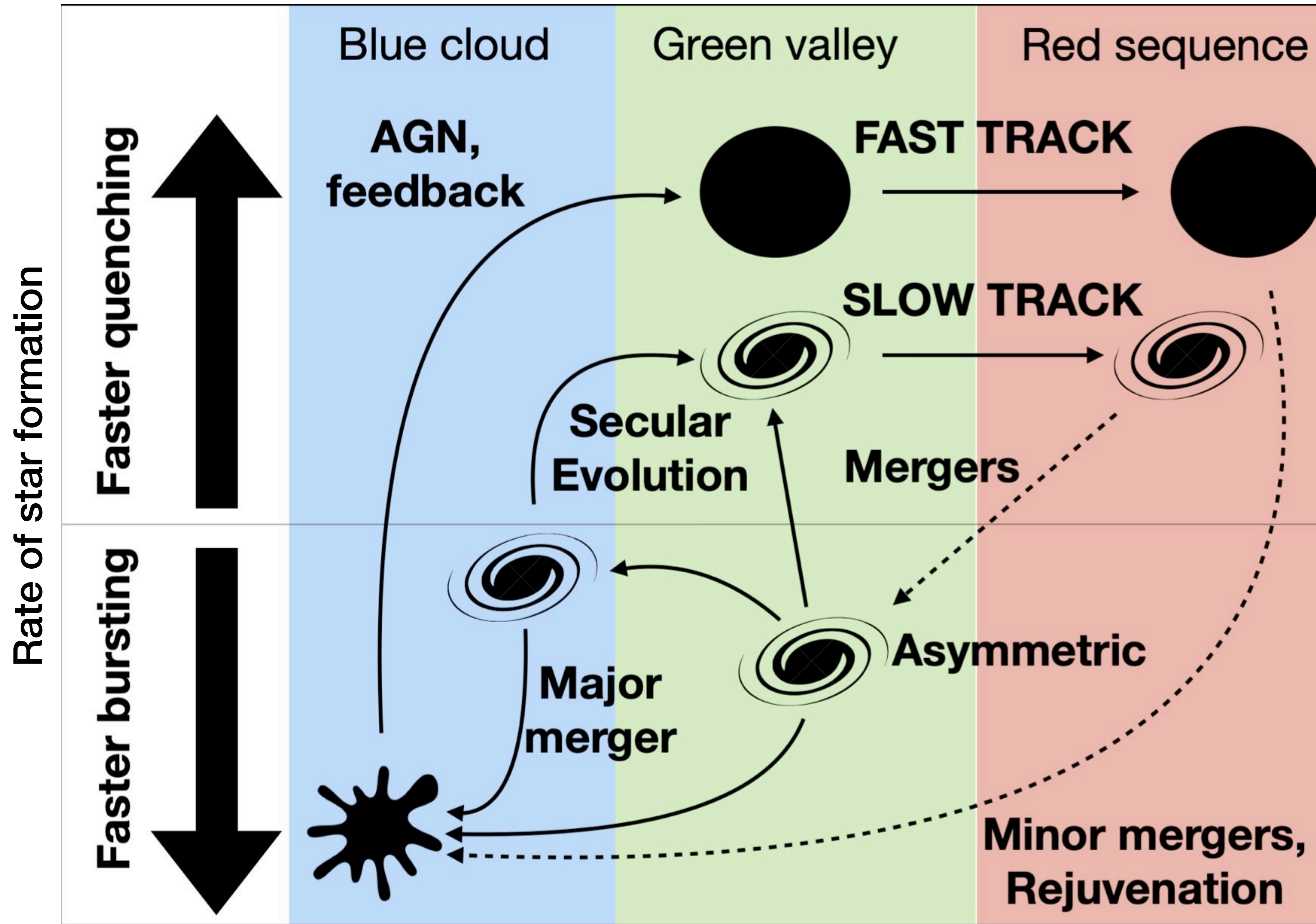


SDSS, $0.02 < z < 0.03$, $10.6 < \text{Log } M/M_{\odot} < 10.8$ Milky Way-mass



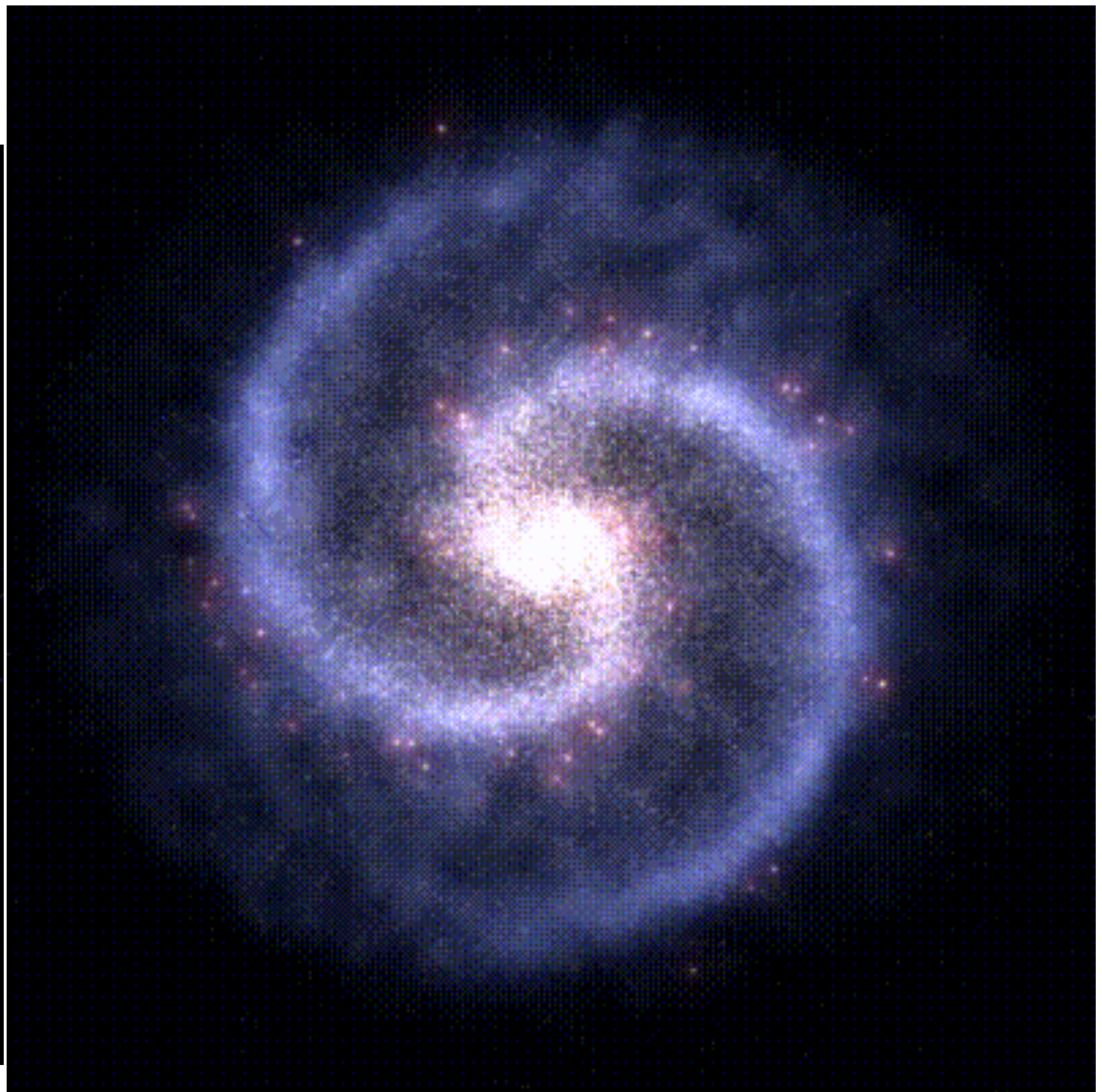
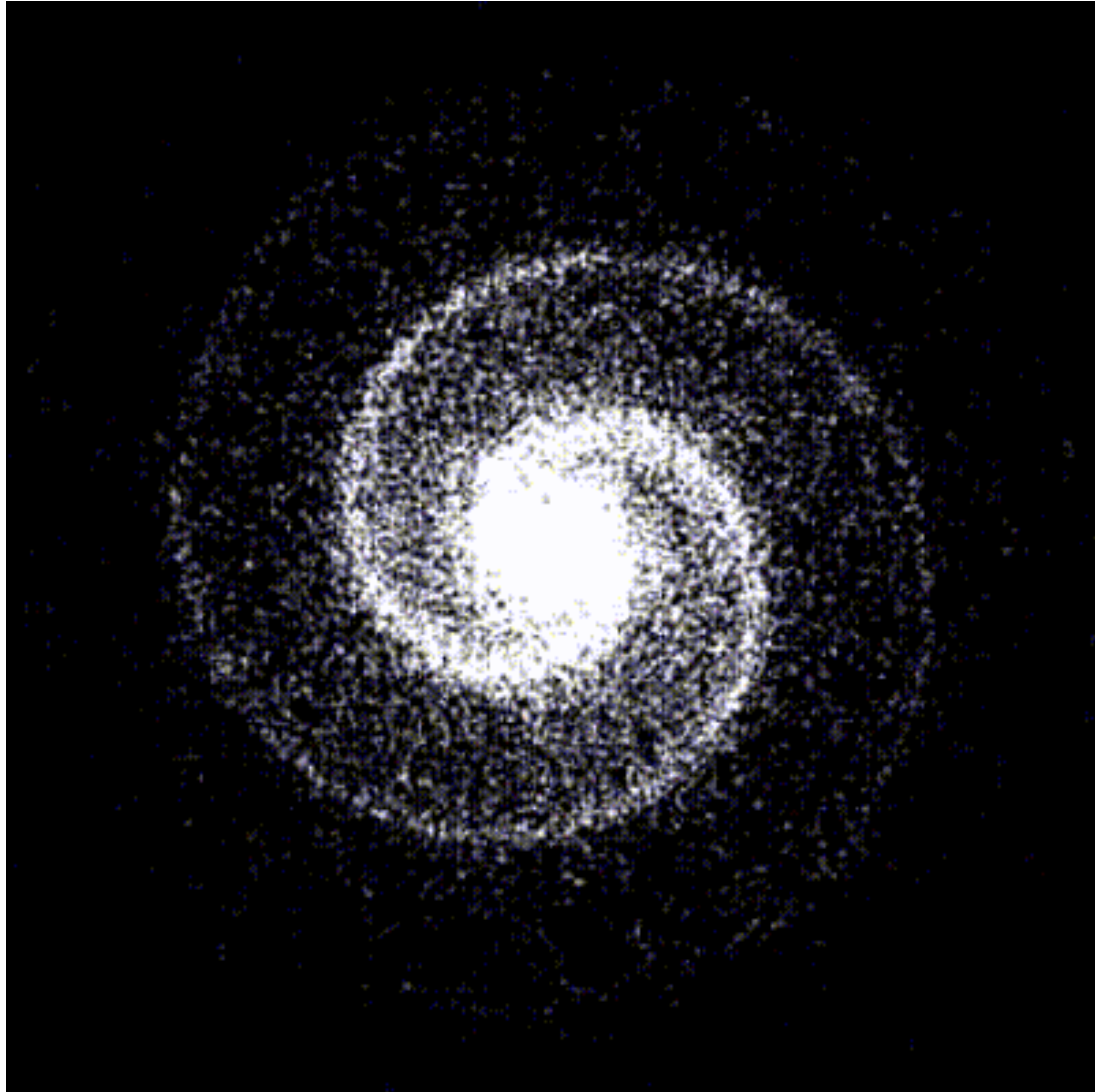
Created by Zolt Frei and James E. Gunn Copyright © 1999 Princeton University Press

Color (current amount of star formation)



Growth / Evolution of Galaxies

Spiral Arms



Galaxies are not isolated





and R frames. After sky subtraction, the coadded frames were then normalized to their respective exposure times, resulting in pixel values in ADUs/second.

