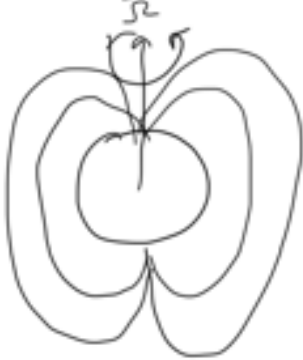


## ASTR 5590 - Pulsars III

### Pulsar Magnetosphere



- Complex problem to find a solution for
- take  $d=0$  ( $\Omega$  &  $B$  aligned)
  - $B$  field is frozen in

- expect an electric field to therefore be present: **Why?**

Get an induced electric field

from motion:  $\vec{E}_i = \vec{v} \times \vec{B}$

But, charges move to remove field, so

$$\vec{E} + (\vec{v} \times \vec{B}) = \vec{E} + [(\vec{\Omega} \times \vec{r}) \times \vec{B}] = 0$$

↑ charges moving to cancel out  $\vec{E}_i$

**\*** Recall, what's conductivity of plasma?  
 → high; for NS assume  $\infty$

For a dipole, this leads to an external electrostatic potential (quadrupolar)

$$\phi_E(r, \theta) = - \frac{B_0 \Omega R^5}{6 r^3} (3 \cos^2 \theta - 1)$$

(volts)

$$R = \frac{m_0 \mu}{\rho} \text{ (mks) or } 2\mu r \text{ . .}$$

$\sim 4\pi R^2 \dots \frac{1}{R^2} \text{ (cgs)}$   
 $\hookrightarrow$  polar field,  $\sim 10^{12} \text{ G}$

$$E = \nabla \phi, \text{ so } \nabla \phi \sim \frac{B_0 \Omega R^5}{r^4}$$

$$E(R) \sim B_0 \Omega R$$

Cheb pulsar surface E&M vs. grav.

$$\frac{e(\vec{v} \times \vec{B})}{6m/r^2} \sim \frac{e B_0 \Omega R}{6m/r^2} \approx 10^{12}$$

so charge motion determined only  
 by  $E \rightarrow e^-$  will be dragged  
 off the surface  $\rightarrow$  origin of PWN

HERE?

B/c quadrupolar, get charges  
 separated by a critical angle



$$3 \cos^2 \theta_c - 1 = 0$$

$$\theta_c = 55^\circ$$

$\rightarrow$  charge separates to  
 create  $E$  that cancels  
 out induced field  $\vec{E}_i$

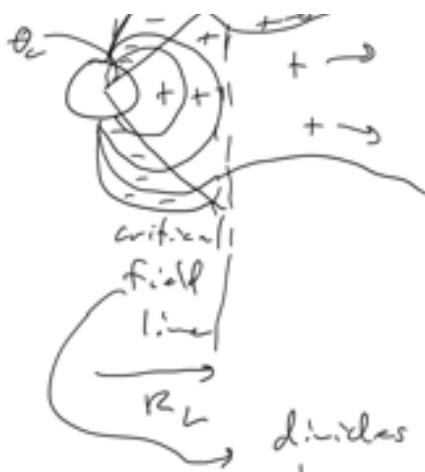
Remember the light cylinder, where  
 you get open field lines

B lines inside LC trap particles  
 - open field lines allow particles  
 to escape (causing  $E$  loss), but...

★ Fig. 13.23



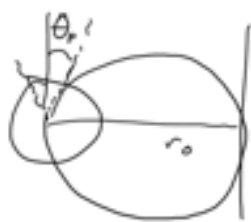
$$r = r_0 \sin^2 \theta$$



radius of line @ equator  
 $r_0 > R_L \rightarrow$  open field line

$$R_L = \frac{c}{\Omega}$$

divides whether + or - charge leaves pulsar



angle associated w/ last closed field line gives size of conical beam

$$R \approx R_L \sin^2 \theta_p$$

- assume small,  $\theta_p \approx \left(\frac{R}{R_L}\right)^{1/2} = \left(\frac{\Omega R}{c}\right)^{1/2}$

$$\theta_p = 1^\circ \left(\frac{R}{10 \text{ km}}\right)^{1/2} \left(\frac{P}{1 \text{ s}}\right)^{-1/2}$$

so for  $P \sim 0.1 \text{ s}$ ,  $\theta_p \sim 3^\circ$

Nicely explains short pulse width in assuming radio emission produced at surface

BUT, what is the radio emission??

- can't be incoherent given the brightness temperature of emission

(Rayleigh Jeans tail of BB)

$$T_{\text{RB}} = \frac{\lambda^2}{2k} \left(\frac{S_{\nu}}{\Omega}\right)$$

flux density  
solid angle of pulse

$\rightarrow > \dots 231 \dots$

$13 \sim 10^4 \text{ K} \rightarrow 100 \text{ eV}$  ---,  
particle E above esc. vel.

Need coherent emission, like a laser,  
which allows bunches of particles  
to emit simultaneously,  
e.g., some kind of "maser" process  
in the plasma within the magnetosphere

$T_B \sim 10^{10} \text{ K}$  in this case, comparable  
to radio transmitters on Earth!

@  $\uparrow E$ , like X-rays, the brightness  
temp. is lower,  $\sim 10^8 \text{ K}$  (10 MeV),  
so incoherent emission is allowed

To set  $e^-$  to v.  $\uparrow E$ , need to accel.  
them (need to get rid of + charges)  
 $\rightarrow$  inner/outer gaps potential sites

Fig. 13.15

$e^-$  more easily ripped off, radiate  
synchrotron until  $\vec{v} \parallel \vec{B}$   
- emits via curvature radiation  
(cyclotron, but due to curve of  $B$ )

- these photons interact w/ strong  $B$   
& produce  $e^-, e^+$ : pair cascade

$\gamma + B \rightarrow e^- e^+ \rightarrow \gamma\gamma \rightarrow e^- e^+ \rightarrow \dots$

$e^-$  &  $e^+$  are accel. in opposite directions

- So don't annihilate very fast  $\rightarrow$
- $\hookrightarrow$  cascades could produce coherent emission needed in radio called a "spark", creates the gap region above the pole
  - $\hookrightarrow$  sparks could cause fine structure <sup>seen</sup> in pulses or  $t \lesssim 100 \mu\text{s}$

Also could explain why pop. has

$$B \sim 10^{12} \text{G}, P \sim 10$$

$\hookrightarrow$  need this for pair cascades <sup>to</sup> occur

$\rightarrow$  predicts the death line:  
 $\sim \dots \sim |P|^{-2}$