## Homework 11

## Due April 22 at 10:45am via Canvas

Please show all work, writing solutions/explanations clearly, or no credit will be given. You are encouraged to work together, but everyone must turn in independent solutions: do not copy from others or from any other sources.

- 1. For the Schechter luminosity function of galaxies, find the total luminosity density  $\Psi$  as a function of  $L^*$ ,  $\Phi^*$ , and  $\alpha$ . What is the numerical value of the luminosity density  $\Psi_V$  in the V band, given  $L_V^* = 2 \times 10^{10} L_{\odot,V}$ ,  $\Phi^* = 0.005 \text{ Mpc}^{-3}$ , and  $\alpha = -1$ ?
- 2. On a mass scale  $M = 10^{17} \text{ M}_{\odot}$ , the root mean square mass fluctuation is  $\delta M/M = 0.12$  today (inferred from Ryden Figure 11.5). Do you expect to see any gravitationally collapsed structures with that mass in the directly visible universe today? Explain why or why not.
- 3. Consider the Big Rip scenario from HW 4, Problem #2 (http://www.astro.utah.edu/~wik/courses/astr4080spring2021/hw/hw04. pdf). Since the energy density has the dependence  $\varepsilon \propto a^{-3(1+w)}$ , the energy density  $\varepsilon_p$  (with  $w_p < -1$ ) of phantom energy increases as the universe expands; when  $\varepsilon_p/c^2$  becomes larger than the mass density of a bound object, the object will be ripped apart. Suppose that the universe contains both matter and phantom energy with  $w_p = -1.1$ .
  - (a) If the density parameters of the two components are  $\Omega_{m,0} = 0.3$  and  $\Omega_{p,0} = 0.7$ , at what scale factor  $a_{\text{gal}}$  will the descendent galaxy of the merger between the Milky Way and M31 be ripped apart?
  - (b) At what scale factor  $a_*$  will a Sun-like star be ripped apart?
  - (c) Using the result from the HW 4 problem, how many years before the Big Rip will that Sun-like star be ripped apart? [Hint: the Big Rip is defined as the time  $t_{\rm rip}$  when  $a \to \infty$ .]