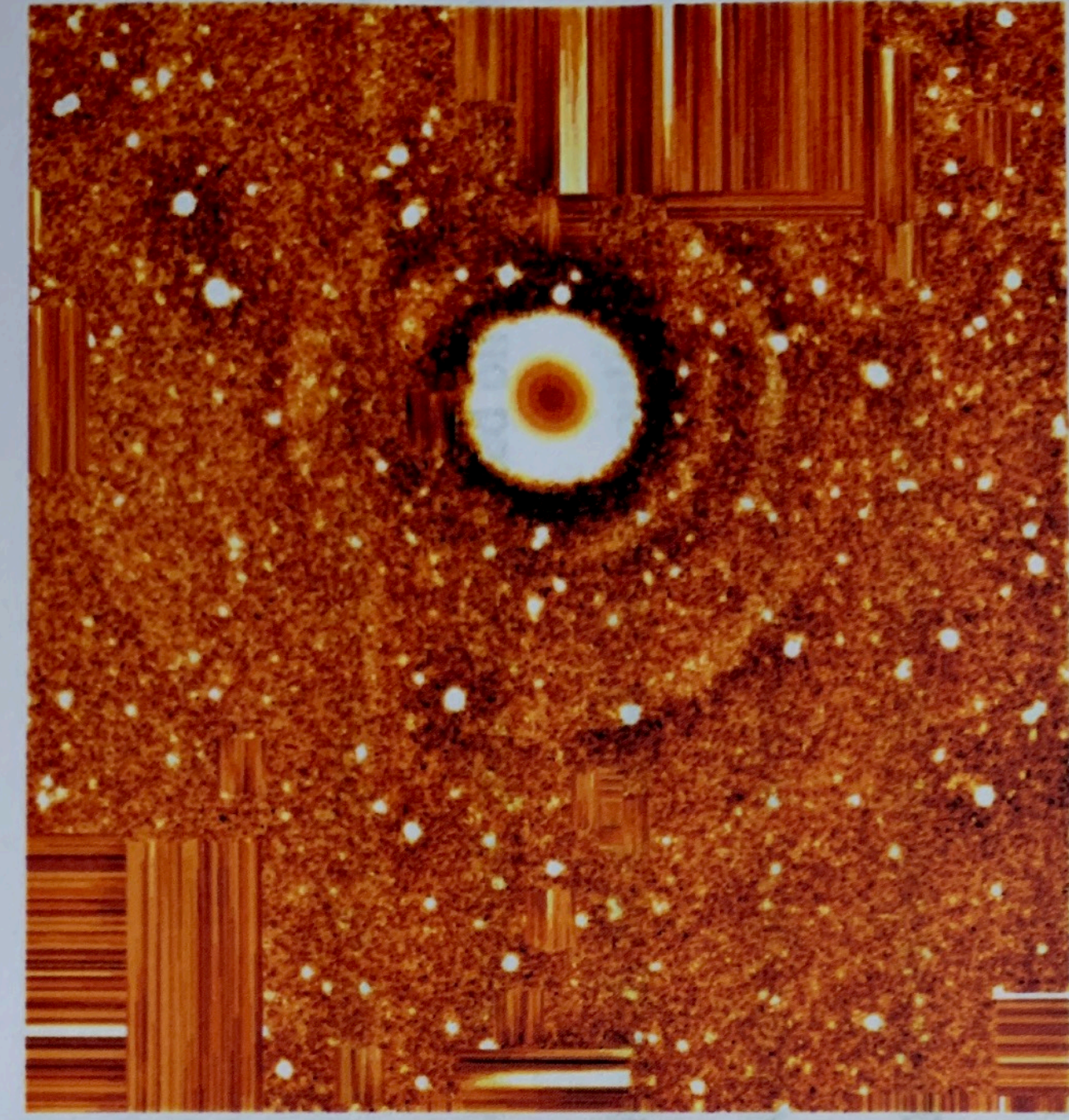
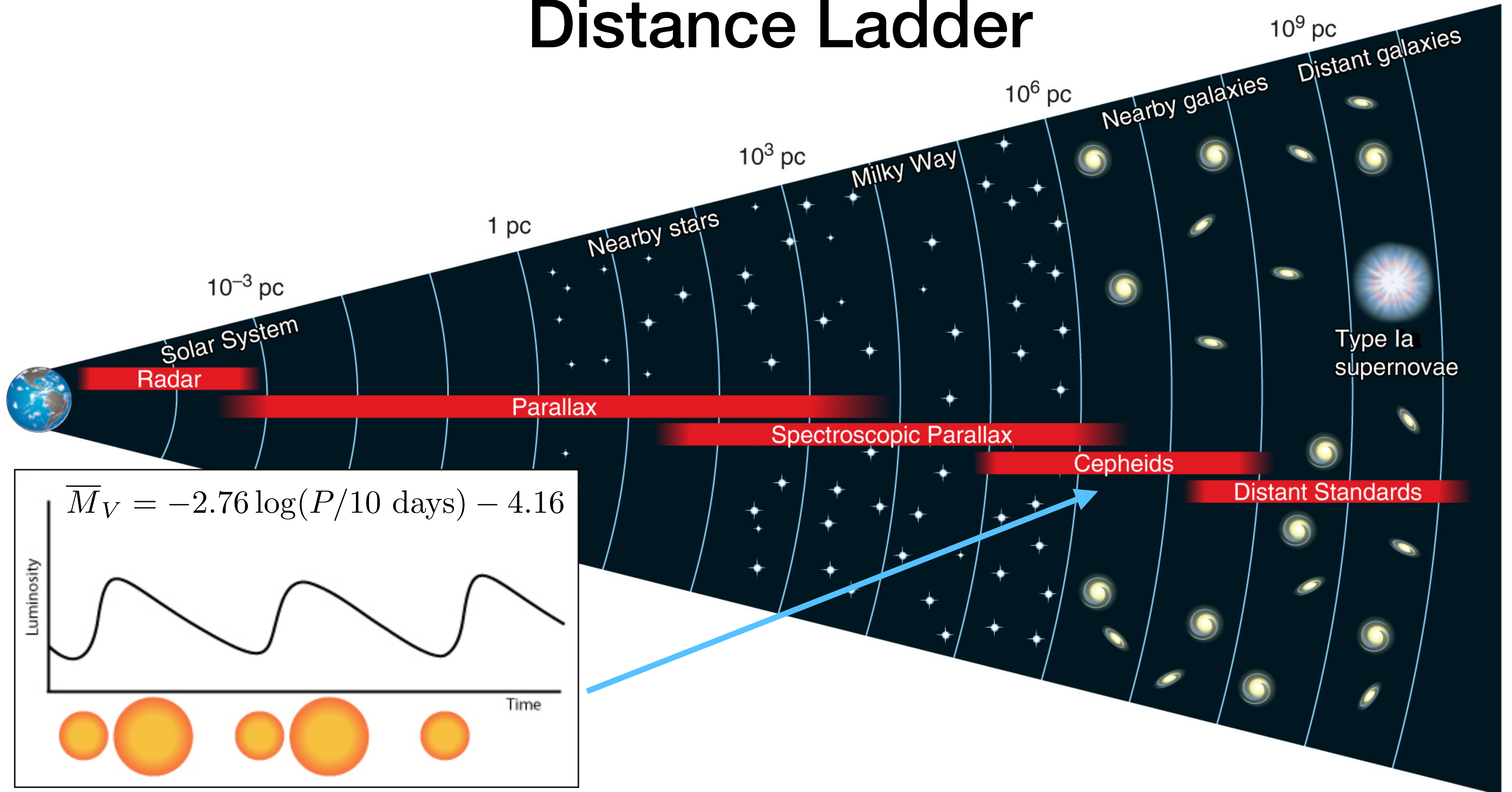


