

# ASTR/PHYS 4080: Introduction to Cosmology



## Week 4

Read thru Chapter 5  
Also read the Key Concepts for those chapters

Today: Eqn of States & Model Universes

HW 3 due Thursday  
HW 4 available

Only available first 20min  
on Tuesday

Office Hours:  
Tuesdays 2-3pm  
Wednesdays 11am-12pm

# Equations of the Universe

Friedmann Eqn: gravity

$$\left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3c^2}\epsilon - \frac{\kappa c^2}{R_0^2 a^2} + \cancel{\frac{\Lambda}{3}}$$

$$\epsilon_\Lambda \equiv \frac{c^2}{8\pi G}\Lambda$$

Ideal Gas Law

$$PV = NRT$$

$$P = nkT$$

$$P = \frac{\rho}{\mu}kT$$

$$\epsilon = \rho c^2$$

Fluid Eqn: conserve energy

$$\dot{\epsilon} + 3\frac{\dot{a}}{a}(\epsilon + P) = 0$$

$$P \approx \frac{kT}{\mu c^2}\epsilon$$

Eqn of State:

relations b/t thermodynamic properties

$$P = w\epsilon$$

# Acceleration Equation

$$\left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3c^2}\varepsilon - \frac{\kappa c^2}{R_0^2 a^2} + \frac{\Lambda}{3} \quad \xrightarrow{+} \quad = \quad \xleftarrow{+} \quad \dot{\varepsilon} + 3\frac{\dot{a}}{a}(\varepsilon + P) = 0$$

$$\frac{\ddot{a}}{a} = \frac{4\pi G}{3c^2}(\varepsilon + 3P) + \frac{\Lambda}{3}$$

$$P < -\frac{1}{3}\varepsilon \quad \longrightarrow \quad \ddot{a} > 0$$

$$\Lambda = 0$$

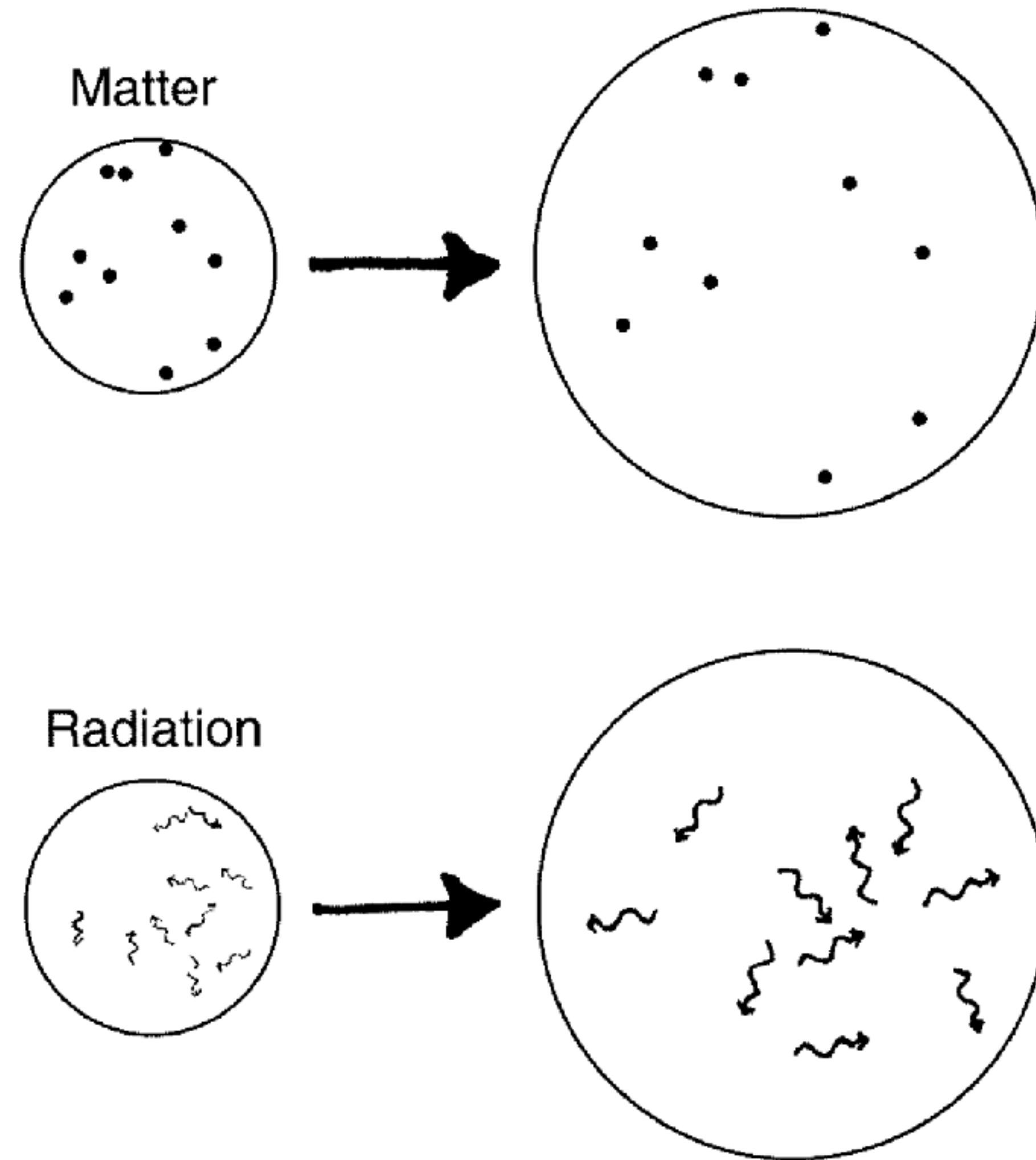
If  $\Lambda$  const,  $\varepsilon$  also const b/c  $\varepsilon_\Lambda \equiv \frac{c^2}{8\pi G}\Lambda$

