

ASTR 5590 - Dispersion measure & FR measure

Recall the plasma freq. is the natural freq. of a plasma, where separated charge restores itself

$$\nu_p = \left(\frac{e^2 n_e}{4\pi^2 \epsilon_0 m_e} \right)^{1/2} \approx 9 \left(\frac{n_e}{1 \text{ cm}^{-3}} \right)^{1/2} \text{ kHz}$$

Radio freq. generally $> 10 \text{ MHz}$, so

$$\nu_p / \nu \ll 1$$

Group velocity of light thru the medium

$$\text{is } v_{gr} = c \left(1 - \left(\frac{\nu_p}{\nu} \right)^2 \right)^{1/2}$$

$$\approx c \left[1 - \frac{1}{2} \left(\frac{\nu_p}{\nu} \right)^2 \right]$$

If you have a pulse of emission, the pulse travels to us w/ a speed that depends on freq.

Arrival time:

$$t_a = \int_0^l \frac{dl}{v_{gr}} = \int_0^l \frac{dl}{c} \left[1 + \frac{1}{2} \left(\frac{\nu_p}{\nu} \right)^2 \right]$$

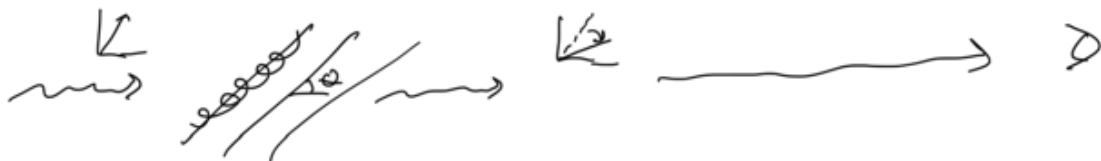
$$= \frac{l}{c} + \frac{e^2}{8\pi^2 \epsilon_0 m_e c} \frac{1}{\nu^2} \int_0^l n_e dl$$

$$\approx 4.2 \times 10^2 \left(\frac{\nu}{\text{Hz}} \right)^{-2} \left[\int \left(\frac{ne}{\text{m}^{-3}} \right) \frac{dl}{\text{parsec}} \right] S$$

Allas measurement of $\int n_e dl \rightarrow$ possible
 toward ~ 2000 Galactic pulsars
 If n_e assumed, can solve for distance

FR Measure

Because the plasma freq. + gyro freq.
 are both much less than radio
 observation freq., linear polarized
 emission gets rotated along the
 direction of the magnetic field



Decompose \vec{E} field vector in left/right
 handed ell. polarizations, which rotate
 in opposite senses thru the plasma

$$n^2 = 1 - \frac{(\nu_p/\nu)^2}{1 \pm (\nu_s/\nu) \cos \theta}$$

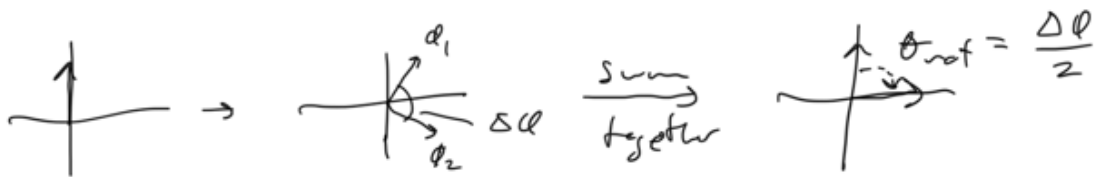
n : index of refraction

For $\nu_p/\nu \ll 1$ + $\nu_s/\nu \ll 1$, the diff.
 $\dots \nu_p^2 \nu_s \dots A$

in n is $\Delta n = \frac{1}{n^3} \cos \theta$

Phase diff. $\Delta \phi$ b/w waves is

$$\Delta \phi = \frac{2\pi n}{c} \Delta n dl$$



$$\Delta \theta = \frac{\pi n_p^2 n_s}{c n^2} \cos \theta dl$$

$$n_s = 2.8 \times 10^{10} \left(\frac{B_{||}}{T} \right) \text{Hz}, \quad n_p = 9 \left(\frac{n_e}{\text{m}^{-3}} \right)^{1/2} \text{Hz}$$

$$\theta = \frac{\pi}{c n^2} \int_0^L n_p^2 n_s \cos \theta dl$$

$$= 8.1 \times 10^3 \left(\frac{\Delta}{\text{m}} \right)^2 \int_0^L \left(\frac{n_e}{\text{m}^{-3}} \right) \left(\frac{B_{||}}{T} \right) \frac{dl}{\text{pc}}$$

★ Product of $n_e B_{||}$ along l.o.s.

Can get avg. B along l.o.s. if measure both rot. measure & disp. measure

$$\langle B_{||} \rangle \propto \frac{RM}{\text{Disp.} \cdot M} \propto \frac{\int n_e B_{||} dl}{\int n_e dl}$$

BUT, \int is weighted by the density along the l.o.s., so the $\uparrow n_e$ regions contribute more to $\langle B_{||} \rangle$

→ if w_e & B_{it} correlate, then
 $\langle B_{it} \rangle$ will be biased
& won't reflect true value of B