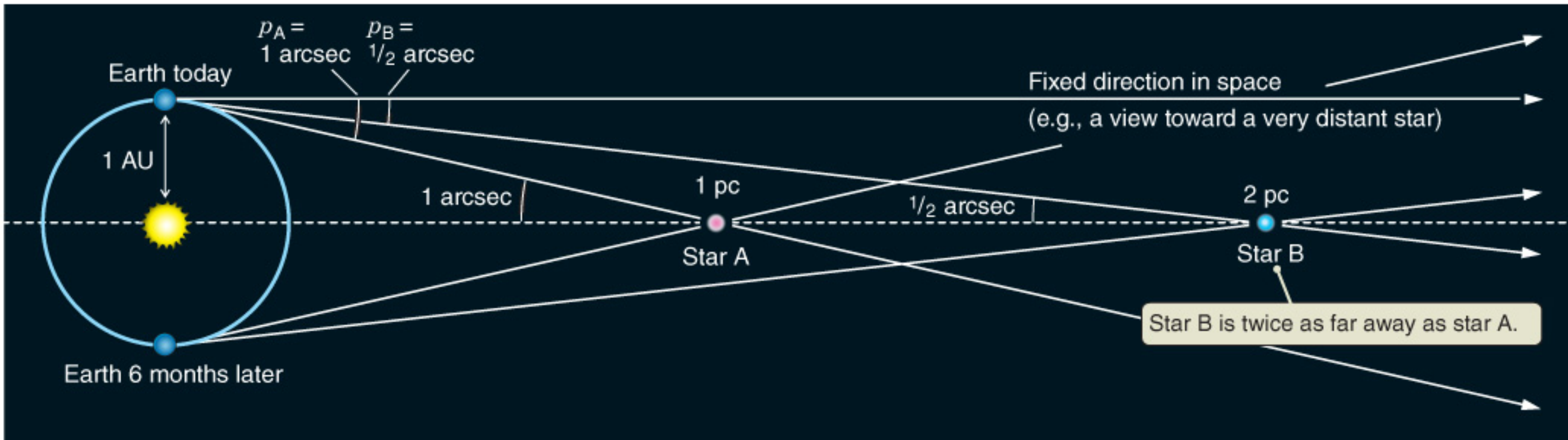
Figure 1.1

Figure 1.2

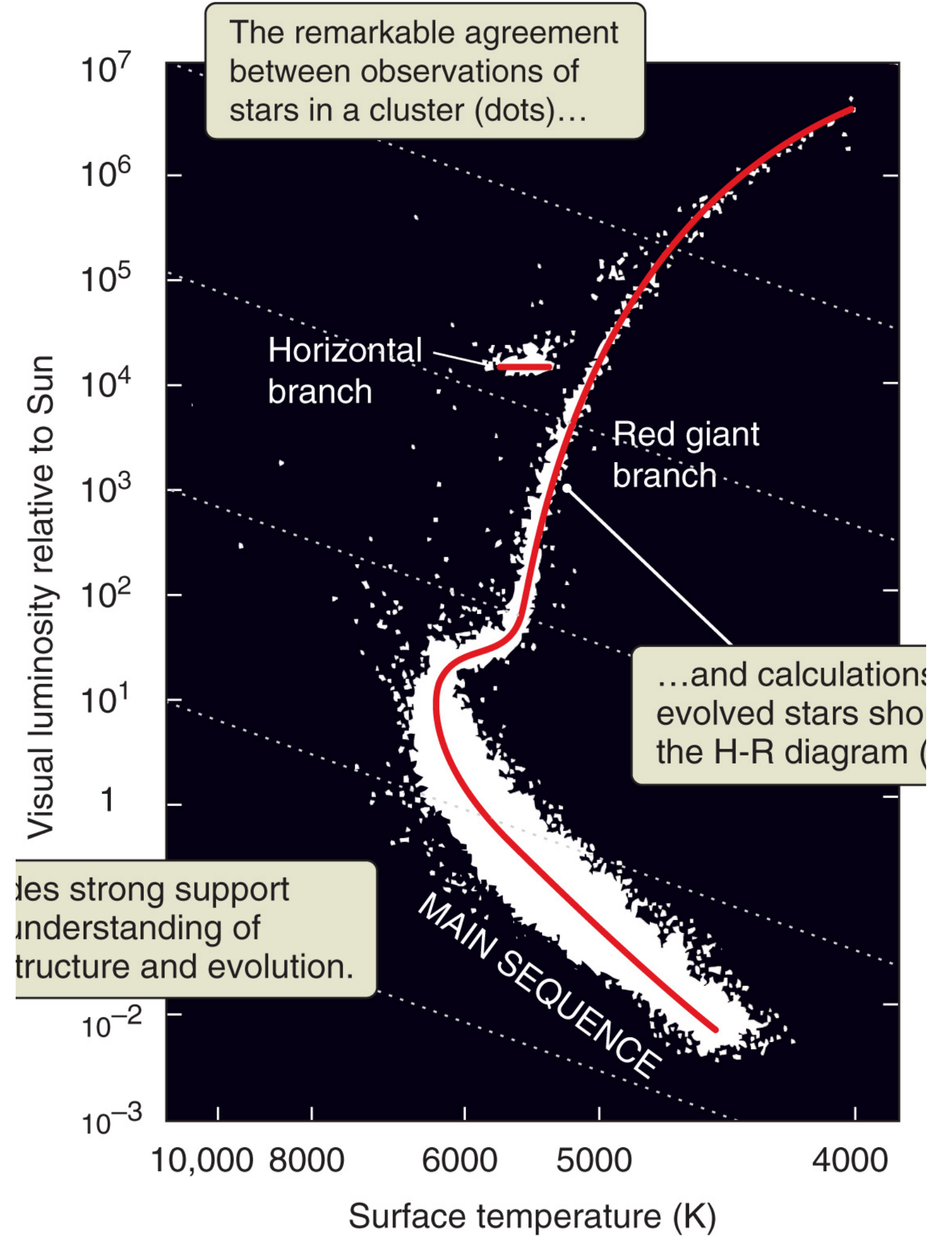
Distance Ladder



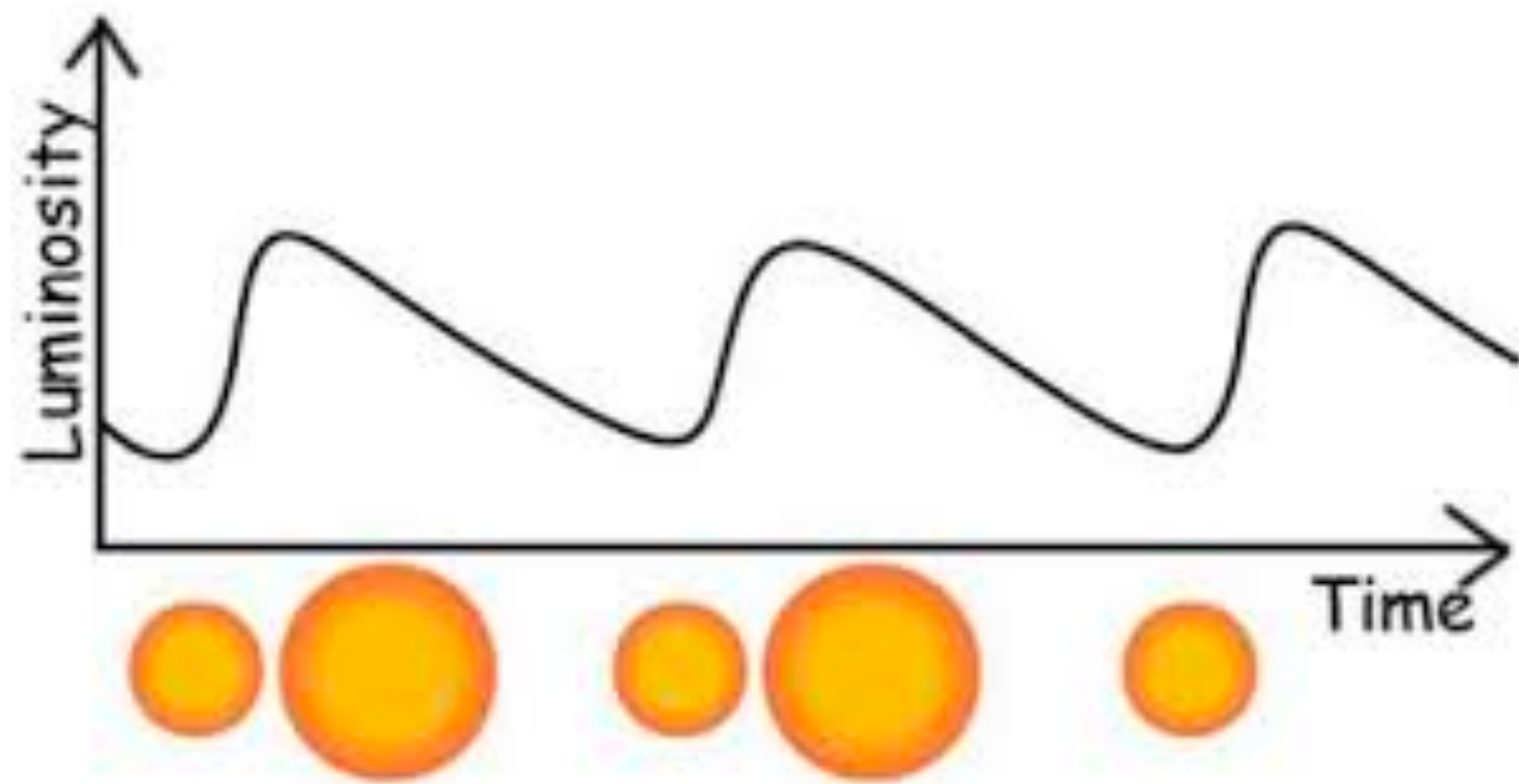
Parallax