so  $P = -\varepsilon_\Lambda = -\frac{c^2}{8\pi G}\Lambda$

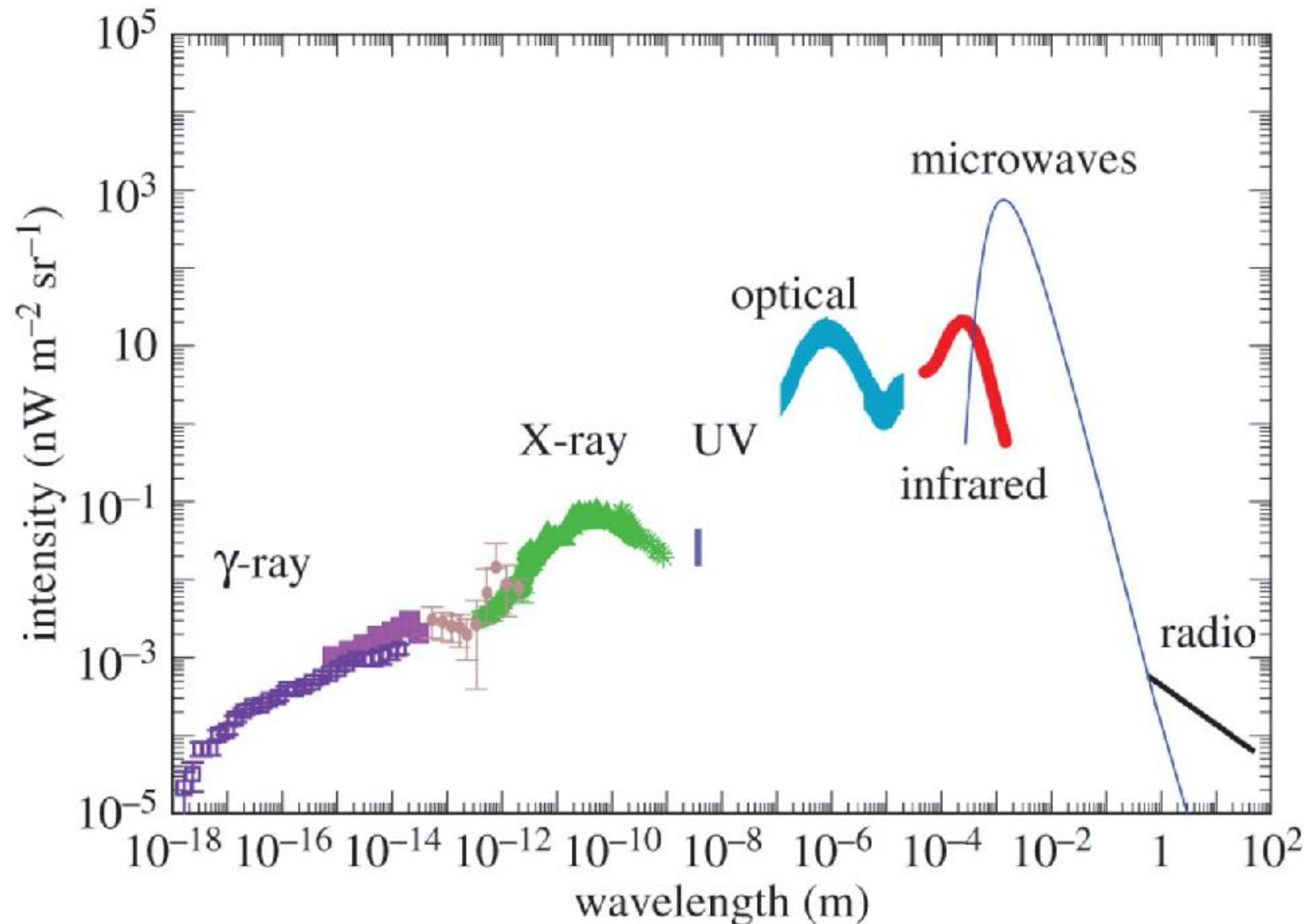
# Evolution of Components

$$\varepsilon = \sum_i \varepsilon_i$$

$$P = \sum_i w_i \varepsilon_i$$



# Relativistic Component: “Radiation” = Photons + Neutrinos



CMB dominates energy density of photons in universe today

Don't know the mass of neutrinos, but energy density dominated by momentum and decoupling in early universe. Calculated to be:

$$\varepsilon_\nu = \frac{7}{8} \left( \frac{4}{11} \right)^{4/3} \varepsilon_{\text{CMB}} = 0.227 \varepsilon_{\text{CMB}}$$

True for each neutrino flavor, so total is 3x above.

$$\varepsilon_{\text{CMB},0} = \alpha T_0^4 = 0.2606 \text{ MeV m}^{-3}$$

# Relativistic, Non-rel., & $\Lambda$ Components

Radiation

$$\varepsilon_{\text{CMB},0} = \alpha T_0^4 = 0.2606 \text{ MeV m}^{-3}$$

$$\varepsilon_{\text{crit},0} = 4870 \text{ MeV m}^{-3}$$

$$\Omega_{\text{r},0} = \frac{1.681 \varepsilon_{\text{CMB},0}}{\varepsilon_{\text{crit},0}} = 9 \times 10^{-5}$$

(neutrinos have ~68% of the energy density of CMB photons)

Matter:

$$\Omega_{\text{m},0} \approx 0.31$$

Dark Energy:

$$\Omega_{\Lambda,0} \approx 0.69$$

Using evolution of each component with  $a$ , can compute when matter-dark energy and matter-radiation components were comparable

# Model Universes

Can now solve for  $a(t)$  generically, if not necessarily analytically:

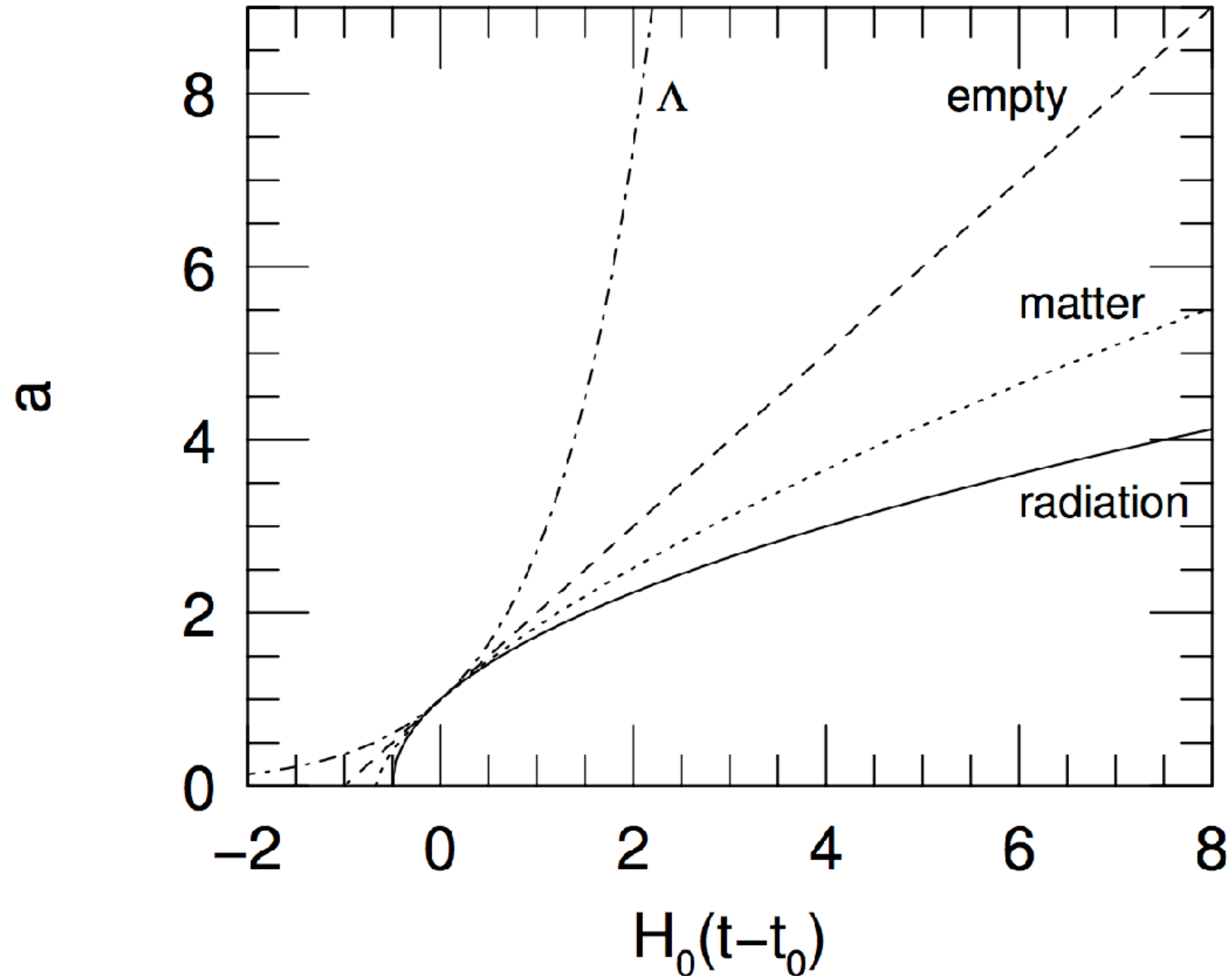
$$\dot{a}^2 = \frac{8\pi G}{3c^2} \sum_i \varepsilon_{i,0} a^{-1-3w_i} - \frac{\kappa c^2}{R_0^2}$$

- Empty
- Matter only
  - classic case: open, closed, flat
- Radiation only (+curvature)
- Lambda only (+curvature)
- Various Combinations!