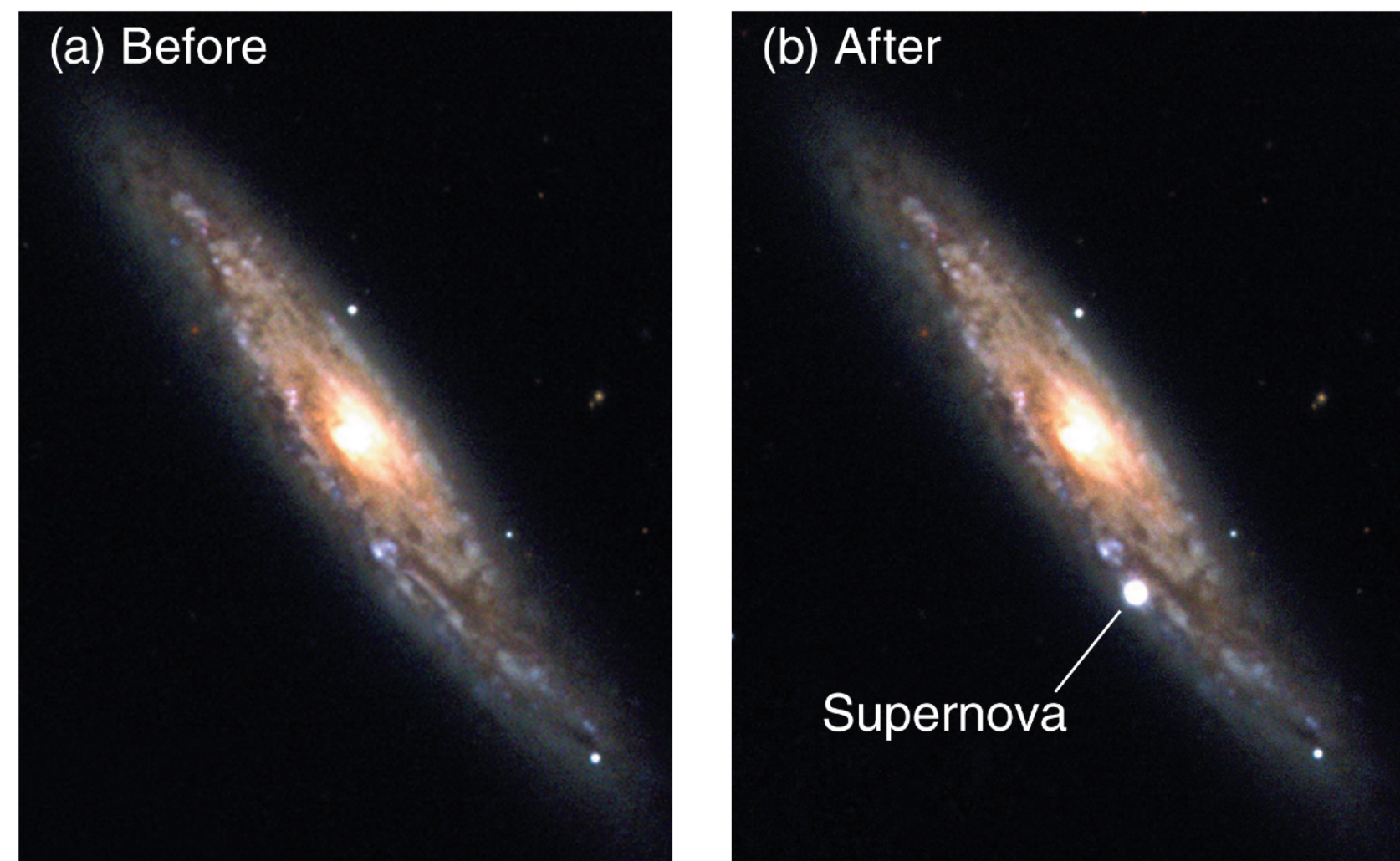
Spectroscopic Parallax



Cepheid Variables

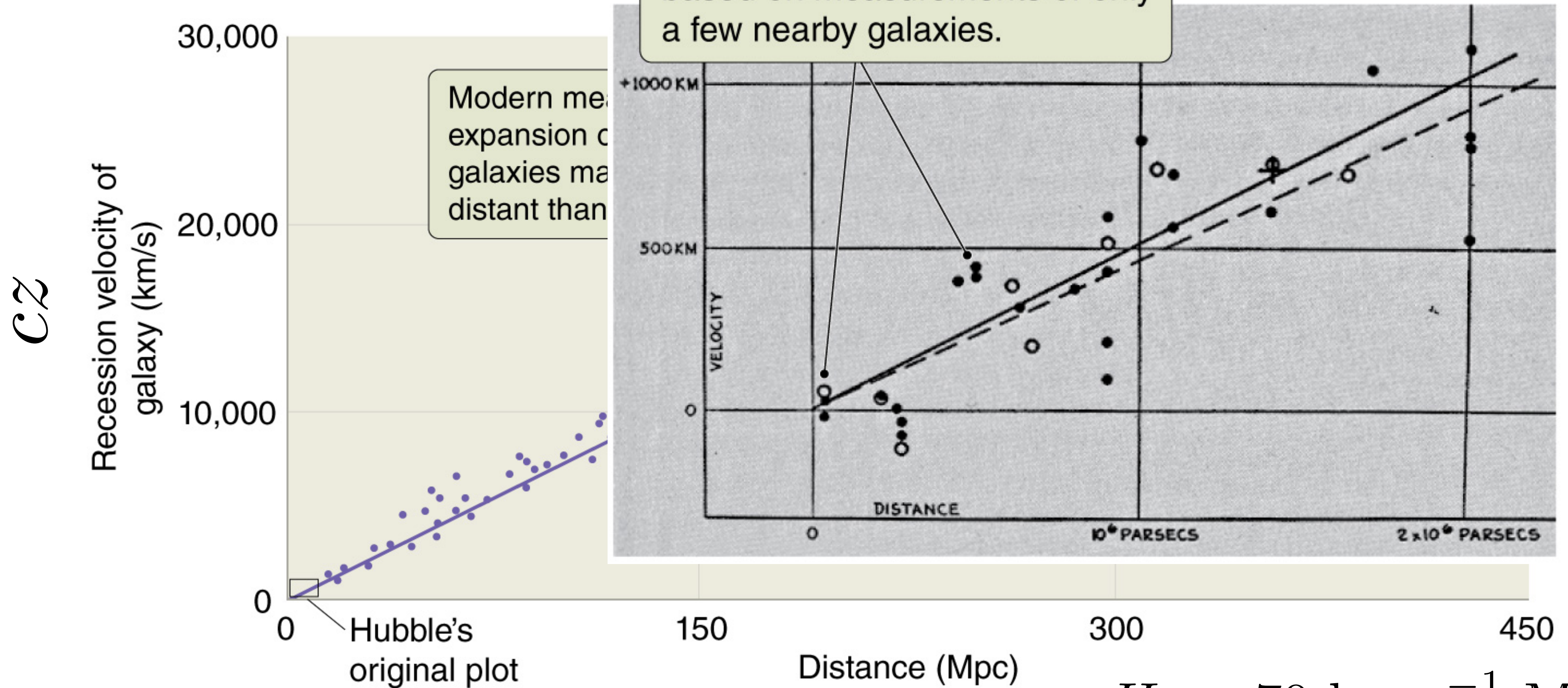


Type Ia SNe



Hubble's Law

$$cz = H_0 d$$

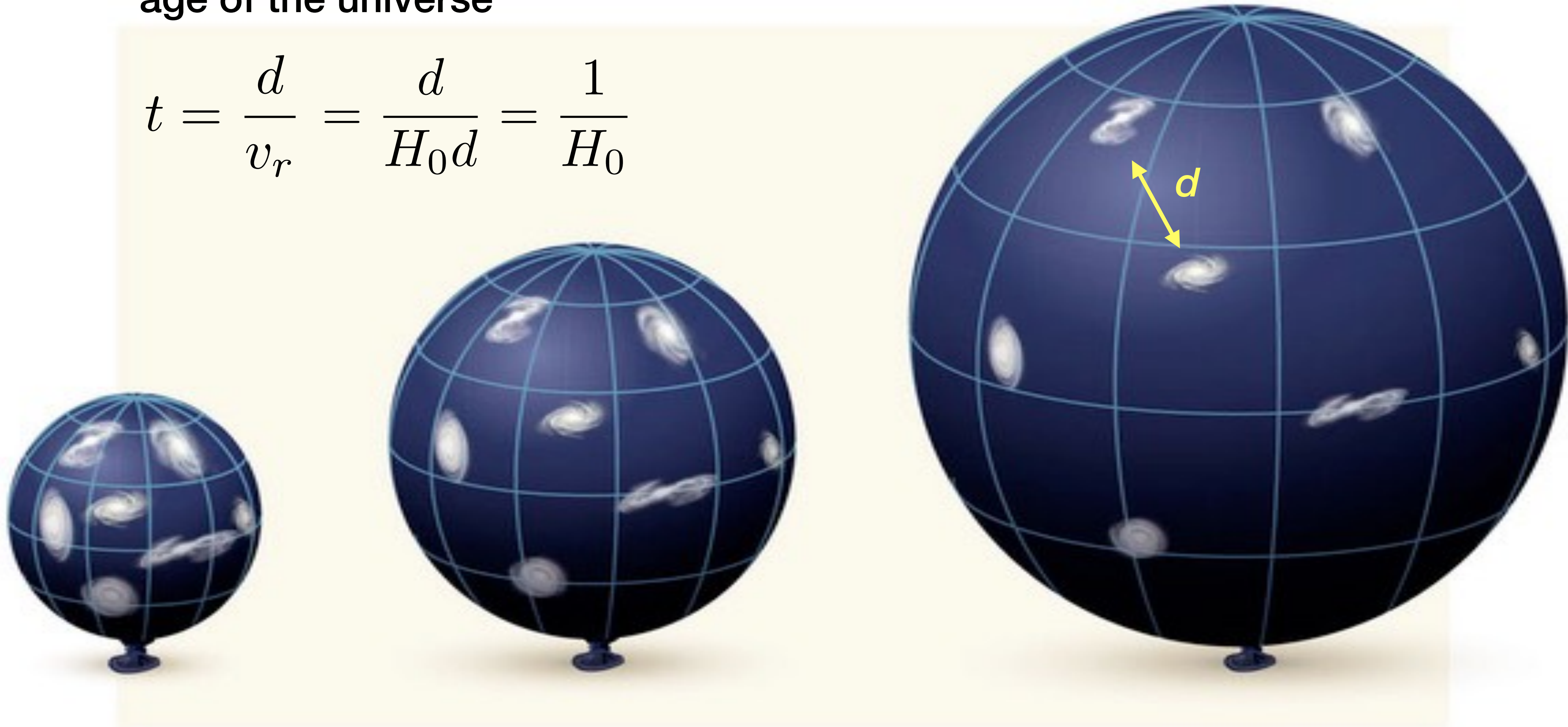


$$H_0 \approx 70 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

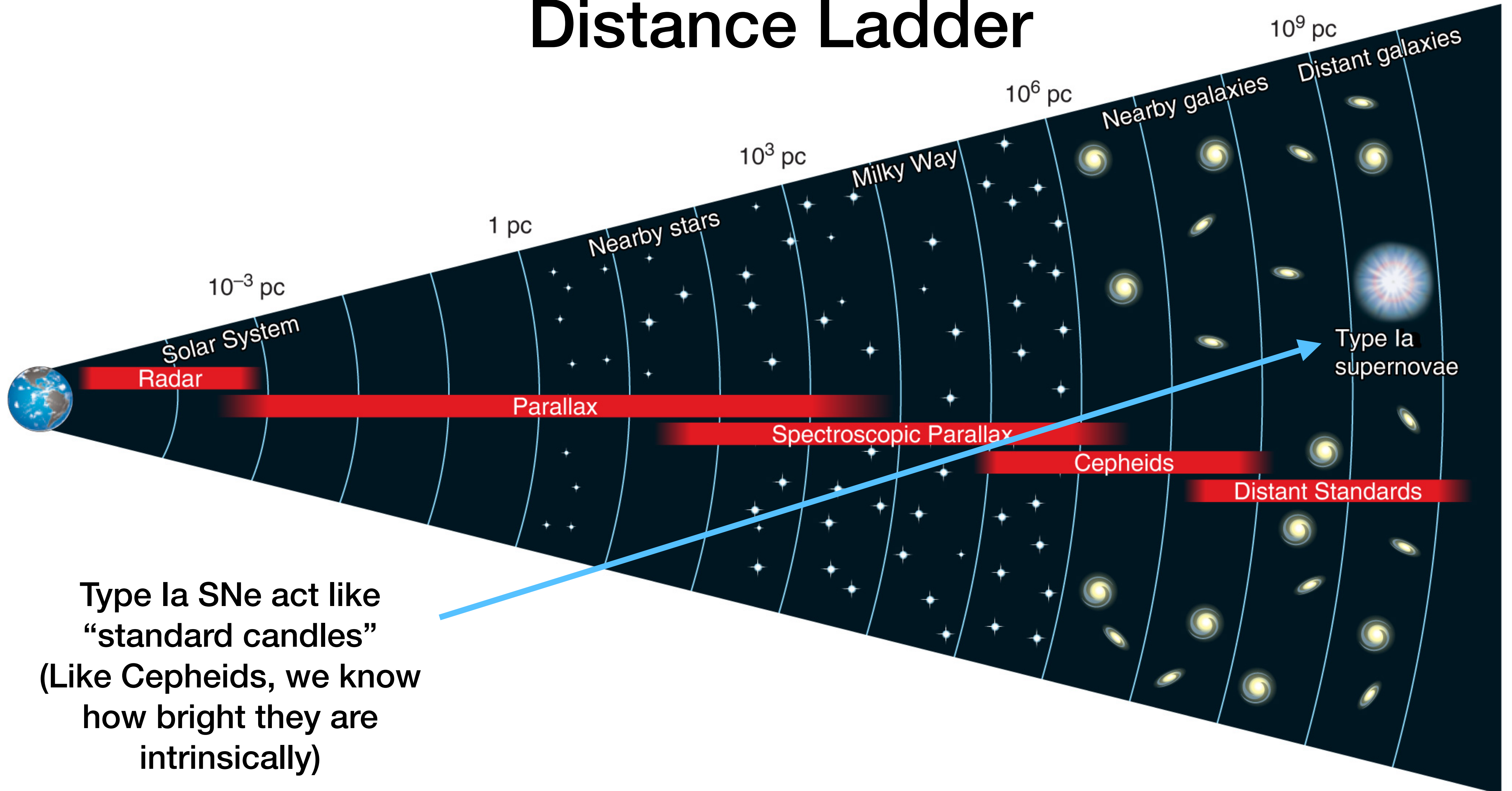
We live in an expanding “balloon universe”