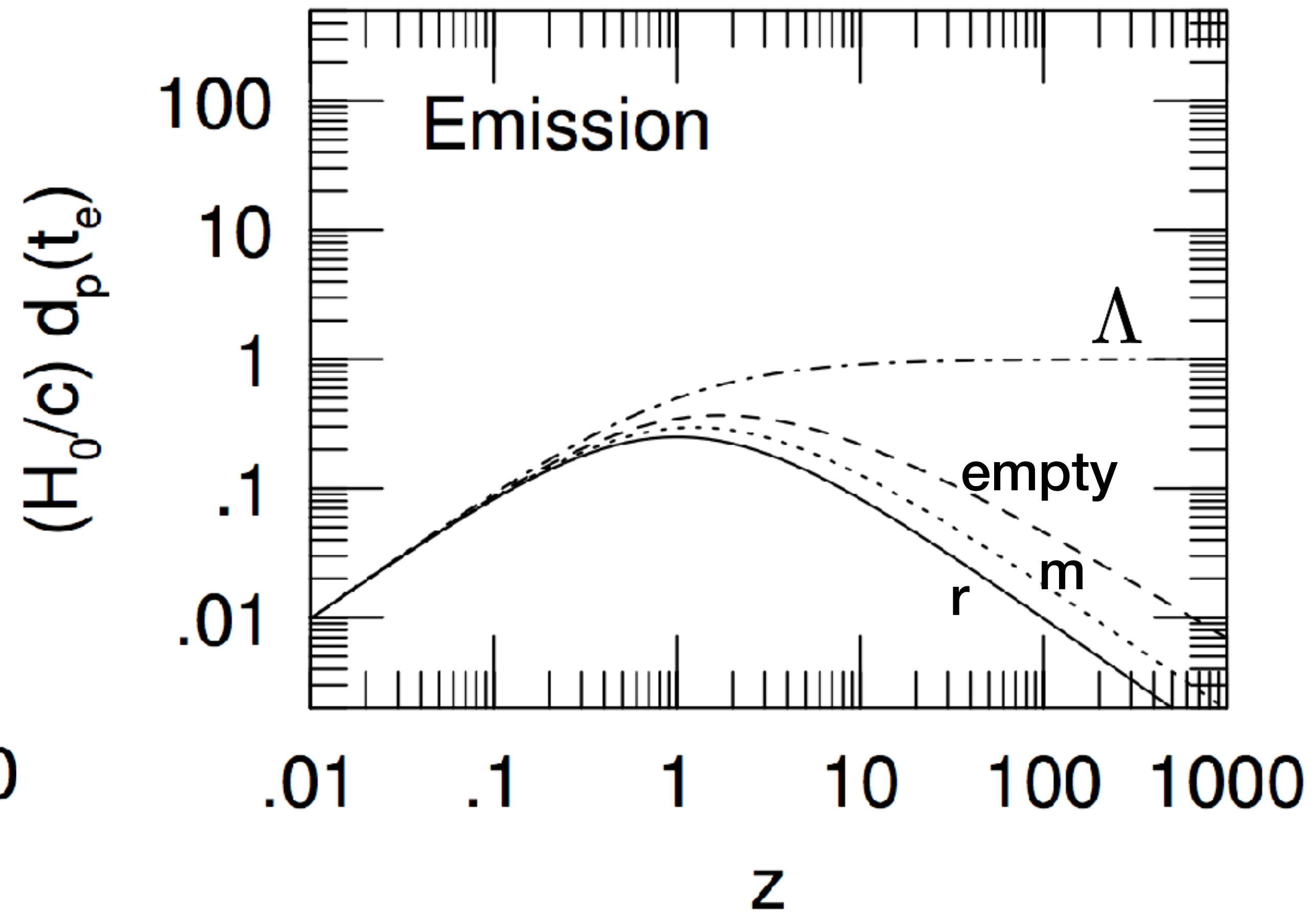
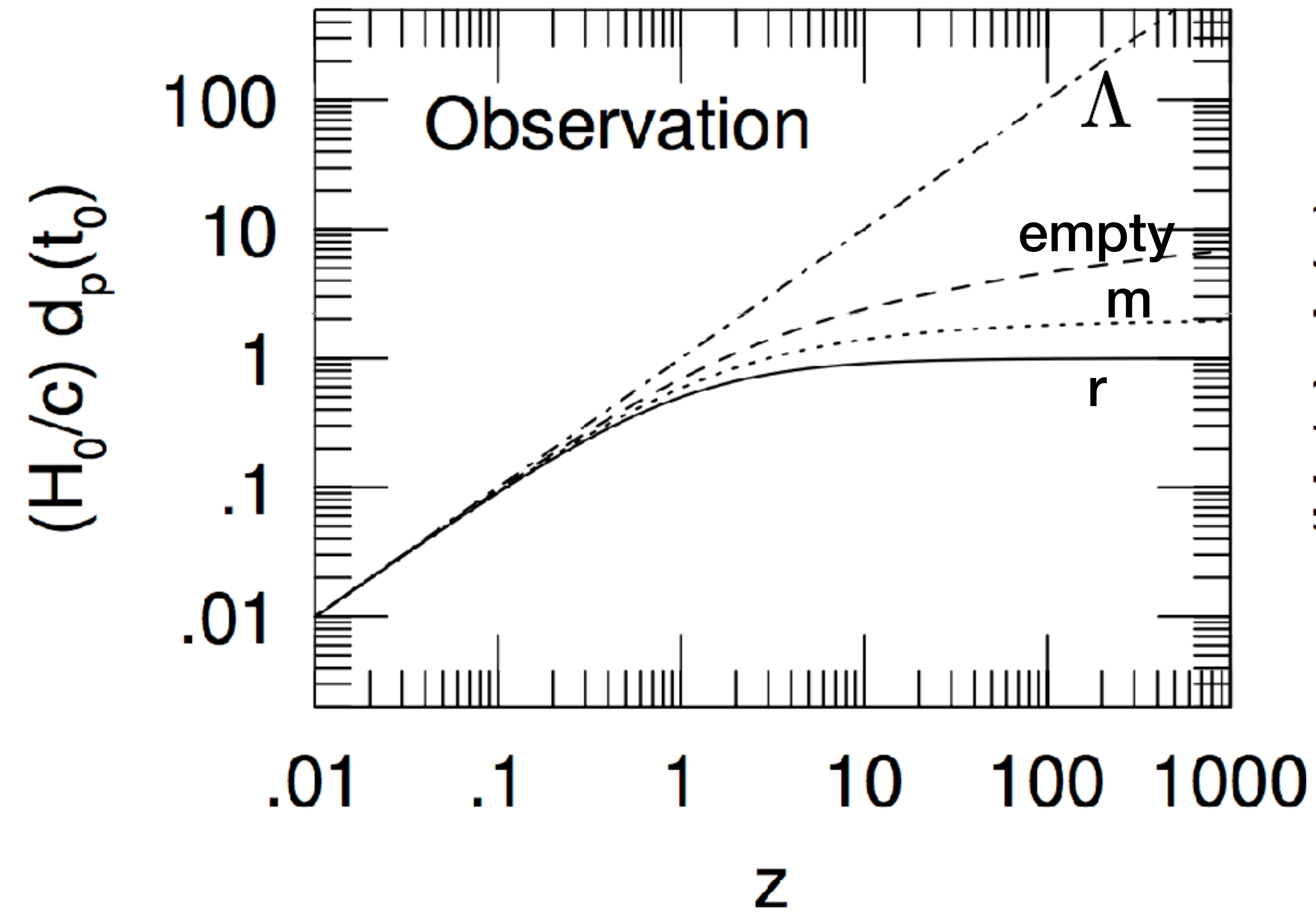
# Empty



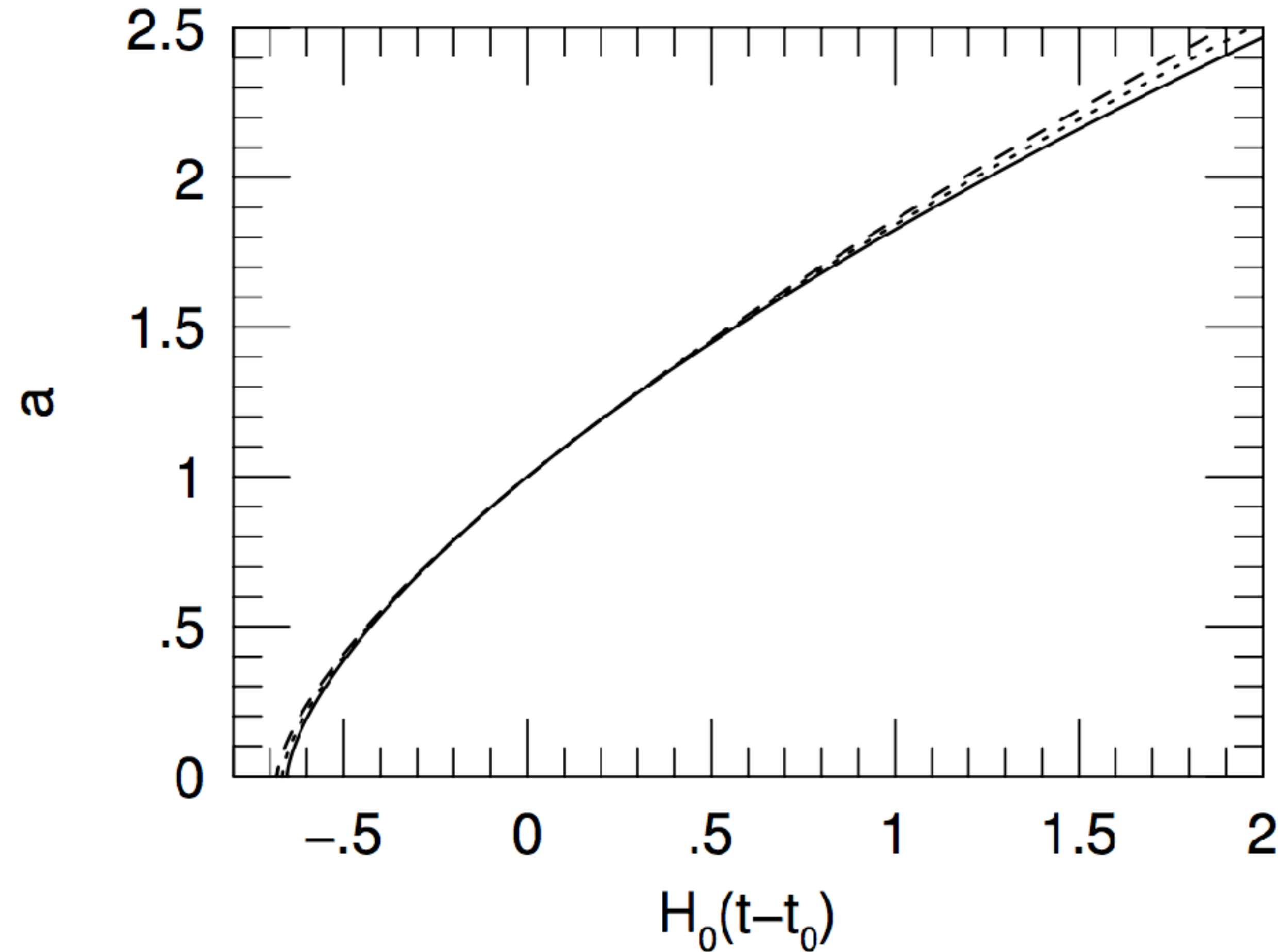
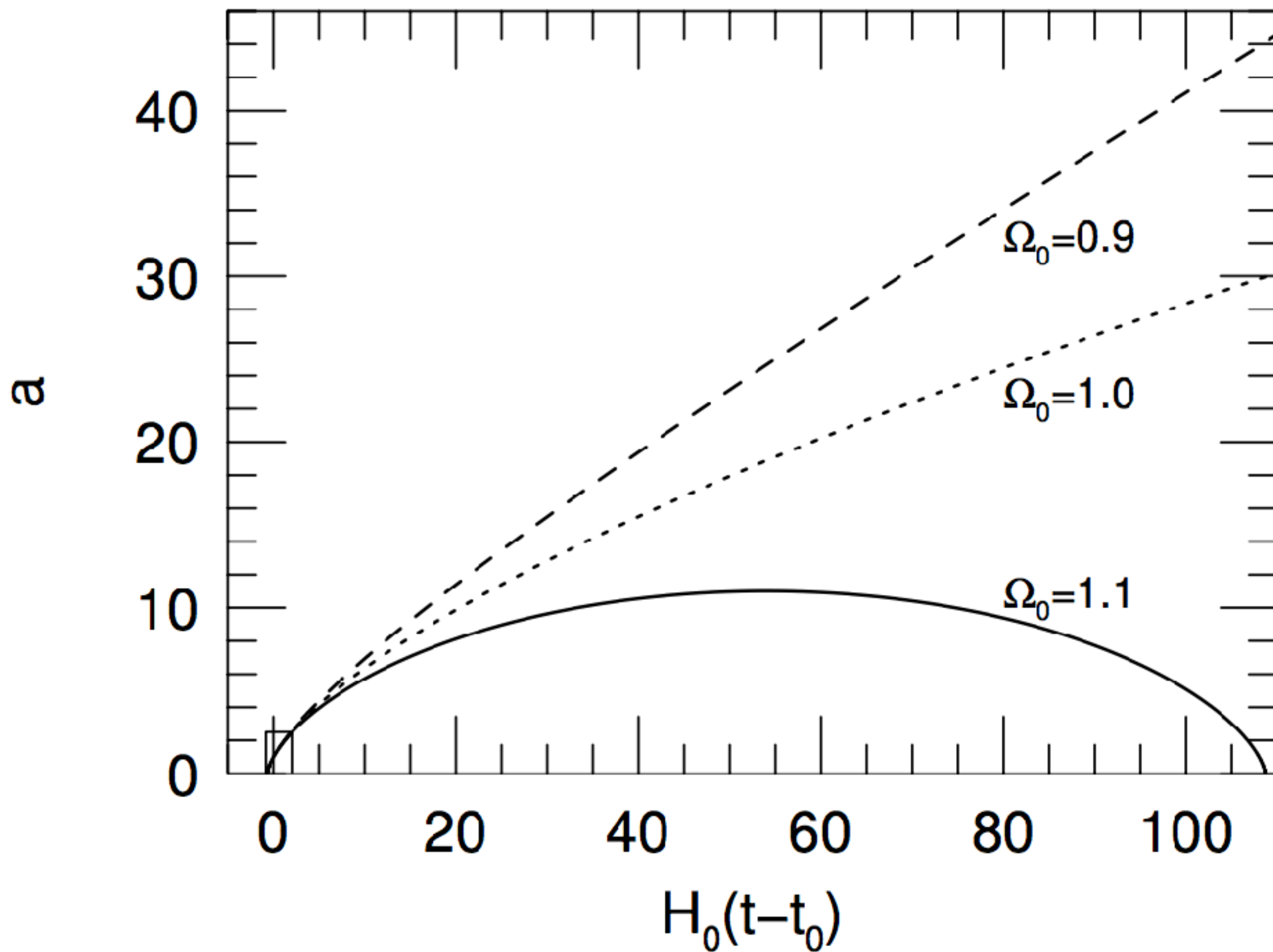
# Only 1 Constituent in a Flat Spacetime