If expansion constant, then can estimate the age of the universe

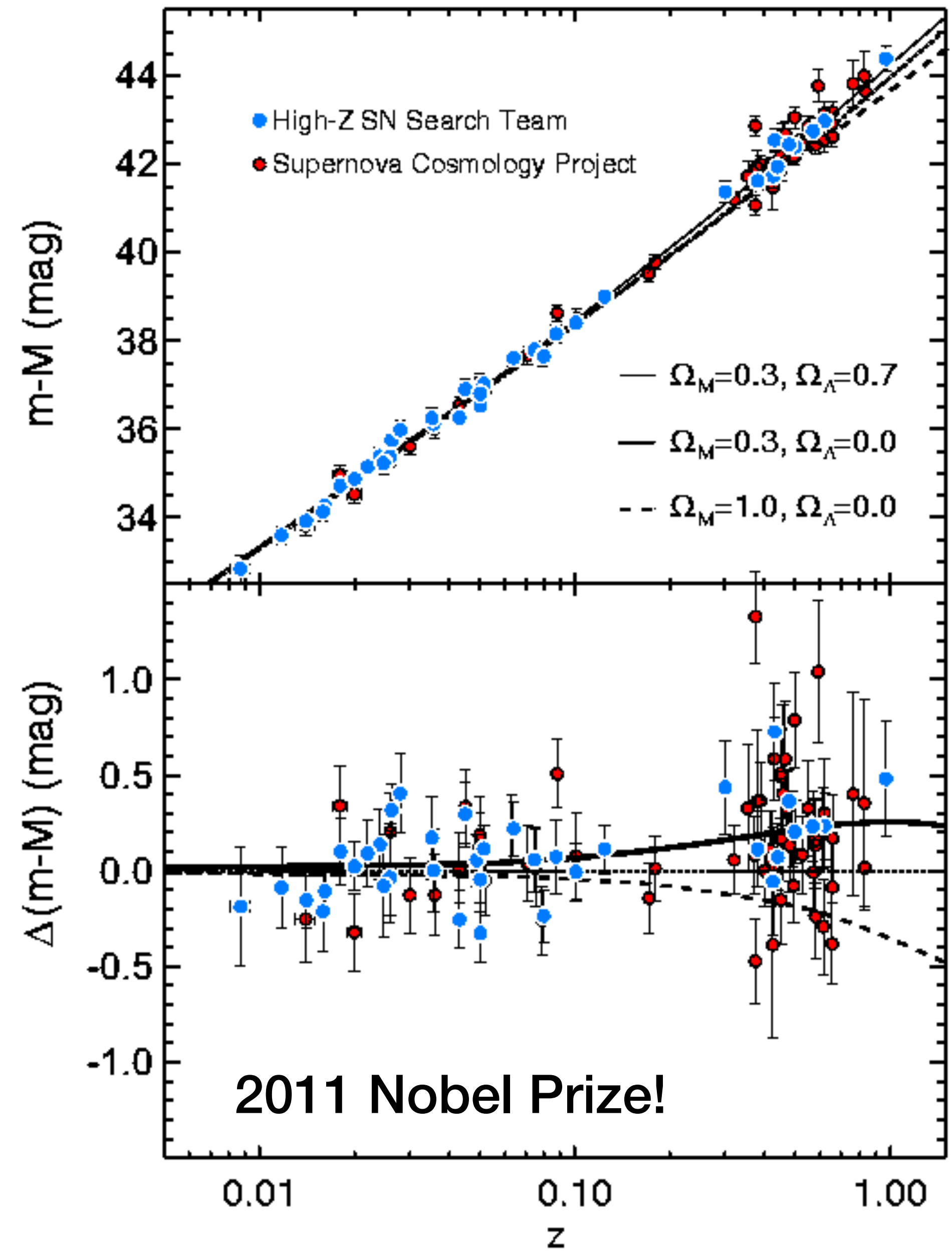
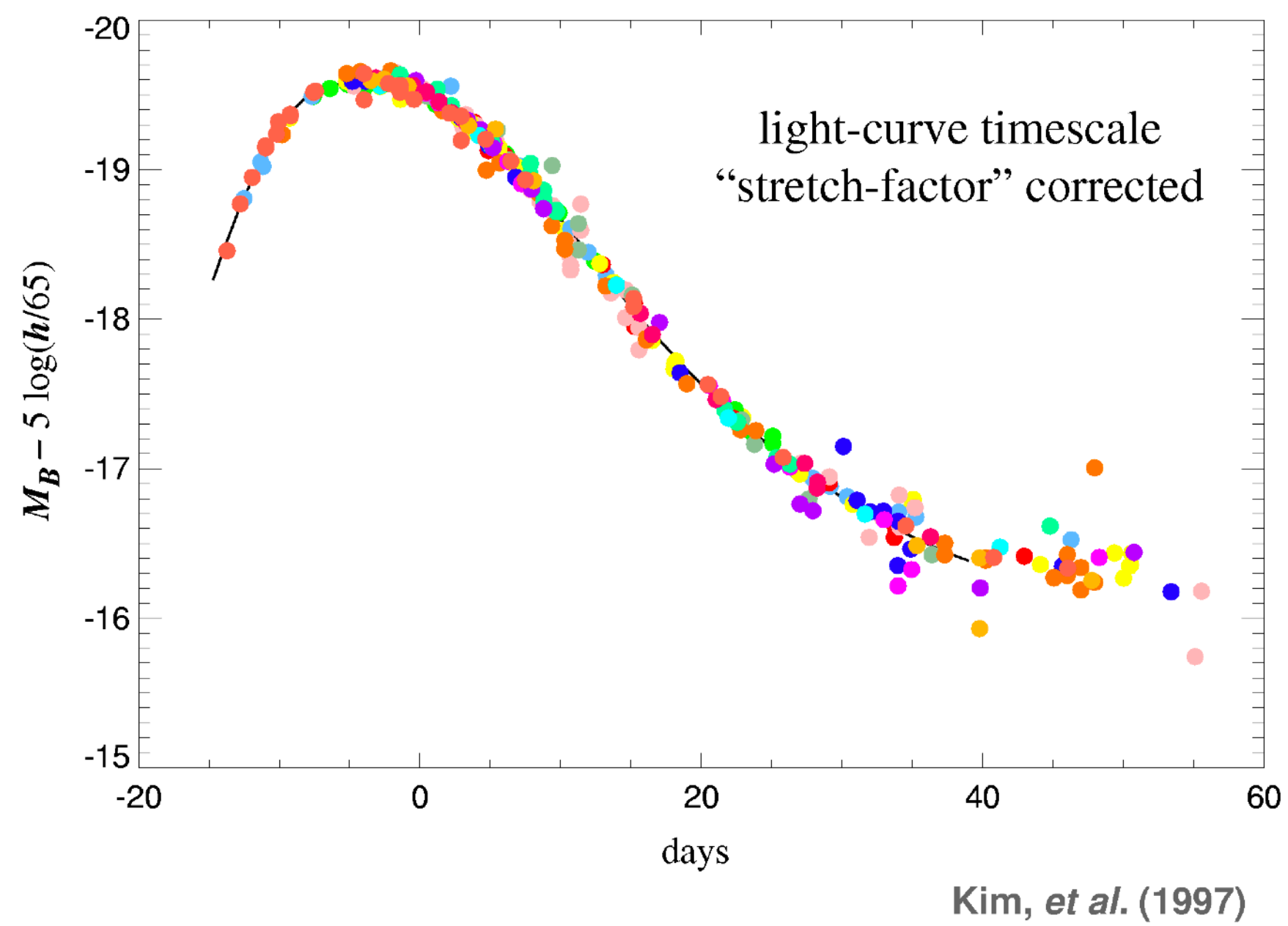
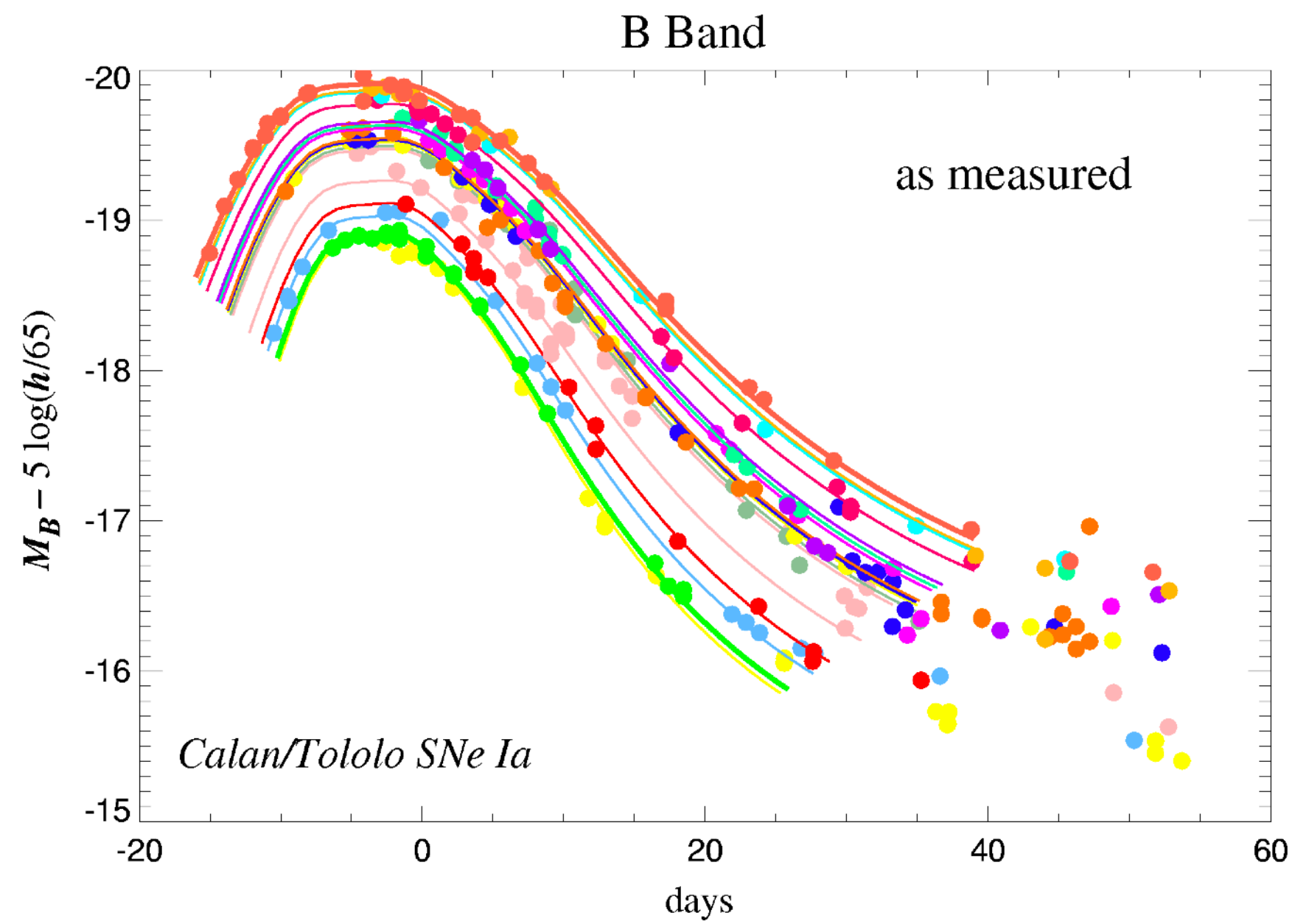
$$t = \frac{d}{v_r} = \frac{d}{H_0 d} = \frac{1}{H_0}$$