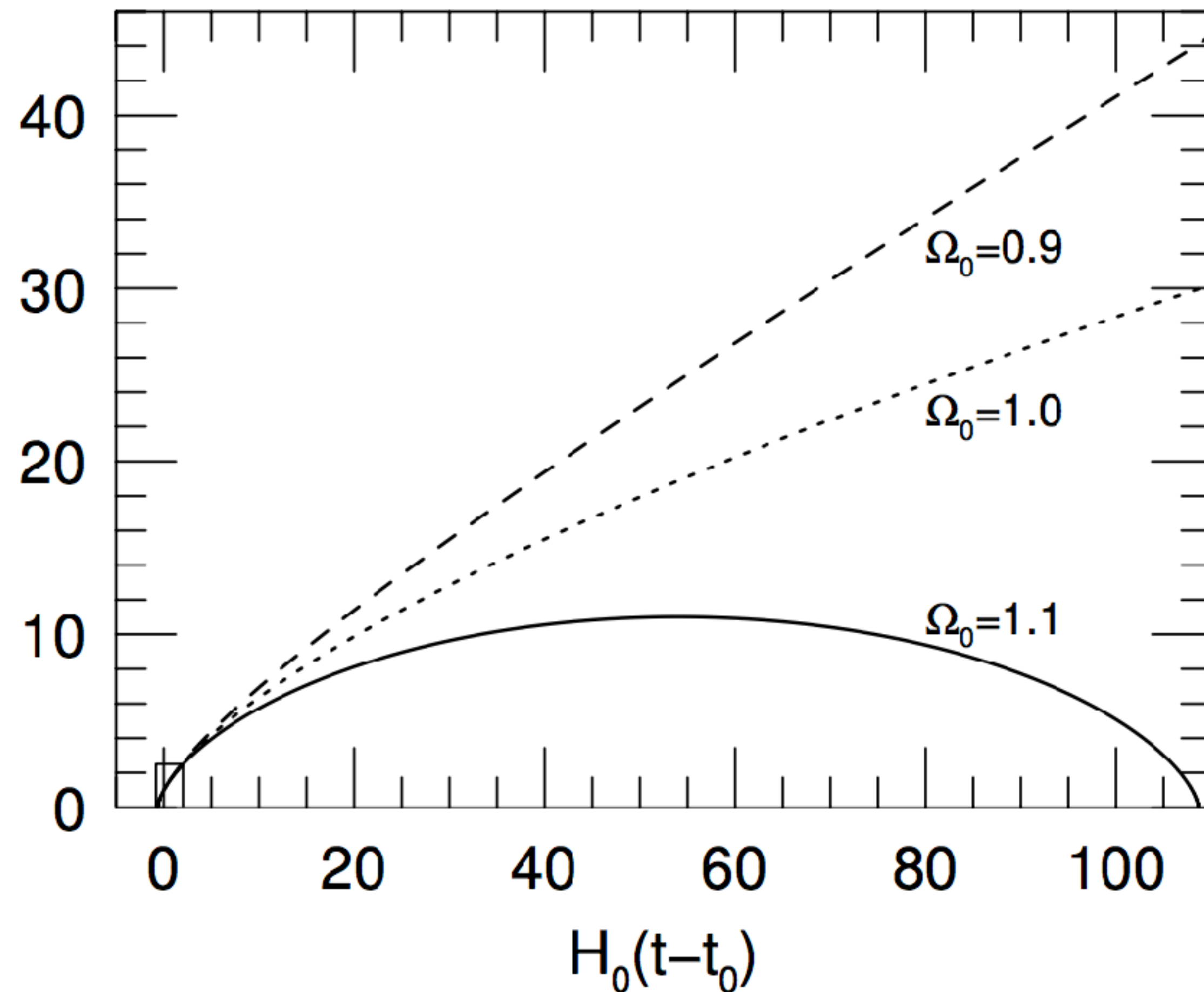
# Only 1 Constituent in a Flat Spacetime



# Classical Cosmology: Matter + Curvature



# Classical Cosmology: Matter + Curvature



$$\kappa = 0, \quad \Omega_0 = 1; \quad a(t) = (t/t_0)^{2/3}$$

$$\kappa = +1, \quad \Omega_0 > 1; \quad a(\theta) = \frac{1}{2} \frac{\Omega_0}{\Omega_0 - 1} (1 - \cos \theta)$$

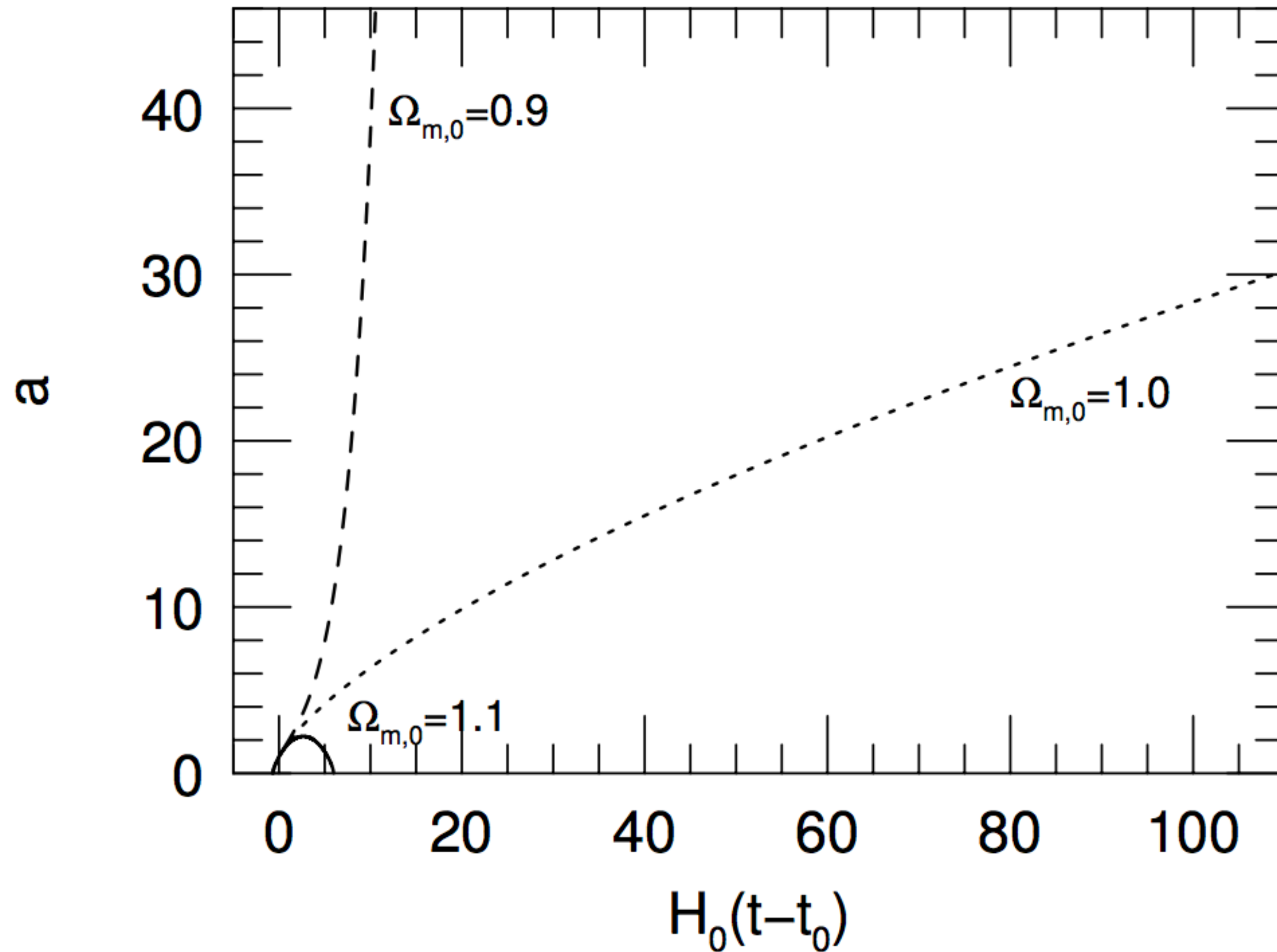
$$t(\theta) = \frac{1}{2H_0} \frac{\Omega_0}{(\Omega_0 - 1)^{3/2}} (\theta - \sin \theta)$$

$$\kappa = -1, \quad \Omega_0 < 1; \quad a(\eta) = \frac{1}{2} \frac{\Omega_0}{1 - \Omega_0} (\cosh \eta - 1)$$

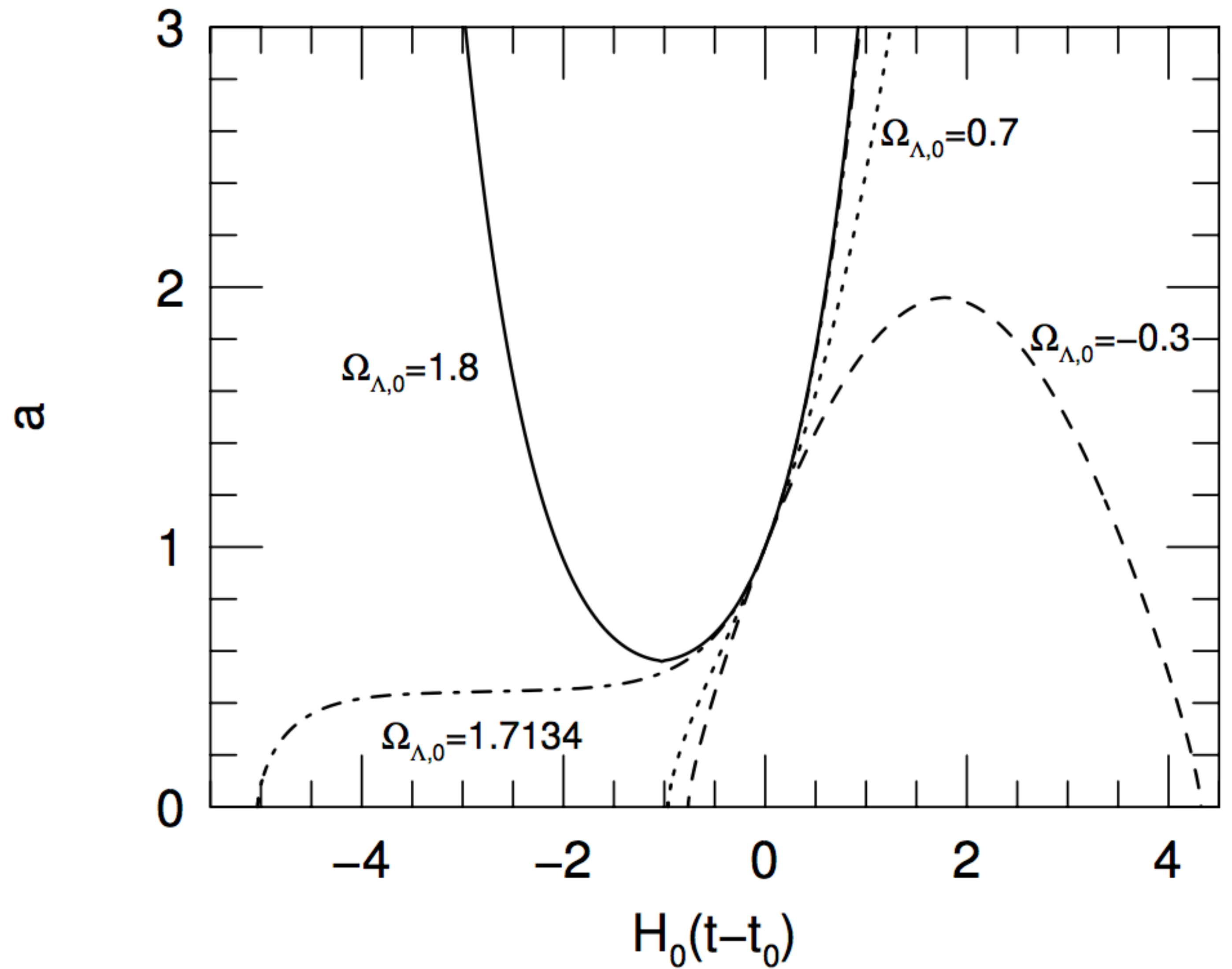
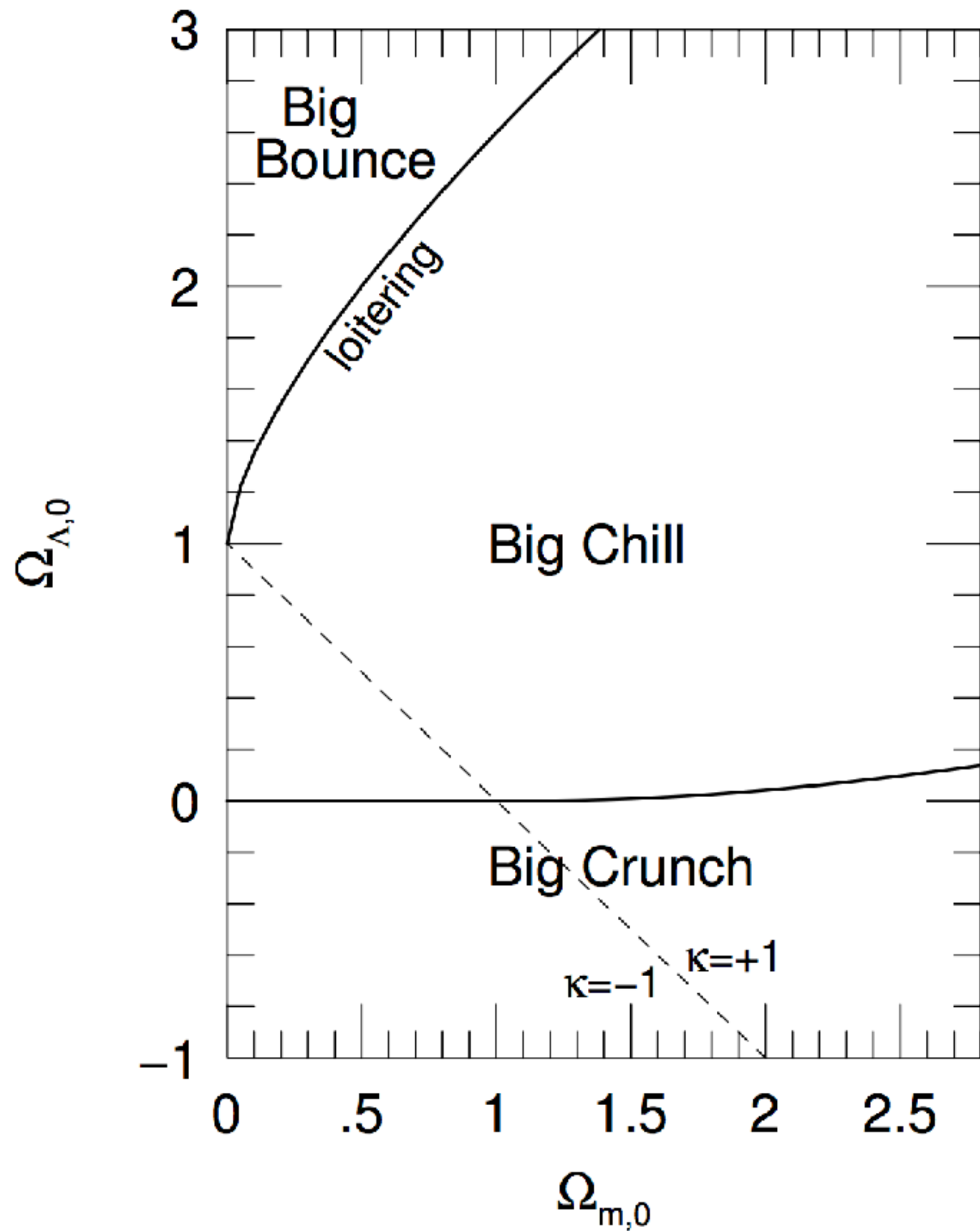
$$t(\eta) = \frac{1}{2H_0} \frac{\Omega_0}{(1 - \Omega_0)^{3/2}} (\sinh \eta - \eta)$$