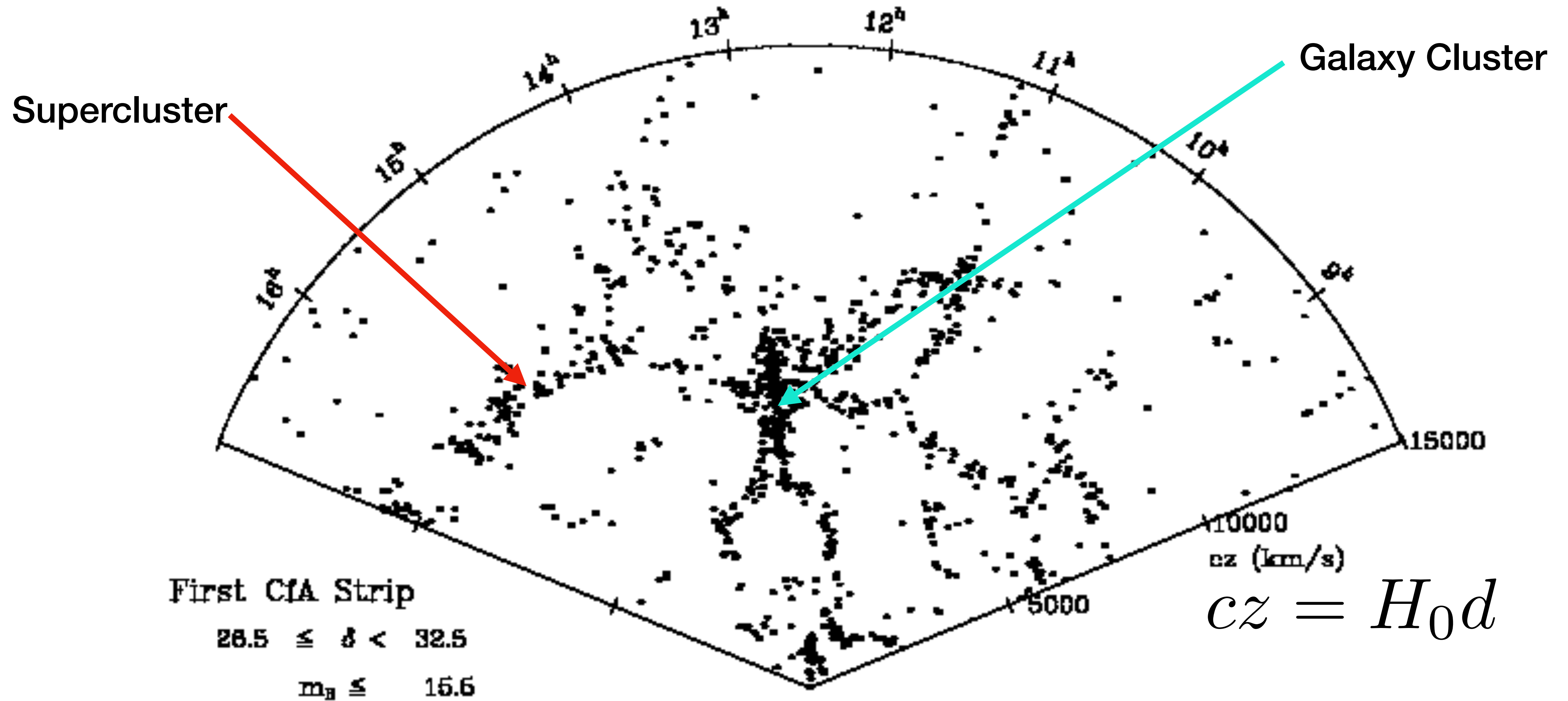
Distance Ladder



Type Ia SNe act like “standard candles” (Like Cepheids, we know how bright they are intrinsically)



Finger of God: the Coma Cluster



Galaxy Surveys

Why does the pattern look different at high redshifts?