**Density  $\rightarrow$  Destiny**

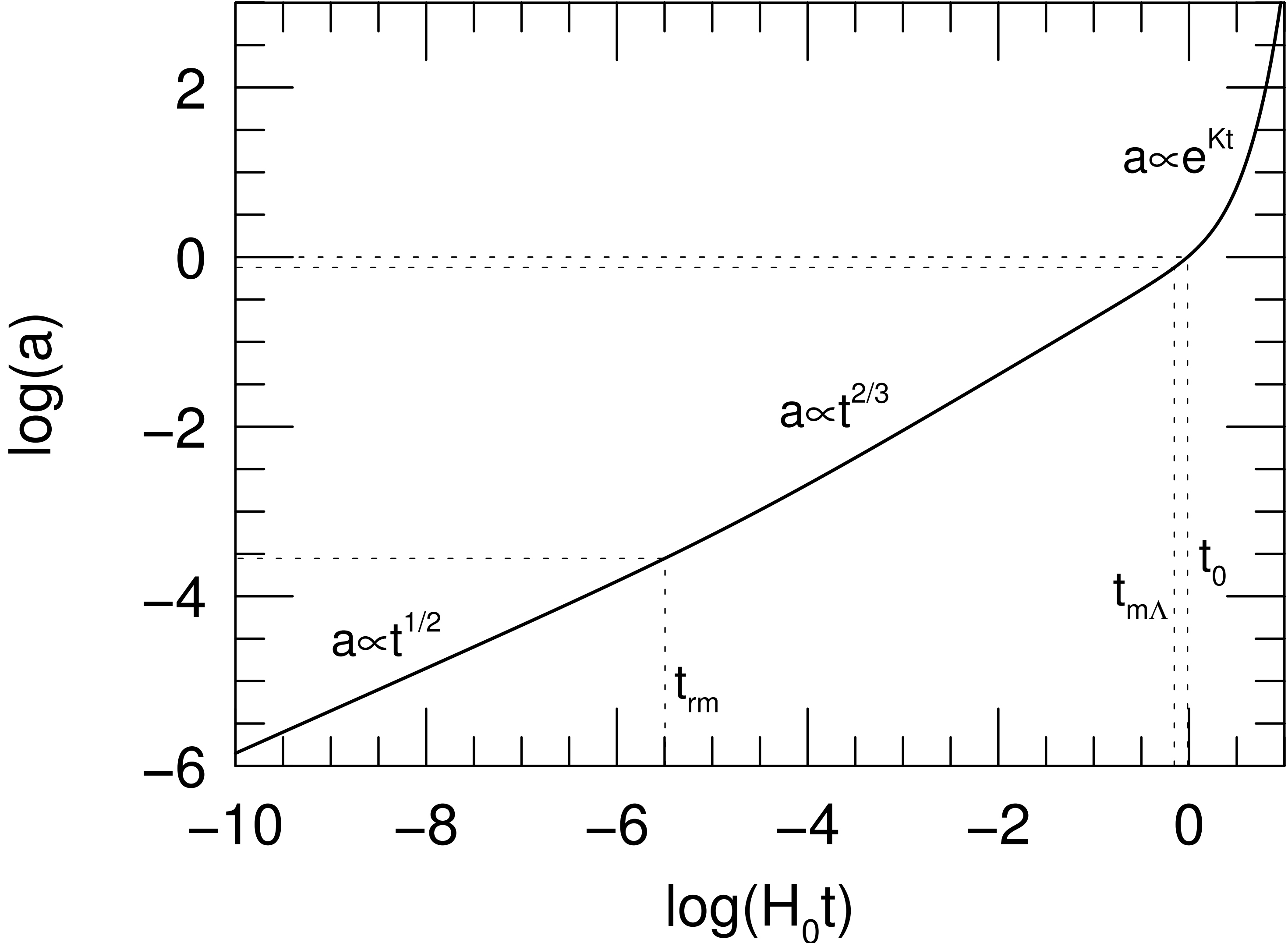
# Flat: Matter + Cosmological Constant



# Matter + Lambda + Curvature



# Benchmark Model



$$\kappa = 0$$

$$\Omega_r + \Omega_m + \Omega_\Lambda = 1$$

$$\Omega_r = \Omega_\gamma + \Omega_\nu$$

$$\Omega_m = \Omega_{\text{bary}} + \Omega_{\text{dm}}$$

$$\Omega_{r,0} = 0.31$$

$$\Omega_{\text{bary},0} = 0.048$$

$$\Omega_{\text{dm},0} = 0.262$$

$$\Omega_{r,0} = 9.0 \times 10^{-5}$$

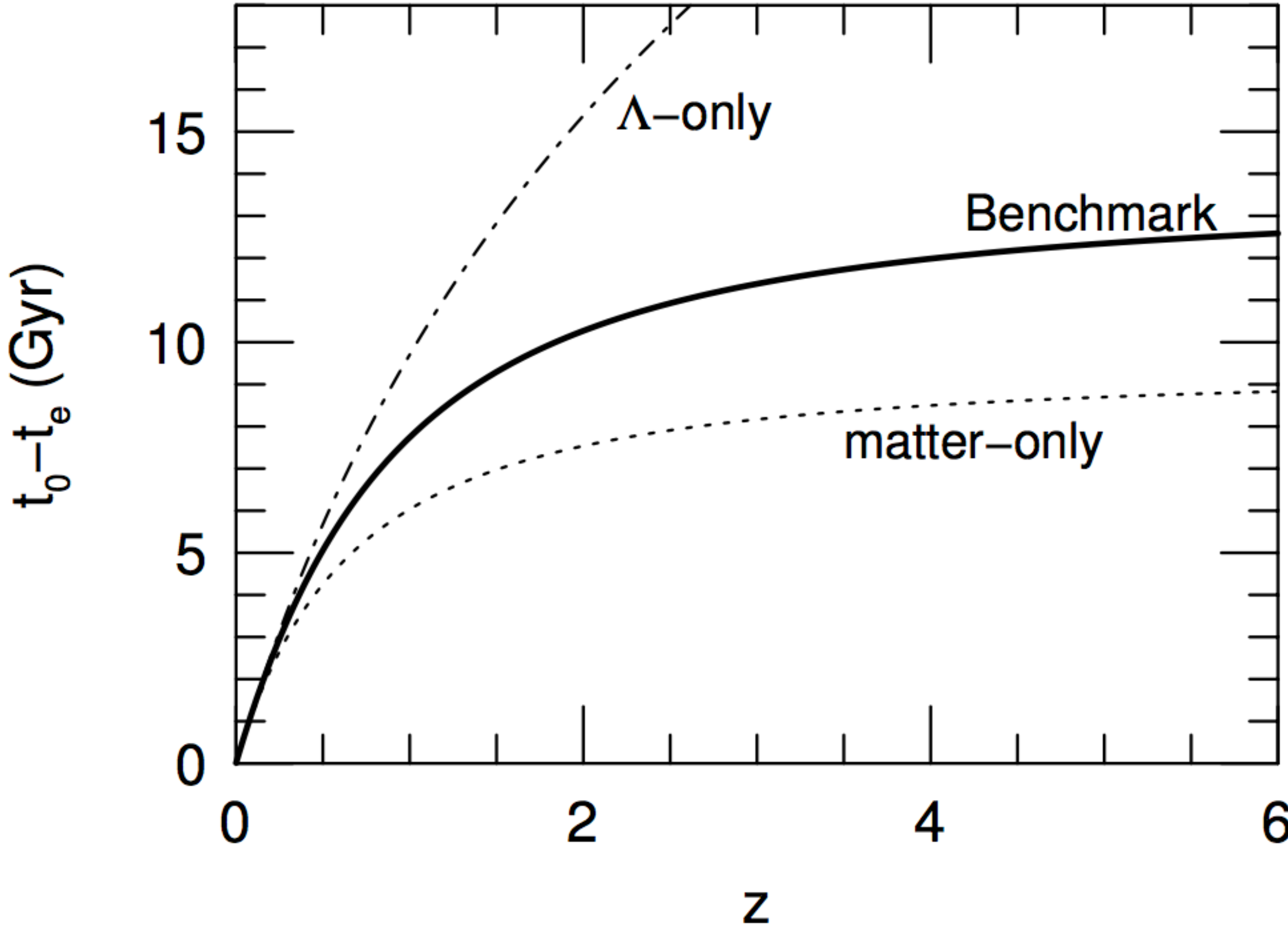
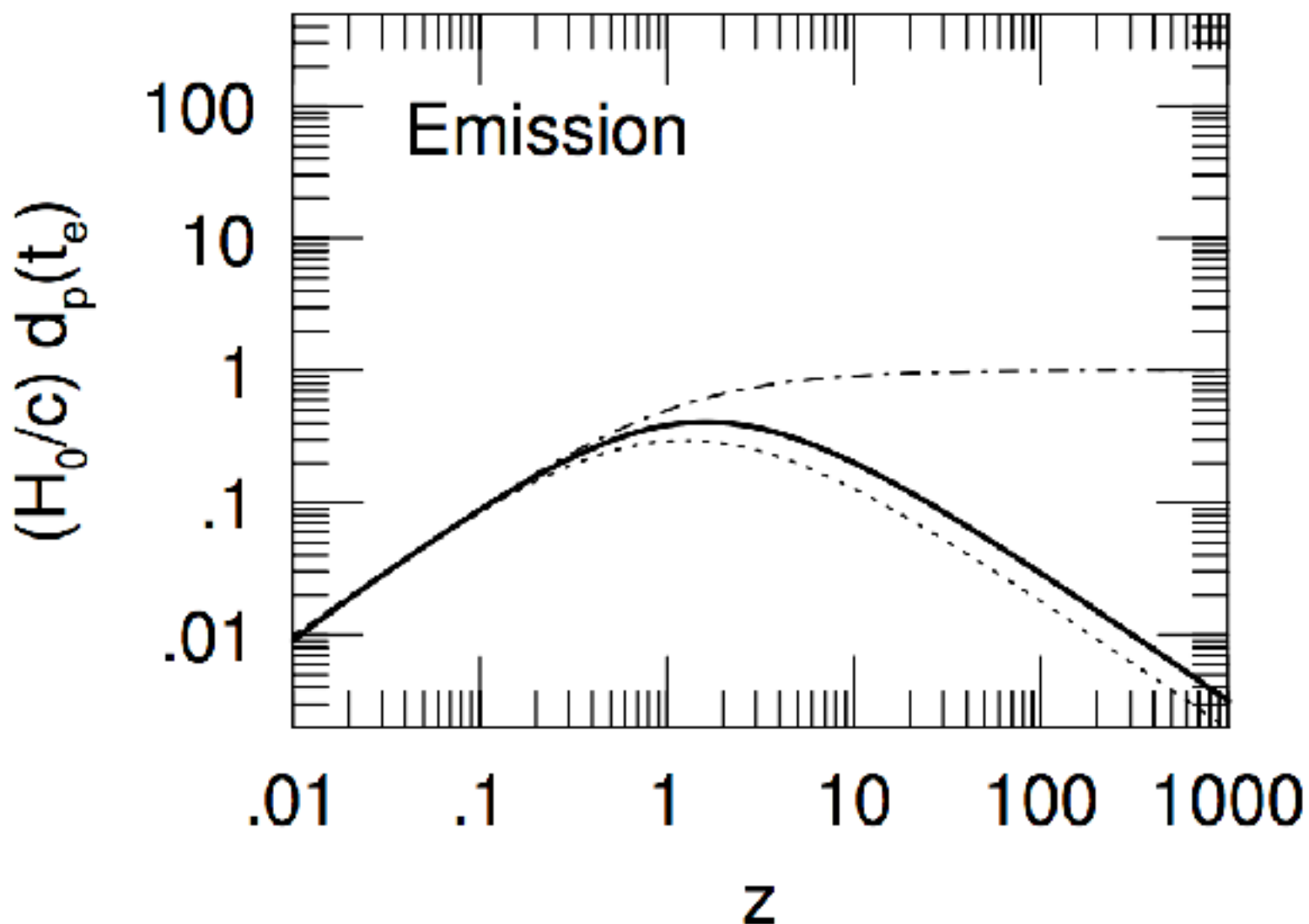
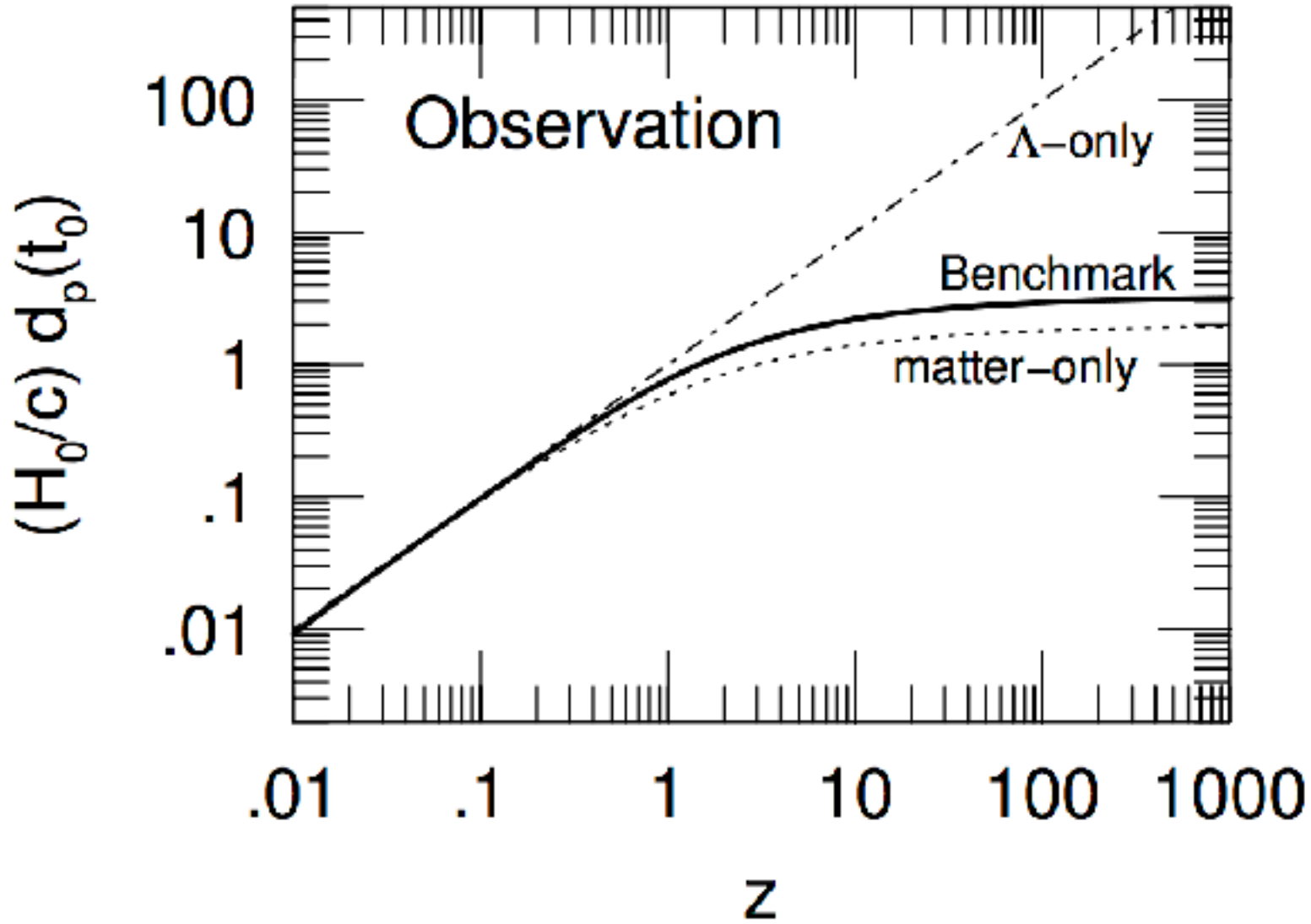
$$\Omega_{\gamma,0} = 5.35 \times 10^{-5}$$

$$\Omega_{\nu,0} = 3.65 \times 10^{-5}$$

$$\Omega_{\Lambda,0} \approx 0.69$$

Rad. – Matter :	$a_{\text{rm}} = 2.9 \times 10^{-4}$	$t_{\text{rm}} = 50 \text{ kyr}$
Matter – $\Lambda$ :	$a_{\text{m}\Lambda} = 0.77$	$t_{\text{m}\Lambda} = 10.2 \text{ Gyr}$
Now :	$a_0 = 1$	$t_0 = 13.7 \text{ Gyr}$

# Benchmark Model





**<https://tritonstation.wordpress.com/2019/01/28/a-personal-recollection-of-how-we-learned-to-stop-worrying-and-love-the-lambda/>**